## Conversion of Bound Muons: Lepton Flavour and Number Violation

## Tanja Geib

+ Alexander Merle: Phys. Rev. D93 (2016) $055039 \rightarrow$ technical details on $\mu^{-}-e^{-}$ + Stephen King, Alexander Merle, Jose Miguel No, Luca Panizzi: Phys. Rev. D93 (2016) $073007 \rightarrow$ complementarity of $\mu^{-}-e^{-}$with LHC
+ Alexander Merle, Kai Zuber: Phys. Lett. B764 (2017) $157 \rightarrow$ 'appetiser' $\mu^{-}-e^{+}$
+ Alexander Merle: arXiv:1612.00452 $\rightarrow$ technical details on $\mu^{-}-e^{+}$

Max Planck Institute for Physics

PSI Seminar, December 16, 2016

## Today's Agenda:

- What happens in a $\mu-e$ conversion?
- What are similarities and differences when considering $\mu^{-}-e^{-}$and $\mu^{-}-e^{+}$conversion?
- How to tackle $\mu^{-}-e^{+}$conversion (using the example of a realisation via doubly charged scalars)?
- Discovery potential for $\mu^{-}-e^{+}$conversion
- Open issues $\rightarrow$ where do we need to improve in order to get reliable predictions?
- Summary and Outlook


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## $\mu-e$ Conversion

What happens in a $\mu^{-}-e^{ \pm}$conversion ?? $\rightarrow$ experimentally a two-step process


First Step: $\mu^{-}$is captured in an 'outer'
atomic shell, and subsequently de-excites
to the 1 s ground state


Second Step: $\mu^{-}$is captured by the nucleus and reemits an $e^{ \pm}$
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## Energy Scales of the Process

- muon bound in 1s state with binding energy
$\epsilon_{B} \simeq \frac{m_{\mu}}{m_{e}} \cdot 13.6 \mathrm{eV} \cdot Z \ll m_{\mu} \xrightarrow{Z \leq 100}$ non-relativistic
 state
+ in good approximation: both nuclei at rest

$\Rightarrow e^{ \pm}$is relativistic particle under influence of Coulomb potential:
$E_{e} \simeq E_{\mu} \simeq m_{\mu}$ and $m_{e} \simeq 0$
- for 4-momentum transfer $q^{\prime}=p_{e}-p_{\mu}$



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## $\mu^{-}-e^{-}$vs $\mu^{-}-e^{+}$Conversion


from
TG, Merle, Zuber Phys.Lett. B764 (2017) 157

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\mu^{-}-\boldsymbol{e}^{-}
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further investigations needed:
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- needs to occur at two nucleons to achieve $\Delta Q=2 \rightarrow$ similar to $0 \nu \beta \beta$
- around $40 \%$ of the process total are g.s. $\rightarrow$ g.s.


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$\rightarrow$ derive a more involved spectrum for the positrons


## Improvements from Upcoming Experiments

Snapshot on current limits and sensitivities of upcoming experiments:
Future sensitivity for $\mu-e$ conversion

| $\mathrm{Pb}-208$ prese |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
| Au-197 |  |  |  |  |
| Ti-48 |  |  |  |  |
| $\mathrm{Si}-28$ |  |  |  |  |
| Al-27 twe experimental figures |  |  |  |  |
| $10^{-19}$ | $10^{-17}$ | $10^{-15}$ | $10^{-13}$ | $10^{-11}$ |
|  | $\mathrm{BR}\left(\mu^{-}-e^{-}\right.$conversion) |  |  |  |

past: SINDRUM II for ${ }^{48} \mathrm{Ti}$ (1993), ${ }^{208} \mathrm{~Pb}$ (1995), ${ }^{197} \mathrm{Au}$ (2006)
future: DeeMee for ${ }^{28} \mathrm{Si}$, COMET and Mu2e (taking data $\sim 2018$ ) for ${ }^{27} \mathrm{Al}$, PRISM/PRIME for ${ }^{48} \mathrm{Ti}$
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$\rightarrow$ sensitivities on both processes will increase by several orders of magnitude in the foreseeable future
$\rightarrow$ target both processes with the same experimental setup
$\Rightarrow$ it's time to investigate these bound muon conversions to describe them within a general framework independent of the respective particle physics realisation

# How to tackle $\boldsymbol{\mu}^{-} \boldsymbol{-} \boldsymbol{e}^{-}$conversion (using the example of a realisation via doubly charged scalars)? 

## Effective theory of a doubly charged scalar singlet

 based on King, Merle, Panizzi JHEP 1411 (2014) 124Minimal extension of SM:

- only one extra particle: $S^{++}$
$\rightarrow$ lightest of possible new particles (UV completion e.g. Cocktail model)
$\rightarrow$ reduction of input parameters
- tree-level coupling to SM (to charged right-handed leptons) $\rightarrow$ LNV and LFV!
- effective Dim-7 operator (necessary to generate neutrino mass)
$\mathcal{L}=\mathcal{L}_{\mathrm{SM}}-V(H, S)$
$+\left(D_{\mu} S\right)^{\dagger}\left(D^{\mu} S\right)$

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\begin{equation*}
+\left(D_{\mu} S\right)^{\dagger}\left(D^{\mu} S\right)+f_{a b} \overline{\left(\ell_{R a}\right)^{c}} \ell_{R b} S^{++}+\text {h.c. } \tag{2}
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$\mu^{-}-e^{-}$Conversion: Universally Valid for Models Involving Doubly Charged Singlet Scalars based on TG, Merle Phys.Rev. D93 (2016) 055039
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relevant diagrams

## Different Contributions to $\mu^{-}-e^{-}$Conversion

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\sim \mathcal{O}(\overbrace{}^{-15} \mathrm{~m})
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- estimate nuclear radius: $R=\overbrace{r_{0}} A^{1 / 3} \sim \mathcal{O}\left(10^{-15} \mathrm{~m}\right)$
- reduced Bohr radius: $\underset{\mathcal{O}\left(10^{-10} \mathrm{~m}\right)}{a_{0}} \frac{m_{e}}{m_{\mu}} \sim \mathcal{O}\left(10^{-13} \mathrm{~m}\right)$


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$\Rightarrow$ contributions need to be treated qualitatively differently!!


## Photonic Contribution


$\rightarrow$ wave functions for $\mu^{-}$and $e^{-}$obtained by solving modified Dirac equation ( + Coulomb potential)
$\rightarrow$ Most general (Lorentz-) invariant expression for $\mid$

with $q^{\prime}=p_{e}-p_{\mu}$.
In non-relativistic limit:
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\mathcal{M} \propto \int \mathrm{d}^{3} r \overline{\psi_{j l m}^{e}}\left(p_{e}, r\right) \Gamma^{\nu} \psi_{j_{\mu} l_{\mu} m_{\mu}}^{\mu}\left(p_{\mu}, r\right) \underbrace{\langle N| \bar{q} \gamma_{\nu} q|N\rangle}_{Z e \rho^{(P)}(r) \delta_{\nu 0}}
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Write branching ratio as product of nuclear and particle physics parts

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\operatorname{BR}\left(\mu^{-} N \rightarrow e^{-} N\right)=\frac{8 \alpha^{5} m_{\mu} Z_{\text {eff }}^{4} Z F_{p}^{2}}{\Gamma_{\text {capt }}} \equiv^{2}
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see Kuno, Okada
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$\rightarrow$ factorisation works perfectly for photonic contributions
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## Photonic Contribution: Results

