

Measuring the gravitational interaction
between matter and antimatter
(and some tests of CPT)

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What measurements are we talking about?

1) Measurement of the gravitational behavior of antimatter

tests of the Weak Equivalence Principle

2) Precise spectroscopic comparison between H and \bar{H}

tests of fundamental symmetry (CPT)

3) related measurements in antihydrogen(-like) systems

antiprotonic helium, positronium, protonium, ...

Gravity...

- General relativity is a classical (non quantum) theory;
- EEP violations may appear in some quantum theory
- New quantum scalar and vector fields are allowed in some models (Kaluza Klein)

Einstein field: tensor graviton (Spin 2, “Newtonian”)
+ Gravi-vector (spin 1)
+ Gravi-scalar (spin 0)

- These fields may mediate interactions violating the equivalence principle

M. Nieto and T. Goldman, Phys. Rep. 205, 5 221-281,(1992)

Scalar: “charge” of particle equal to “charge of antiparticle” : **attractive force**
Vector: “charge” of particle opposite to “charge of antiparticle”: **repulsive/attractive force**

$$V = - \frac{G_{\infty}}{r} m_1 m_2 \left(1 \mp a e^{-r/v} + b e^{-r/s} \right) \quad \text{Phys. Rev. D 33 (2475) (1986)}$$

Cancellation effects in matter experiment if $a \approx b$ and $v \approx s$

although CPT is part of the “standard model”,
the SM can be extended to allow CPT violation

CPT violation and the standard model

Phys. Rev. D 55, 6760–6774 (1997)

Don Colladay and V. Alan Kostelecký
Department of Physics, Indiana University, Bloomington, Indiana 47405
(Received 22 January 1997)

Modified Dirac eq. in SME

$$(i\gamma^\mu D_\mu - m_e - a_\mu^e \gamma^\mu - b_\mu^e \gamma_5 \gamma^\mu - \frac{1}{2} H_{\mu\nu}^e \sigma^{\mu\nu} + ic_{\mu\nu}^e \gamma^\mu D^\nu + id_{\mu\nu}^e \gamma_5 \gamma^\mu D^\nu) \psi = 0.$$

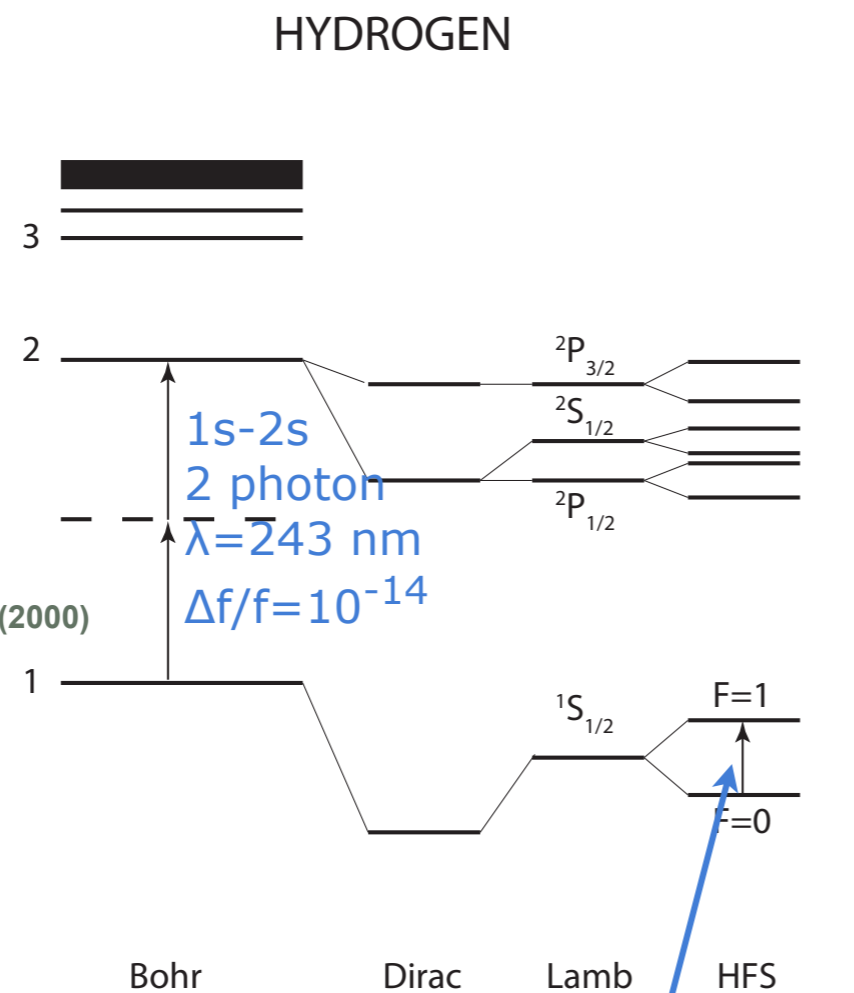
CPT & Lorentz violation

Lorentz violation

- Spontaneous Lorentz symmetry breaking by (exotic) string vacua
- Note: if there is a preferred frame, sidereal variation due to Earth's rotation might be detectable