In good approximation (up to a few per cent), we use
$F_{1}\left(q^{\prime 2}\right)=G_{1}\left(q^{\prime 2}\right)=-f_{e a}^{*} f_{a \mu}\left[\frac{2 m_{a}^{2}+m_{\mu}^{2} \log \left(\frac{m_{a}}{M_{S}}\right)}{12 \pi^{2} M_{S}^{2}}+\frac{\sqrt{m_{\mu}^{2}+4 m_{2}^{2}}\left(m_{\mu}^{2}-2 m_{a}^{2}\right)}{12 \pi^{2} m_{\mu} M_{S}^{2}} \operatorname{Arctanh}\left(\frac{m_{\mu}}{\sqrt{m_{\mu}^{2}+4 m_{a}^{2}}}\right)\right]$
$F_{2}\left(q^{\prime 2}\right)=-G_{2}\left(q^{\prime 2}\right)=f_{e a}^{*} f_{a \mu} \frac{m_{\mu}^{2}}{24 \pi^{2} M_{S}^{2}}$
with $q^{\prime 2}=-m_{\mu}^{2}$ for the particle physics factor:

$$
\begin{aligned}
& \left.\equiv_{\text {photonic }}^{2}=\frac{1}{288 \pi^{4} m_{\mu}^{2} M_{S}^{4}} \right\rvert\, \sum_{a=e, \mu, \tau} f_{e a}^{*} f_{a \mu}\left(4 m_{a}^{2} m_{\mu}-m_{\mu}^{3}+2\left(-2 m_{a}^{2}+m_{\mu}^{2}\right) \sqrt{4 m_{a}^{2}+m_{\mu}^{2}}\right. \\
&\text { Arctanh } \left.\left[\frac{m_{\mu}}{\sqrt{4 m_{a}^{2}+m_{\mu}^{2}}}\right]+m_{\mu}^{3} \ln \left[\frac{m_{a}^{2}}{M_{S}^{2}}\right]\right)\left.\right|^{2}
\end{aligned}
$$

## $\rightarrow$ while $F_{2}$ is independent of $m_{a},\left|F_{1}\right|$ decreases with increasing $m_{a}$

 $\rightarrow$ hierarchy: $\left|F_{2}\right|<\left|F_{1}\right|$ but for $M_{S} \sim 10 \mathrm{GeV}$ of order $10 \%$ $\rightarrow$ compare to $\mu \rightarrow$ e $: F_{1}\left(q^{\prime 2}=0\right)=G_{1}\left(q^{\prime 2}=0\right)=0$ and
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& \text { Arctanh } \left.\left[\frac{m_{\mu}}{\sqrt{4 m_{a}^{2}+m_{\mu}^{2}}}\right]+m_{\mu}^{3} \ln \left[\frac{m_{a}^{2}}{M_{S}^{2}}\right]\right)\left.\right|^{2}
\end{aligned}
$$

$\rightarrow$ while $F_{2}$ is independent of $m_{a},\left|F_{1}\right|$ decreases with increasing $m_{a}$
$\rightarrow$ hierarchy: $\left|F_{2}\right|<\left|F_{1}\right|$ but for $M_{S} \sim 10 \mathrm{GeV}$ of order $10 \%$

## Photonic Contribution: Results

In good approximation (up to a few per cent), we use
$F_{1}\left(q^{\prime 2}\right)=G_{1}\left(q^{\prime 2}\right)=-f_{e a}^{*} f_{a \mu}\left[\frac{2 m_{a}^{2}+m_{\mu}^{2} \log \left(\frac{m_{a}}{M_{S}}\right)}{12 \pi^{2} M_{S}^{2}}+\frac{\sqrt{m_{\mu}^{2}+4 m_{2}^{2}}\left(m_{\mu}^{2}-2 m_{a}^{2}\right)}{12 \pi^{2} m_{\mu} M_{S}^{2}} \operatorname{Arctanh}\left(\frac{m_{\mu}}{\sqrt{m_{\mu}^{2}+4 m_{a}^{2}}}\right)\right]$
$F_{2}\left(q^{\prime 2}\right)=-G_{2}\left(q^{\prime 2}\right)=f_{e a}^{*} f_{a \mu} \frac{m_{\mu}^{2}}{24 \pi^{2} M_{S}^{2}}$
with $q^{\prime 2}=-m_{\mu}^{2}$ for the particle physics factor:

$$
\begin{aligned}
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$\rightarrow$ hierarchy: $\left|F_{2}\right|<\left|F_{1}\right|$ but for $M_{S} \sim 10 \mathrm{GeV}$ of order $10 \%$
$\rightarrow$ compare to $\mu \rightarrow e \gamma: F_{1}\left(q^{\prime 2}=0\right)=G_{1}\left(q^{\prime 2}=0\right)=0$ and
$F_{2}\left(q^{\prime 2}=0\right)=-G_{2}\left(q^{\prime 2}=0\right)=F_{2}\left(q^{\prime 2}=-m_{\mu}^{2}\right) \Rightarrow \mu^{-}-e^{-}$conversion enhanced by $F_{1}$ contribution

## Non-Photonic Contribution

Short-range $\leftrightarrow$ takes place inside the nucleus: EFT treatment $\Rightarrow$ Integrating out the Z-boson:

$\rightarrow$ four-point vertices
$\rightarrow$ consider operators up to dimension six
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\mathcal{L}_{\text {short-range }}=-\frac{G_{F}}{\sqrt{2}} \underbrace{\frac{2\left(1+k_{q} \sin ^{2} \theta_{W}\right) \cos \theta_{W}}{g} A_{R}\left(q^{\prime 2}\right)}_{\operatorname{g}_{R V(q)}} \overline{e_{R}} \gamma_{\nu} \mu_{R} \bar{q} \gamma^{\nu} q
$$

## Non-Photonic Contribution

We can write the branching ratio as

$$
\operatorname{BR}\left(\mu^{-} N \rightarrow e^{-} N\right)=\frac{8 \alpha^{5} m_{\mu} Z_{\text {eff }}^{4} Z F_{p}^{2}}{\Gamma_{\text {capt }}} \Xi_{\text {non-photonic }}^{2}\left(Z, N, A_{R}\left(q^{\prime 2}\right)\right)
$$

$\rightarrow$ no perfect factorisation anymore: 三 modified to be function of nuclear characteristics
$\rightarrow$ instead of lines we do have bands with finite widths for $\overline{\text { E }}$
$\Rightarrow$ determine
from amputated diagrams with off-shell
Z-Boson
Combining photonic and non-photonic contributions: $\overline{\bar{p}}_{\text {particle }} \rightarrow \overline{\bar{c}}_{\text {combined }}(Z, N)=\bar{\Xi}_{\text {photonic }}+\bar{E}_{\text {non-photonic }}(Z, N)$

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$\rightarrow$ dependence on nuclear characteristics

## Combining the Contributions: Results

 see TG, Merle Phys.Rev. D93 (2016) 055039

Benchmark Points:
$f_{a b}$ such that LFV/LNV bounds fulfilled + suitable neutrino mass matrix reproduced

- 'red': $f_{e e} \simeq 0$ and $f_{e \tau} \simeq 0$
- 'purple': $f_{e e} \simeq 0$ and $f_{e \mu} \simeq \frac{f_{\mu \tau}^{*}}{f_{\mu \mu}^{*}} f_{e \tau}$
- 'blue': $f_{e \mu} \simeq \frac{f_{\mu \tau}^{*}}{f_{\mu \mu}^{*}} f_{e \tau}$
choose representative 'average' set for each scenario to display $M_{S}$ dependence


## Combining the Contributions: Results

see TG, Merle Phys.Rev. D93 (2016) 055039

$\rightarrow$ widths of the bands so small that appear as lines
$\rightarrow$ non-photonic (DASHED) contributions negligibly small
$\rightarrow$ approximate process by its purely photonic (SOLID) contribution
$\rightarrow$ factorisation: dependence on isotope only in width of limit

## Results: Photonic Contribution vs $\mu \rightarrow e \gamma$

see TG, Merle Phys.Rev. D93 (2016) 055039 and King, Merle, Panizzi JHEP 1411 (2014) 124

strongest bound for red, weakest for blue points

## Results: Photonic Contribution vs $\mu \rightarrow e \gamma$

see TG, Merle Phys.Rev. D93 (2016) 055039 and King, Merle, Panizzi JHEP 1411 (2014) 124
For $\mu^{+} \rightarrow \boldsymbol{e}^{+} \gamma$ : strongest bound for red, weakest for blue points

$$
\mathcal{A} \propto\left|f_{e e} f_{e \mu}^{*}+f_{e \mu} f_{\mu \mu}^{*}+f_{e \tau} f_{\tau \mu}^{*}\right| \cdot C
$$

$\rightarrow$ some amount of cancellation
For $\mu^{-}-e^{-}$conversion
$\rightarrow$ flavour-dependent coefficients:

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!! other way around !!
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$$

$\rightarrow$ flavour-dependent coefficients:
prevent similar cancellations $\rightarrow$ shape of amplitude leads to drastical change (not mainly off-shell contributions)

## Results: Complementarity

see TG, King, Merle, No, Panizzi Phys.Rev. D93 (2016) 073007




From 'average scenarios' (depicted by lines), we can estimate the lower limits on $\mathbf{M}_{S}$ resulting from $\mu$-e conversion:

COMET I (Al-27) [GeV]

| blue curve | $M_{S}>131.9-447.1$ | $M_{S}>1031.5-13271.3$ | $M_{S}>1954.1$ |
| :---: | :---: | :---: | :---: | :---: |
| purple curve | $M_{S}>42.5-152.3$ | $M_{S}>360.7-4885.2$ | $M_{S}>694.5$ |
| red curve | $M_{S}>33.9-118.1$ | $M_{S}>276.3-3656.1$ | $M_{S}>528.0$ |

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# How to tackle $\boldsymbol{\mu}^{-}-\boldsymbol{e}^{+}$conversion (using the example of a realisation via doubly charged scalars)? 

## Motivation

In the following, we perform the computation for the decay rate for one particular short-range operator $\epsilon_{3}^{L L z}$. But why?!
$\rightarrow$ has the nuclear matrix elements (for ${ }^{48} \mathrm{Ti}$ ) that we use: $\epsilon_{3}^{L L z}$
$\rightarrow$ explicit computation focussing on the nuclear physics
$\Rightarrow$ includes the formalism that we want make accessible to the
particle physics community

- many aspects do not change if another operator was realised
guideline how to use existing results and establish a general formalism
to replicate such a computation for different scenarios


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- There are a few earlier references available focussing on $\mu^{-}-e^{+}$ conversion from Majorana neutrinos but no uniform formalism is used:
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## $\mu^{-}-e^{+}$Conversion from doubly charged scalars

- formalism to describe $\mu^{-}-e^{+}$conversions within general framework
- use EFT to neatly separate the nuclear physics from the respective particle physics realisation of the conversion $\rightarrow$ factorisation


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$\rightarrow$ map the model onto short-range operators


## General Formalism for $\mu^{-}-e^{+}$Conversion from

 Short-Range Operators based on Päs et al. Phys.Lett. B498 (2001) 35, and TG, Merle, Zuber Phys.Lett. B764 (2017) 157Employ EFT formalism to generally describe $\mu^{-}-e^{+}$conversion $\Rightarrow \operatorname{dim}-9$ short-range operators:

$$
\begin{aligned}
\mathcal{L}_{\text {short-range }}^{\mu e} & =\frac{G_{F}^{2}}{2 m_{\rho}} \sum_{x, y, z=L, R}\left[\epsilon_{1}^{x y z} J_{x} J_{y} j_{z}+\epsilon_{2}^{x y z} J_{x}^{\nu \rho} J_{y, \nu \rho} j_{z}+\epsilon_{3}^{x y z} J_{x}^{\nu} J_{y, \nu} j_{z}+\epsilon_{4}^{x y z} J_{x}^{\nu} J_{y, \nu \rho} j_{z}^{\rho}\right. \\
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\end{aligned}
$$

using the hadronic currents:

$$
J_{R, L}=\bar{d}\left(1 \pm \gamma_{5}\right) u, \quad J_{R, L}^{\nu}=\bar{d} \gamma^{\nu}\left(1 \pm \gamma_{5}\right) u, \quad J_{R, L}^{\nu \rho}=\bar{d} \sigma^{\nu \rho}\left(1 \pm \gamma_{5}\right) u,
$$

and the leptonic currents:

$$
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$$

$\Rightarrow$ derive the decay rate using the example of doubly charged scalars

## Deriving the Decay Rate for $\epsilon_{3}$ based on TG, Merle arXiv:1612.00452

 Start with the amplitude obtained from EFT diagram
which is

$$
\begin{aligned}
\left\langle N^{\prime}, f\right| S_{\mathrm{eff}}^{(1)}|N, i\rangle & =-i\left\langle N^{\prime}, f\right| \int \mathrm{d}^{4} x \widehat{T}\left\{\mathcal{L}_{\mathrm{eff}}(x)\right\}|N, i\rangle \\
& =-i \frac{G_{F}^{2}}{2 m_{p}} \epsilon_{3}^{L L R} \int \mathrm{~d}^{4} x\left\langle N^{\prime}, f\right| \widehat{T}\left\{J_{L, \nu}(x) J_{L}^{\nu}(x) j_{R}(x)\right\}|N, i\rangle
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Structure can be split into hadronic and leptonic parts:

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Leptonic part:

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- positron propagates freely under the influence of the nucleus' Coulomb potential
$\Rightarrow$ need to modify the free spinors $u$ and $v$ respectively
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\langle f| j_{R}(x)|i\rangle=2 \mathrm{e}^{i k_{e} \cdot x} \mathrm{e}^{-i E_{\mu} \cdot x^{0}} \sqrt{F\left(Z-2, E_{e}\right)} \phi_{\mu}(\vec{x}) \overline{v_{e}}\left(k_{e}\right) \mathrm{P}_{\mathrm{R}} u_{\mu}\left(k_{\mu}\right)
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## Deriving the Decay Rate for $\epsilon_{3}$ based on TG, Merle arXiv:1612.00452

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- hadronic currents can be approximated by their non-relativistic versions $J_{\nu}(\vec{x})$
- need to account for quarks' distribution within the nucleus $\rightarrow$ dipole parametrisation factor $\tilde{F}\left(\vec{k}^{2}, \wedge_{i}\right)$
- two nucleon interactions $\rightarrow$ take place with finite distance $\rightarrow$ introduce second location $\tilde{x}$ over which we also "sum" $\int \mathrm{d}^{3} \tilde{x}$
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- hadronic currents can be approximated by their non-relativistic versions $J_{\nu}(\vec{x})$