Goal of comparative spectroscopy: test CPT symmetry

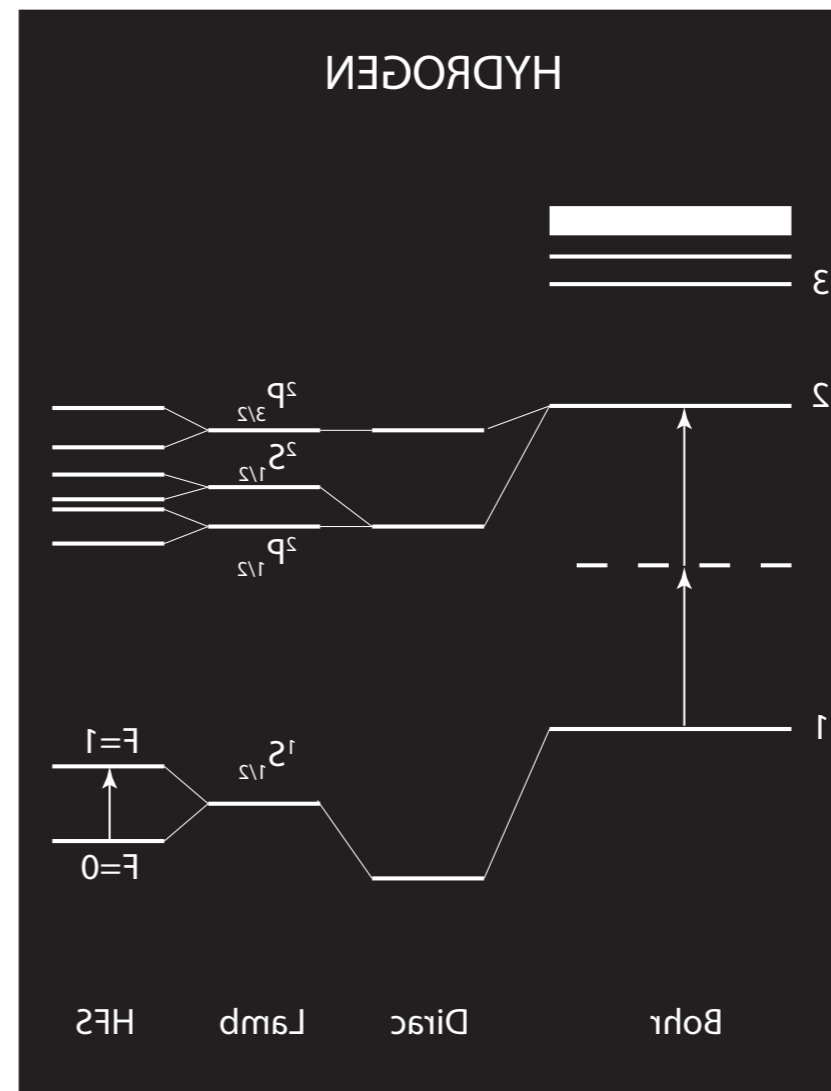
Hydrogen and Antihydrogen