- need to account for quarks' distribution within the nucleus $\rightarrow$ dipole parametrisation factor $\tilde{F}\left(\vec{k}^{2}, \Lambda_{i}\right)$
- two nucleon interactions $\rightarrow$ take place with finite distance $\rightarrow$ introduce second location $\tilde{x}$ over which we also "sum" $\int \mathrm{d}^{3} \tilde{x}$
$\Rightarrow$ need to modify hadronic currents $J_{\nu}$ respectively

$$
\left\langle N^{\prime}\right| \widehat{T}\left\{J_{L, \nu}(x) J_{L}^{\nu}(x)\right\}|N\rangle \rightarrow \int \mathrm{d}^{3} \tilde{x} \int \frac{\mathrm{~d}^{3} k}{(2 \pi)^{3}}\left\langle N^{\prime}\right| \mathrm{e}^{i \vec{k} \cdot(\vec{x}-\vec{x}) \tilde{F}^{2}\left(\vec{k}^{2}, \Lambda_{i}\right) J_{L, \nu}(\overrightarrow{\tilde{x}}) J_{L}^{\nu}(\vec{x})|N\rangle . N . . . \mid}
$$

## Deriving the Decay Rate for $\epsilon_{3}$ based on TG, Merle arXiv:1612.00452

## Next:

- perform $x^{0}$ integration
$\rightarrow$ conservation of external energies $2 \pi \delta\left(E_{i}+E_{\mu}-E_{f}-E_{e}\right)$
- write non-relativistic currents in term of effective transition operators:
 with nuclear isospin raising operator $\tau_{m}^{-}$and the dominant spin structures given by the Fermi onerator and the Gamoun-Teller onerator

physics model:

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with nuclear isospin raising operator $\tau_{m}^{-}$and the dominant spin structures given by the Fermi operator and the Gamow-Teller operator
$\Rightarrow$ allows for factorisation of nuclear physics from respective particle physics model:

$$
\mathcal{M}=\frac{G_{F}^{2} \epsilon_{3}^{L R} g_{A}^{2} m_{e}}{2 R} \sqrt{F\left(Z-2, E_{e}\right)} \delta\left(E_{f}-E_{i}+E_{e}-E_{\mu}\right) \overline{v_{e}}\left(k_{e}\right) \mathrm{P}_{\mathrm{R}} u_{\mu}\left(k_{\mu}\right) \mathcal{M}^{\left(\mu^{-}, e^{+}\right) \phi}
$$

with $\mathcal{M}^{\left(\mu^{-}, e^{+}\right) \phi}$ being the nuclear matrix element.

## Deriving the Decay Rate for $\epsilon_{3}$ based on TG, Merle arXiv:1612.00452

From amplitude to decay rate using Fermi's Golden rule:

$$
\Gamma=2 \pi \frac{1 / T}{(2 \pi)^{3}} \int \mathrm{~d}^{3} k_{e}|\mathcal{M}|^{2}
$$

So, we need to

- spin sum/average $\rightarrow 1 / 4$
- rewrite nuclear matrix element using that the muon wave function varies only slowly within nucleus:
- square delta-function: " $\delta\left(E_{f}-E_{i}+E_{e}-E_{\mu}\right)^{2 "}=\frac{T}{2 \pi} \delta\left(E_{f}-E_{i}+E_{e}-E_{\mu}\right)$

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and obtain the



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$$
\Gamma=\frac{g_{A}^{4} G_{F}^{4} m_{e}^{2} m_{\mu}^{2}\left|\epsilon_{3}^{L L R}\right|^{2}}{32 \pi^{2} R^{2}}\left|F\left(Z-2, E_{e}\right)\right|\left\langle\phi_{\mu}\right\rangle^{2}\left|\mathcal{M}^{\left(\mu^{-}, e^{+}\right)}\right|^{2}
$$

$\rightarrow$ can be generalised to $\epsilon_{3}^{x y z}$ for $x=y$
$\rightarrow$ for $x \neq y$ there is a relative sign switched in the nuclear matrix element

## Further Realisations of $\epsilon_{3}$

## Cheng-Geng-Ng model Cheng, Geng, Ng Phys.Rev. D75 (2007) 053004



Heavy Majorana neutrinos
Domin, Kovalenko, Faessler,
Simkovic Phys.Rev. C70 (2004) 065501

Left-Right symmetric

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EFT with doubly charged scalar King, Merle, Panizzi
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Left-Right symmetric models Pritimita, Dash, Patra JHEP 1610 (2016) 147

## Reach of Future Experiments for $\epsilon_{3}$

based on TG, Merle, Zuber Phys.Lett. B764 (2017) 157


- obvious: limits on $0 \nu \beta \beta$ are superior to those of $\mu^{-}-e^{+}$ conversion by orders of magnitude
- but also apparent: there are models where LNV is much more prominent in e $\mu$ instead of ee sector
- there are much more settings/operators which are likely to sit within reach for the next generation of experiments


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## Reach of Future Experiments for $\epsilon_{3}$

 based on TG, Merle, Zuber Phys.Lett. B764 (2017) 157LNV discovery potential


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- but also apparent: there are models where LNV is much more prominent in $e \mu$ instead of ee sector
- there are much more settings/operators which are likely to sit within reach for the next generation of experiments
$\Rightarrow$ valuable new information from $\mu^{-}-e^{+}$conversion experiments


## Open Issues of $\mu^{-}-e^{+}$based on TG, Merle, Zuber Phys.Lett. B764 (2017) 157

However: several key pieces of information are missing!! We are in dire need of improvements from different areas:

- Experiment: more detailed sensitivity studies for $\mu^{-}-e^{+}$conversion
- Nuclear Matrix Elements:
- detailed study on percentage of process that is "coherent'
- hardly any nuclear matrix elements (NMEs) are available $\rightarrow$ need for NMEs for further element, e. g. ${ }^{27} \mathrm{Al}$, and for other operators like $\epsilon_{1,2}$
$\Rightarrow$ there are promising models but we cannot judge them properly
- Particle Physics: for many models there are no (detailed) studies on LNV in the $e \mu$ sector and no information on which effective operators are realised


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- Particle Physics: for many models there are no (detailed) studies on LNV in the e $\mu$ sector and no information on which effective operators are realised
$\Rightarrow$ Only if all three communities pull together, advances will be achieved!!


## Summary and Outlook

- orders of magnitude improvement of sensitivities in near-future experiments

```
- }\mp@subsup{\mu}{}{-}-\mp@subsup{e}{}{-}\mathrm{ conversion:
```

- FIRST work that treats $\mu^{-}-e^{-}$conversion in such detail,
 $\rightarrow$ analytic expression for $\overline{\text { }}_{\text {particle }}$
- complementarity: rich phenomenology of loop models $\rightarrow$ high- and low-energy processes $\rightarrow \mu^{-}-e^{-}$conversion important part of study
- $\mu^{-}-e^{+}$conversion:
- complete computation of the rate for the lepton flavour and number violating conversion process, mediated by the effective operator $\epsilon_{3}$
- pointed out open issues and further models/operators
- LNV possibly more prominent in $e \mu$ sector $\rightarrow$ experiments could make a countable physics impact
- open issues need to be addressed in order to proceed
- COMET: expecting to take first data in 2018


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# Thank you for your attention!! 

## Any questions?

## Backup Slides

## Generating the Neutrino Mass

The mass is generated at two-loop level via the diagram

which leads to the neutrino mass

$$
\mathcal{M}_{\nu, a b}^{2 \text {-loop }}=\frac{2 \xi m_{a} m_{b} M_{S}^{2} g_{a b}\left(1+\delta_{a b}\right)}{\Lambda^{3}} \mathcal{I}\left[M_{W}, M_{S}, \mu\right]
$$

$\longrightarrow$ Majorana mass term
$\longrightarrow$ further LNV processes

## Testing the Model

Selection of interesting processes: low energy physics

- neutrinoless double beta decay:

$$
\frac{\xi f_{e e}}{M_{S}^{2} \Lambda^{3}}<\frac{4.0 \cdot 10^{-3}}{T e V^{5}}
$$



- $\mu^{-} \rightarrow e^{-} \gamma$ :

$$
\left|f_{e e}^{*} f_{e \mu}+f_{e \mu}^{*} f_{\mu \mu}+f_{e \tau}^{*} f_{\mu \tau}\right|<3.2 \cdot 10^{-4} M_{S}^{2}[\mathrm{TeV}]
$$

## Testing the Model

## benchmark points:

$f_{a b}$ such that bounds fulfilled + suitable light neutrino mass matrix reproduced

- 'red': $f_{e e} \simeq 0$ and $f_{e \tau} \simeq 0$
- 'purple': $f_{e e} \simeq 0$ and $f_{e \mu} \simeq \frac{f_{\mu \tau}^{*}}{f_{\mu \mu}^{*}} f_{e \tau}$
- 'blue': $f_{e \mu} \simeq \frac{f_{\mu \tau}^{*}}{f_{\mu \mu}^{*}} f_{e \tau}$
complementary check with high energy experiments:
compute cross sections for e.g.



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complementary check with high energy experiments: compute cross sections for e.g.
- $S^{ \pm \pm} \rightarrow W^{ \pm \pm}$
- $S^{ \pm \pm} \rightarrow l_{a}^{ \pm \pm} l_{b}^{ \pm \pm}$
$\rightarrow$ some of the benchmark points already excluded by LHC data (7 TeV run)


## Photonic Contribution: Cross Check via UV Divergences

 In form of $i \mathcal{M}=e f_{e a}^{*} f_{a \mu} A_{\nu}\left(q^{\prime}\right) \bar{u}_{e}\left(p_{e}\right) \mathcal{I}^{\nu} u_{\mu}\left(p_{\mu}\right)$ :

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$$
-4 Q_{S} \int \frac{\mathrm{~d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} k\left(2 p_{\mu}-2 k+q^{\prime}\right)^{\nu}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k+q^{\prime}\right)^{2}-M_{S}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]} \stackrel{\text { div }}{\longrightarrow} \frac{2 i}{(4 \pi)^{2} \varepsilon} Q_{S} P_{L} \gamma^{\nu}
$$



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$$
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$$



$$
\begin{gathered}
-4 Q_{S} \int \frac{\mathrm{~d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} k\left(2 p_{\mu}-2 k+q^{\prime}\right)^{\nu}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k+q^{\prime}\right)^{2}-M_{S}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]} \stackrel{\text { div }}{\longrightarrow} \frac{2 i}{(4 \pi)^{2} \varepsilon} Q_{S} P_{L} \gamma^{\nu} \\
-4 Q_{I+} \int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L}\left(k+q^{\prime}+m_{a}\right) \gamma^{\nu}\left(k+m_{a}\right) P_{R}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]\left[\left(k+q^{\prime}\right)^{2}-m_{a}^{2}\right]} \stackrel{\text { div }}{\longrightarrow} \frac{-i}{(4 \pi)^{2} \varepsilon} Q_{I+} P_{L} \gamma^{\rho} \gamma^{\nu} \gamma_{\rho} P_{R}
\end{gathered}
$$

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& -4 Q_{I+} \int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L}\left(k+\phi^{\prime}+m_{a}\right) \gamma^{\nu}\left(k+m_{a}\right) P_{R}}{\left.\left[k^{2}-m_{a}^{2} I\left(\rho_{\mu}-k\right)^{2}-M_{S}^{2}\right]\left(k+q^{\prime}\right)^{2}-m_{a}^{2}\right]} \xrightarrow{\text { div }} \frac{-i}{(4 \pi)^{2} \varepsilon} Q_{l+} P_{L} \gamma^{\rho} \gamma^{\nu} \gamma_{\rho} P_{R}
\end{aligned}
$$



$$
-4 Q_{\mu}-\int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} k\left(\phi_{e}+m_{\mu}\right) \gamma^{\nu}}{\left[p_{e}^{2}-m_{\mu}^{2}\right]\left[\left(p_{e}-k\right)^{2}-M_{S}^{2}\left[k^{2}-m_{a}^{2}\right]\right.} \stackrel{\text { div }}{\longrightarrow} \frac{2 i}{(4 \pi)^{2} \varepsilon} \frac{Q_{\mu}-}{m_{\mu}^{2}} P_{L} \phi_{e}\left(\phi_{e}+m_{\mu}\right) \gamma^{\nu}
$$



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$$
\begin{aligned}
& -4 Q_{S} \int \frac{\mathrm{~d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} k\left(2 p_{\mu}-2 k+q^{\prime}\right)^{\nu}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k+q^{\prime}\right)^{2}-M_{S}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]} \stackrel{\text { div }}{\longrightarrow} \frac{2 i}{(4 \pi)^{2} \varepsilon} Q_{S} P_{L} \gamma^{\nu} \\
& -4 Q_{I}+\int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L}\left(k+q^{\prime}+m_{a}\right) \gamma^{\nu}\left(k+m_{a}\right) P_{R}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]\left[\left(k+q^{\prime}\right)^{2}-m_{a}^{2}\right]} \stackrel{\text { div }}{\longrightarrow} \frac{-i}{(4 \pi)^{2} \varepsilon} Q_{I+} P_{L} \gamma^{\rho} \gamma^{\nu} \gamma_{\rho} P_{R} \\
& -4 Q_{\mu}-\int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} \not k\left(\not p_{e}+m_{\mu}\right) \gamma^{\nu}}{\left[p_{e}^{2}-m_{\mu}^{2}\right]\left[\left(p_{e}-k\right)^{2}-M_{S}^{2}\right]\left[k^{2}-m_{a}^{2}\right]} \stackrel{\text { div }}{\xrightarrow{(4 \pi)^{2} \varepsilon}} \frac{2 i}{Q_{\mu}-}{m_{\mu}^{2}}_{P_{L} \not 中_{e}}\left(\not p_{e}+m_{\mu}\right) \gamma^{\nu} \\
& 4 Q_{e}-\int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{\gamma^{\nu} \not p_{\mu} P_{L} k}{\left[p_{\mu}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]\left[k^{2}-m_{a}^{2}\right]} \stackrel{\text { div }}{\longrightarrow} \frac{-2 i}{(4 \pi)^{2} \varepsilon} \frac{Q_{e}-}{m_{\mu}^{2}} \gamma^{\nu} \not p_{\mu} P_{L} \not p_{\mu}
\end{aligned}
$$

## Photonic Contribution: Cross Check via UV Divergences

 In form of $i \mathcal{M}=e f_{e a}^{*} f_{a \mu} A_{\nu}\left(q^{\prime}\right) \bar{u}_{e}\left(p_{e}\right) \mathcal{I}^{\nu} u_{\mu}\left(p_{\mu}\right)$ :