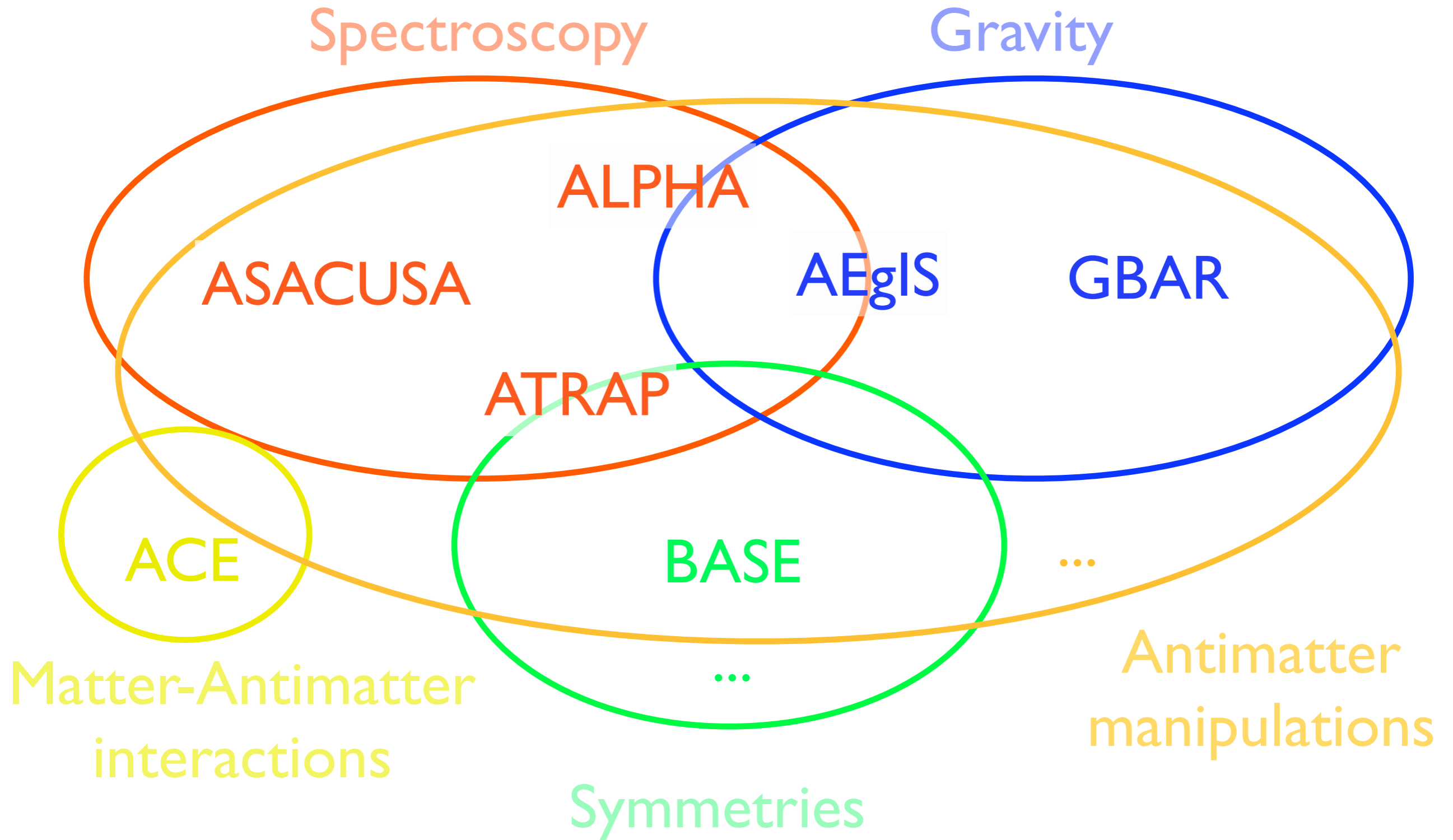
T. Hänsch et al.,
Phys. Rev. Lett. 84, 5496–5499 (2000)