$$
\begin{aligned}
& -4 Q_{S} \int \frac{\mathrm{~d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} \not k\left(2 p_{\mu}-2 k+q^{\prime}\right)^{\nu}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k+q^{\prime}\right)^{2}-M_{S}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]} \stackrel{\text { div }}{\longrightarrow} \frac{2 i}{(4 \pi)^{2} \varepsilon} Q_{S} P_{L} \gamma^{\nu} \\
& -4 Q_{I}+\int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L}\left(\not k+\not q^{\prime}+m_{a}\right) \gamma^{\nu}\left(\not k+m_{a}\right) P_{R}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]\left[\left(k+q^{\prime}\right)^{2}-m_{a}^{2}\right]} \stackrel{\operatorname{div}}{\longrightarrow} \frac{-i}{(4 \pi)^{2} \varepsilon} Q_{I}+P_{L} \gamma^{\rho} \gamma^{\nu} \gamma_{\rho} P_{R} \\
& -4 Q_{\mu}-\int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} \not k\left(\not p_{e}+m_{\mu}\right) \gamma^{\nu}}{\left[p_{e}^{2}-m_{\mu}^{2}\right]\left[\left(p_{e}-k\right)^{2}-M_{S}^{2}\right]\left[k^{2}-m_{a}^{2}\right]} \stackrel{\operatorname{div}}{\longrightarrow} \frac{2 i}{(4 \pi)^{2} \varepsilon} \frac{Q_{\mu}-}{m_{\mu}^{2}} P_{L} \not p_{e}\left(\not p_{e}+m_{\mu}\right) \gamma^{\nu} \\
& 4 Q_{e}-\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\gamma^{\nu} \not p_{\mu} P_{L} \not k}{\left[p_{\mu}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]\left[k^{2}-m_{a}^{2}\right]} \stackrel{\operatorname{div}}{\longrightarrow} \frac{-2 i}{(4 \pi)^{2} \varepsilon} \frac{Q_{e}-}{m_{\mu}^{2}} \gamma^{\nu} \not p_{\mu} P_{L} \not p_{\mu} \\
& \Rightarrow \sum \mathcal{I}^{\nu}=\frac{i}{(4 \pi)^{2} \varepsilon}\left[\left(2 Q_{S}+2 Q_{1+}-Q_{e}-Q_{\mu}-\right) P_{L} \gamma^{\nu}\right]=0
\end{aligned}
$$

## Photonic Contribution: Cross Check via UV Divergences

 In form of $i \mathcal{M}=e f_{e a}^{*} f_{a \mu} A_{\nu}\left(q^{\prime}\right) \bar{u}_{e}\left(p_{e}\right) \mathcal{I}^{\nu} u_{\mu}\left(p_{\mu}\right)$ :

$$
\begin{aligned}
& -4 Q_{S} \int \frac{\mathrm{~d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} k\left(2 p_{\mu}-2 k+q^{\prime}\right)^{\nu}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k+q^{\prime}\right)^{2}-M_{S}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]} \stackrel{\text { div }}{\longrightarrow} \frac{2 i}{(4 \pi)^{2} \varepsilon} Q_{S} P_{L} \gamma^{\nu} \\
& -4 Q_{I+} \int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L}\left(k+\phi^{\prime}+m_{a}\right) \gamma^{\nu}\left(k+m_{a}\right) P_{R}}{\left[k^{2}-m_{a}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]\left[\left(k+q^{\prime}\right)^{2}-m_{a}^{2}\right]} \stackrel{\operatorname{div}}{\xrightarrow{(4 \pi)^{2} \varepsilon}} \frac{-i}{I+} P_{L} \gamma^{\rho} \gamma^{\nu} \gamma_{\rho} P_{R} \\
& -4 Q_{\mu}-\int \frac{\mathrm{d}^{d} k}{(2 \pi)^{d}} \frac{P_{L} k\left(\phi_{e}+m_{\mu}\right) \gamma^{\nu}}{\left[p_{e}^{2}-m_{\mu}^{2}\right]\left[\left(p_{e}-k\right)^{2}-M_{S}^{2}\right]\left[k^{2}-m_{a}^{2}\right]} \xrightarrow{\text { div }} \frac{2 i}{(4 \pi)^{2} \varepsilon} \frac{Q_{\mu}-}{m_{\mu}^{2}} P_{L \phi_{e}}\left(\phi_{e}+m_{\mu}\right) \gamma^{\nu} \\
& 4 Q_{e}-\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\gamma^{\nu} \phi_{\mu} P_{L} k}{\left[p_{\mu}^{2}\right]\left[\left(p_{\mu}-k\right)^{2}-M_{S}^{2}\right]\left[k^{2}-m_{a}^{2}\right]} \xrightarrow{\text { div }} \frac{-2 i}{(4 \pi)^{2} \varepsilon} \frac{Q_{e}-}{m_{\mu}^{2}} \gamma^{\nu} \phi_{\mu} P_{L \phi_{\mu}} \\
& \Rightarrow \Sigma \mathcal{I}^{\nu}=\frac{i}{(4 \pi)^{2} \varepsilon}\left[\left(2 Q_{S}+2 Q_{H_{+}}-Q_{e^{-}}-Q_{\mu^{-}}\right) P_{L} \gamma^{\nu}\right]=0
\end{aligned}
$$

## Photonic Contribution: Results I

Determine form factors with help of Mathematica package Package-X (Patel, arXiv:1503.01469):

$$
\begin{aligned}
& F_{1}\left(-m_{\mu}^{2}\right)=G_{1}\left(-m_{\mu}^{2}\right)= \\
& =-\frac{1}{128 \pi^{2} m_{\mu}^{4}} \sum_{a=e, \mu, \tau} f_{e a}^{*} f_{a \mu}\left[2 m_{\mu}^{2}\left(-5 m_{a}^{2}+6 m_{\mu}^{2}+5 M_{S}^{2}\right)-2 S_{a} m_{\mu}^{2}\left(m_{a}^{2}+3 m_{\mu}^{2}-M_{S}^{2}\right)\right. \\
& \ln \left[\frac{2 m_{a}^{2}}{2 m_{a}^{2}+m_{\mu}^{2}\left(1+S_{a}\right)}\right]+4 S_{S} m_{\mu}^{2}\left(m_{a}^{2}+m_{\mu}^{2}-M_{S}^{2}\right) \ln \left[\frac{2 M_{S}^{2}}{2 M_{S}^{2}+m_{\mu}^{2}\left(1+S_{S}\right)}\right]+\left(3 m _ { a } ^ { 2 } \left(2 m_{a}^{2}-m_{\mu}^{2}\right.\right. \\
& \left.\left.-4 M_{S}^{2}\right)+5 m_{\mu}^{4}-7 m_{\mu}^{2} M_{S}^{2}+6 M_{S}^{4}\right) \ln \left[\frac{m_{a}^{2}}{M_{S}^{2}}\right]+2 T_{a}\left(-6 m_{a}^{2}+m_{\mu}^{2}+6 M_{S}^{2}\right) \ln \left[\frac{2 m_{2} M_{S}}{m_{a}^{2}-m_{\mu}^{2}+M_{S}^{2}-T_{a}}\right] \\
& +2 m_{\mu}^{2}\left[\left(m_{a}^{4}+8 m_{a}^{2} m_{\mu}^{2}+M_{S}^{4}-2 M_{S}^{2}\left(m_{a}^{2}+2 m_{\mu}^{2}\right)\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; m_{a}, M_{S}, m_{a}\right]\right. \\
& \left.\left.+2\left(m_{a}^{4}-2 M_{S}^{2}\left(m_{a}^{2}-2 m_{\mu}^{2}\right)+M_{S}^{4}\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; M_{S}, m_{a}, M_{S}\right]\right]\right]
\end{aligned}
$$

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\begin{aligned}
& \mathrm{F}_{1}\left(-m_{\mu}^{2}\right)=\mathrm{G}_{1}\left(-m_{\mu}^{2}\right)= \\
& =-\frac{1}{128 \pi^{2} m_{\mu}^{4}} \sum_{a=e, \mu, \tau} f_{e a}^{*} f_{a \mu}\left[2 m_{\mu}^{2}\left(-5 m_{a}^{2}+6 m_{\mu}^{2}+5 M_{S}^{2}\right)-2 S_{a} m_{\mu}^{2}\left(m_{a}^{2}+3 m_{\mu}^{2}-M_{S}^{2}\right)\right. \\
& \ln \left[\frac{2 m_{2}^{2}}{2 m_{a}^{2}+m_{\mu}^{2}\left(1+S_{a}\right)}\right]+4 S_{S} m_{\mu}^{2}\left(m_{a}^{2}+m_{\mu}^{2}-M_{S}^{2}\right) \ln \left[\frac{2 M_{S}^{S}}{2 M_{S}^{2}+m_{\mu}^{2}\left(1+S_{S}\right)}\right]+\left(3 m _ { a } ^ { 2 } \left(2 m_{a}^{2}-m_{\mu}^{2}\right.\right. \\
& \left.\left.-4 M_{S}^{2}\right)+5 m_{\mu}^{4}-7 m_{\mu}^{2} M_{S}^{2}+6 M_{S}^{4}\right) \ln \left[\frac{m_{a}^{2}}{M_{S}^{2}}\right]+2 T_{a}\left(-6 m_{a}^{2}+m_{\mu}^{2}+6 M_{S}^{2}\right) \ln \left[\frac{2 m_{a} M_{S}}{m_{a}^{2}-m_{\mu}^{2}+M_{S}^{2}-T_{a}}\right] \\
& +2 m_{\mu}^{2}\left[\left(m_{a}^{4}+8 m_{a}^{2} m_{\mu}^{2}+M_{S}^{4}-2 M_{S}^{2}\left(m_{a}^{2}+2 m_{\mu}^{2}\right)\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; m_{a}, M_{S}, m_{a}\right]\right. \\
& \left.\left.+2\left(m_{a}^{4}-2 M_{S}^{2}\left(m_{a}^{2}-2 m_{\mu}^{2}\right)+M_{S}^{4}\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; M_{S}, m_{a}, M_{S}\right]\right]\right] \\
& \xrightarrow{M_{S} \gg m_{a}} \longrightarrow-f_{e a}^{*} f_{a \mu}\left[\frac{2 m_{a}^{2}+m_{\mu}^{2} \log \left(\frac{m_{a}}{S_{S}}\right)}{12 \pi^{2} M_{S}^{2}}+\frac{\sqrt{m_{\mu}^{2}+4 m_{2}^{2}}\left(m_{\mu}^{2}-2 m_{a}^{2}\right)}{12 \pi^{2} m_{\mu} M_{S}^{2}} \operatorname{Arctanh}\left(\frac{m_{\mu}}{\sqrt{m_{\mu}^{2}+4 m_{a}^{2}}}\right)\right]+\mathcal{O}\left(M_{S}^{-4}\right)
\end{aligned}
$$

## Photonic Contribution: Results I

Determine form factors with help of Mathematica package Package-X (Patel, arXiv:1503.01469):

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& \ln \left[\frac{2 m_{a}^{2}}{2 m_{2}^{2}+m_{\mu}^{2}\left(1+S_{a}\right)}\right]+4 S_{S} m_{\mu}^{2}\left(m_{a}^{2}+m_{\mu}^{2}-M_{S}^{2}\right) \ln \left[\frac{2 M_{S}^{2}}{2 M_{S}^{2}+m_{\mu}^{2}\left(1+S_{S}\right)}\right]+\left(3 m _ { a } ^ { 2 } \left(2 m_{a}^{2}-m_{\mu}^{2}\right.\right. \\
& \left.\left.-4 M_{S}^{2}\right)+5 m_{\mu}^{4}-7 m_{\mu}^{2} M_{S}^{2}+6 M_{S}^{4}\right) \ln \left[\frac{m_{a}^{2}}{M_{S}^{2}}\right]+2 T_{a}\left(-6 m_{a}^{2}+m_{\mu}^{2}+6 M_{S}^{2}\right) \ln \left[\frac{2 m_{2} M_{S}}{m_{a}^{2}-m_{\mu}^{2}+M_{S}^{2}-T_{a}}\right] \\
& +2 m_{\mu}^{2}\left[\left(m_{a}^{4}+8 m_{a}^{2} m_{\mu}^{2}+M_{S}^{4}-2 M_{S}^{2}\left(m_{a}^{2}+2 m_{\mu}^{2}\right)\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; m_{a}, M_{S}, m_{a}\right]\right. \\
& \left.\left.+2\left(m_{a}^{4}-2 M_{S}^{2}\left(m_{a}^{2}-2 m_{\mu}^{2}\right)+M_{S}^{4}\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; M_{S}, m_{a}, M_{S}\right]\right]\right]
\end{aligned}
$$

$$
\xrightarrow{M_{s} \gg m_{a}}-f_{e a}^{*} f_{a \mu}\left[\frac{2 m_{a}^{2}+m_{\mu}^{2} \log \left(\frac{m_{a}}{M_{S}}\right)}{12 \pi^{2} M_{S}^{2}}+\frac{\sqrt{m_{\mu}^{2}+4 m_{2}^{2}}\left(m_{\mu}^{2}-2 m_{a}^{2}\right)}{12 \pi^{2} m_{\mu} M_{s}^{2}} \operatorname{Arctanh}\left(\frac{m_{\mu}}{\sqrt{m_{\mu}^{2}+4 m_{a}^{2}}}\right)\right]+\mathcal{O}\left(M_{S}^{-4}\right)
$$

Note: $\mathcal{O}\left(M_{S}^{-4}\right)$ gives corrections of up to a few per cent

## Photonic Contribution: Results I

Determine form factors with help of Mathematica package Package-X (Patel, arXiv:1503.01469):

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\begin{aligned}
& F_{2}\left(-m_{\mu}^{2}\right)=-G_{2}\left(-m_{\mu}^{2}\right)= \\
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& \ln \left[\frac{2 m_{2}^{2}}{2 m_{a}^{2}+m_{\mu}^{2}\left(1+S_{a}\right)}\right]+4 S_{S} m_{\mu}^{2}\left(-3 m_{a}^{2}+m_{\mu}^{2}+3 M_{S}^{2}\right) \ln \left[\frac{2 M_{S}^{2}}{2 M_{S}^{2}+m_{\mu}^{\prime}\left(1+S_{S}\right)}\right] \\
& +\left(m_{a}^{2}\left(-2 m_{a}^{2}-7 m_{\mu}^{2}+4 M_{S}^{2}\right)+m_{\mu}^{4}+5 m_{\mu}^{2} M_{S}^{2}-2 M_{S}^{4}\right) \ln \left[\frac{m_{a}^{2}}{M_{S}^{2}}\right]+2 T_{a}\left(2 m_{a}^{2}-3 m_{\mu}^{2}-2 M_{S}^{2}\right) \\
& \ln \left[\frac{2 m_{a} M_{S}}{m_{a}^{2}-m_{\mu}^{2}+M_{S}^{2}-T_{a}}\right]+2 m_{\mu}^{2}\left[\left(-3 m_{a}^{4}-3 M_{S}^{4}+2 M_{S}^{2}\left(3 m_{a}^{2}+2 m_{\mu}^{2}\right)\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; m_{a}, M_{S}, m_{a}\right]\right. \\
& \left.\left.+2\left(-3 m_{a}^{4}+2 m_{a}^{2}\left(3 M_{S}^{2}+2 m_{\mu}^{2}\right)-3 M_{S}^{4}\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; M_{S}, m_{a}, M_{S}\right]\right]\right]
\end{aligned}
$$