N. F. Ramsey,
Physica Scripta T59, 323 (1995)

Ground state
hyperfine splitting
 $f = 1.4 \text{ GHz}$
 $\Delta f/f = 10^{-12}$

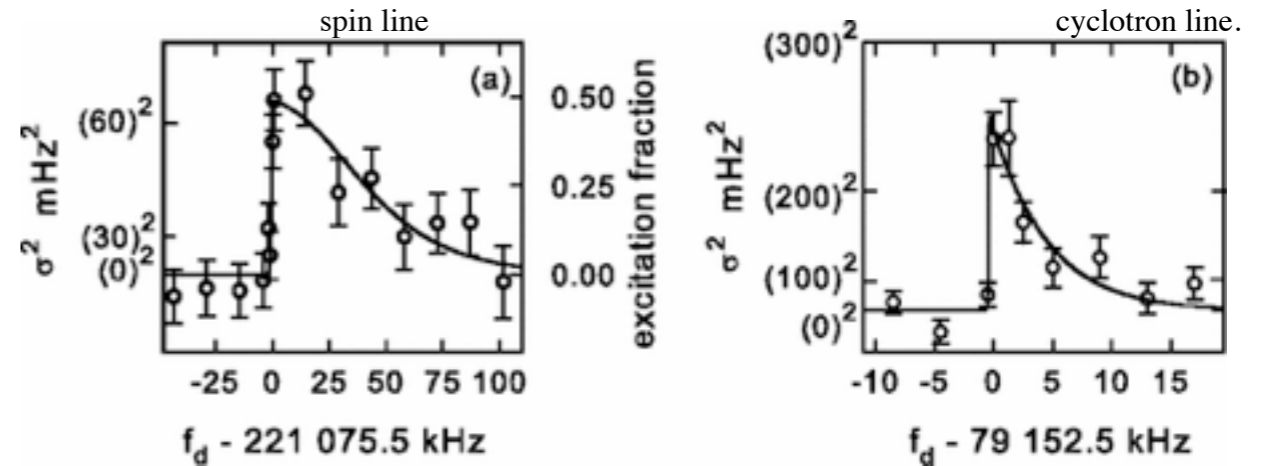
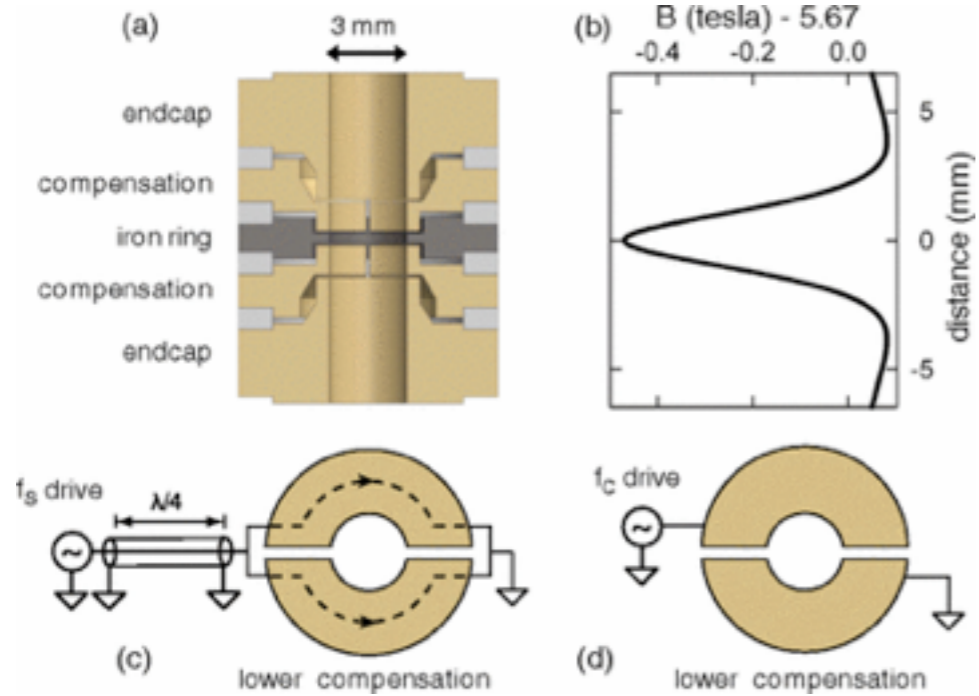


Experiments at the AD (antiprotons and antihydrogen)



ATRAP & BASE

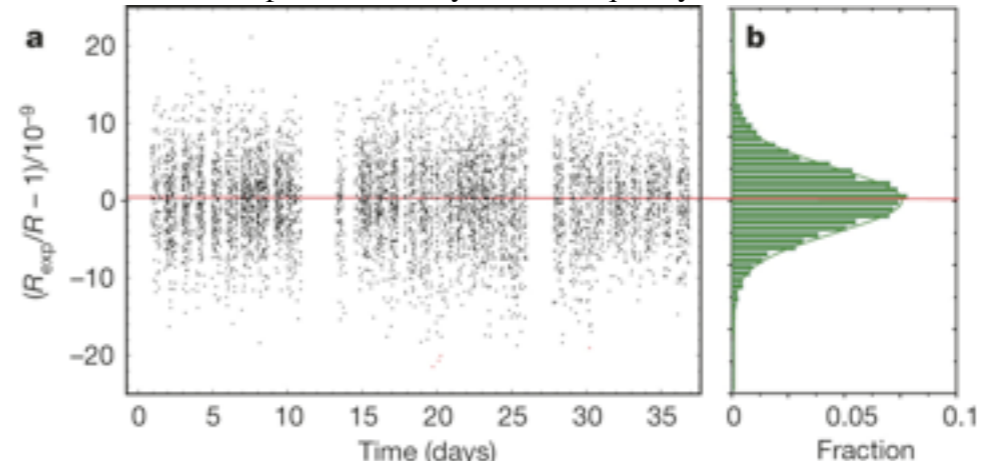
DiSciaccia, J. *et al.* One-particle measurement of the antiproton magnetic moment. Phys. Rev. Lett. 110, 130801 (2013)