## Photonic Contribution: Results I

Determine form factors with help of Mathematica package Package-X (Patel, arXiv:1503.01469):

$$
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& \mathrm{F}_{2}\left(-m_{\mu}^{2}\right)=-G_{2}\left(-m_{\mu}^{2}\right)= \\
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& \ln \left[\frac{2 m_{2}^{2}}{2 m_{a}^{2}+m_{\mu}^{2}\left(1+S_{a}\right)}\right]+4 S_{S} m_{\mu}^{2}\left(-3 m_{a}^{2}+m_{\mu}^{2}+3 M_{S}^{2}\right) \ln \left[\frac{2 M_{S}^{2}}{2 M_{S}^{2}+m_{\mu}^{2}\left(1+S_{S}\right)}\right] \\
& +\left(m_{a}^{2}\left(-2 m_{a}^{2}-7 m_{\mu}^{2}+4 M_{S}^{2}\right)+m_{\mu}^{4}+5 m_{\mu}^{2} M_{S}^{2}-2 M_{S}^{4}\right) \ln \left[\frac{m_{a}^{a}}{M_{S}^{2}}\right]+2 T_{a}\left(2 m_{a}^{2}-3 m_{\mu}^{2}-2 M_{S}^{2}\right) \\
& \ln \left[\frac{2 m_{a} M_{S}}{m_{a}^{2}-m_{\mu}^{2}+M_{S}^{2}-T_{a}}\right]+2 m_{\mu}^{2}\left[\left(-3 m_{a}^{4}-3 M_{S}^{4}+2 M_{S}^{2}\left(3 m_{a}^{2}+2 m_{\mu}^{2}\right)\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; m_{a}, M_{S}, m_{a}\right]\right. \\
& \left.\left.+2\left(-3 m_{a}^{4}+2 m_{a}^{2}\left(3 M_{S}^{2}+2 m_{\mu}^{2}\right)-3 M_{S}^{4}\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; M_{S}, m_{a}, M_{S}\right]\right]\right] \\
& \xrightarrow{M_{S} \gg m_{a}} f_{e a}^{*} f_{a \mu} \frac{m_{\mu}^{2}}{24 \pi^{2} M_{S}^{2}}+\mathcal{O}\left(M_{S}^{-4}\right)
\end{aligned}
$$

## Photonic Contribution: Results I

Determine form factors with help of Mathematica package Package-X (Patel, arXiv:1503.01469):

$$
\begin{aligned}
& \mathrm{F}_{2}\left(-m_{\mu}^{2}\right)=-G_{2}\left(-m_{\mu}^{2}\right)= \\
& =-\frac{1}{128 \pi^{2} m_{\mu}^{4}} \sum_{a=e, \mu, \tau} f_{e a}^{*} f_{a \mu}\left[2 m_{\mu}^{2}\left(-m_{a}^{2}+6 m_{\mu}^{2}+M_{S}^{2}\right)+2 S_{a} m_{\mu}^{2}\left(3 m_{a}^{2}+m_{\mu}^{2}-3 M_{S}^{2}\right)\right. \\
& \ln \left[\frac{2 m_{a}^{2}}{2 m_{a}^{2}+m_{\mu}^{2}\left(1+S_{a}\right)}\right]+4 S_{S} m_{\mu}^{2}\left(-3 m_{a}^{2}+m_{\mu}^{2}+3 M_{S}^{2}\right) \ln \left[\frac{2 M_{S}^{2}}{2 M_{S}^{2}+m_{\mu}^{2}\left(1+S_{S}\right)}\right] \\
& +\left(m_{a}^{2}\left(-2 m_{a}^{2}-7 m_{\mu}^{2}+4 M_{S}^{2}\right)+m_{\mu}^{4}+5 m_{\mu}^{2} M_{S}^{2}-2 M_{S}^{4}\right) \ln \left[\frac{m_{a}^{a}}{M_{S}^{2}}\right]+2 T_{a}\left(2 m_{a}^{2}-3 m_{\mu}^{2}-2 M_{S}^{2}\right) \\
& \ln \left[\frac{2 m_{a} M_{S}}{m_{a}^{2}-m_{\mu}^{2}+M_{S}^{2}-T_{a}}\right]+2 m_{\mu}^{2}\left[\left(-3 m_{a}^{4}-3 M_{S}^{4}+2 M_{S}^{2}\left(3 m_{a}^{2}+2 m_{\mu}^{2}\right)\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; m_{a}, M_{S}, m_{a}\right]\right. \\
& \left.\left.+2\left(-3 m_{a}^{4}+2 m_{a}^{2}\left(3 M_{S}^{2}+2 m_{\mu}^{2}\right)-3 M_{S}^{4}\right) C_{0}\left[0,-m_{\mu}^{2}, m_{\mu}^{2} ; M_{S}, m_{a}, M_{S}\right]\right]\right] \\
& \xrightarrow{M_{S} \gg m_{a}} f_{e a}^{*} f_{a \mu} \frac{m_{\mu}^{2}}{24 \pi^{2} M_{S}^{2}}+\mathcal{O}\left(M_{S}^{-4}\right)
\end{aligned}
$$

Note: $\mathcal{O}\left(M_{S}^{-4}\right)$ gives corrections of up to a few per cent

## 'Average Scenario' Couplings

|  | red | purple | blue |
| :---: | :---: | :---: | :---: |
| $f_{e e}$ | $10^{-16}$ | $10^{-15}$ | $10^{-1}$ |
| $f_{e \mu}$ | $10^{-2}$ | $10^{-3}$ | $10^{-4}$ |
| $f_{e \tau}$ | $10^{-19}$ | $10^{-2}$ | $10^{-2}$ |
| $f_{\mu \mu}$ | $10^{-4}$ | $10^{-3}$ | $10^{-3}$ |
| $f_{\mu \tau}$ | $10^{-5}$ | $10^{-4}$ | $10^{-4}$ |
| $f_{e e} f_{e \mu}$ | $10^{-18}$ | $10^{-18}$ | $\mathbf{1 0}^{-5}$ |
| $f_{e \mu} f_{\mu \mu}$ | $\mathbf{1 0}^{-6}$ | $\mathbf{1 0}^{-6}$ | $10^{-7}$ |
| $f_{e \tau} f_{\mu \tau}$ | $10^{-24}$ | $\mathbf{1 0}^{-6}$ | $10^{-6}$ |

Table: First part: 'average scenario' couplings for the benchmark points as extracted from Tab. 7 in King, Merle, Panizzi: arXiv:1406.4137. Second part: combination of couplings that enter the $\mu-e$ conversion amplitude. The bold values indicate the dominant photonic contribution.

## Non-Photonic Bands

- The amplitude that enters the non-photonic इ takes the form

$$
\mathcal{A} \propto\left|f_{e e}^{*} f_{e \mu} D\left(m_{e}\right)+f_{e \mu}^{*} f_{\mu \mu} D\left(m_{\mu}\right)+f_{e \tau}^{*} f_{\tau \mu} D\left(m_{\tau}\right)\right| .
$$

- The function $D\left(m_{a}\right)$ strongly varies with $m_{a}$. $\rightarrow$ dominant term stems from the tau propagating within the loop, i.e. $\boldsymbol{D}\left(\boldsymbol{m}_{\tau}\right)$
$\rightarrow$ exeeds the muon and electron contribution by three to four orders of magnitude
- blue/purple scenario: neither $f_{e e}^{*} f_{e \mu}$ nor $f_{e \mu}^{*} f_{\mu \mu}$ bypasses this difference + identic $\boldsymbol{f}_{\boldsymbol{e} \tau}^{*} \boldsymbol{f}_{\tau \mu}$ in both scenarios
$\rightarrow$ indistinguishable curves
- red/grey scenario:
dominant contributions: $f_{e \mu}^{*} f_{\mu \mu} D\left(m_{\mu}\right) \sim f_{e \tau}^{*} f_{\tau \mu} D\left(m_{\tau}\right)$
$\rightarrow$ same order of magnitude, i.e. comparable values of non-photonic contribution