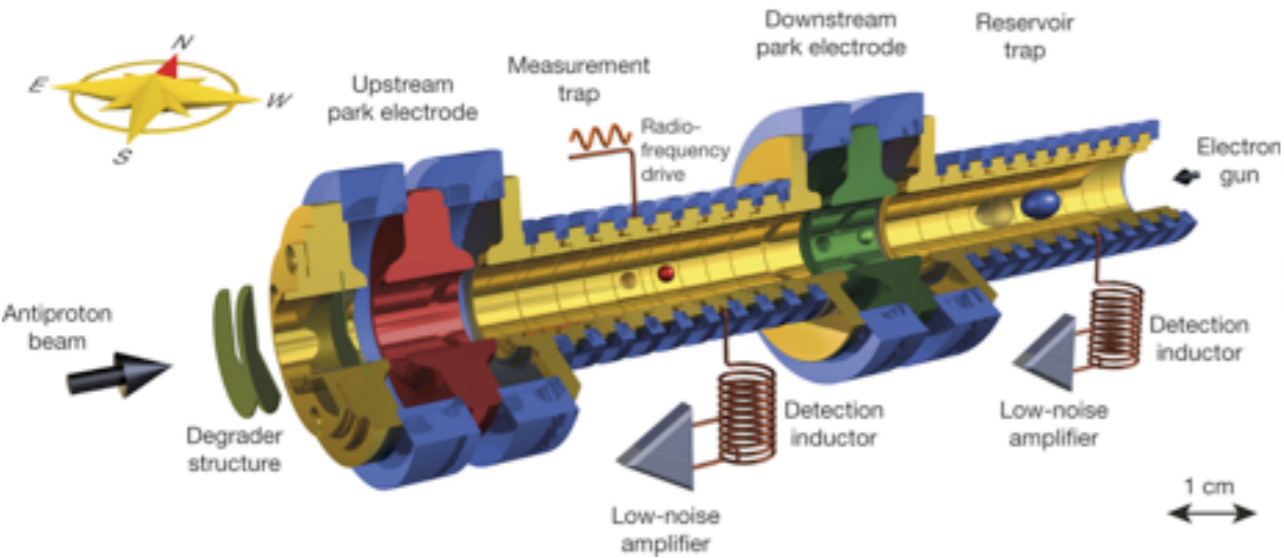
ATRAP: $\mu_{\bar{p}}/\mu_{\bar{p}} = -1.000\,000 \pm 0.000\,005$ (2013)

BASE: $(q/m)_{\bar{p}}/(q/m)_p - 1 = 1(69) \times 10^{-12}$ (2015)

All measured antiproton-to-H- cyclotron frequency ratios as a function of time

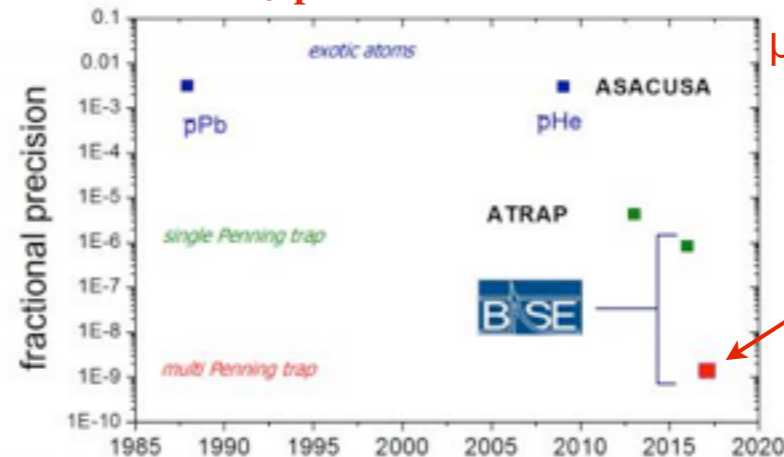


S. Ulmer. *et al.* Nature 524,196–199 (13 August 2015)



BASE: $\mu_{\bar{p}} = -2.792\,847\,344\,1(42)$ (2017)

$\mu_p = 2.792\,847\,350\,0(90)$

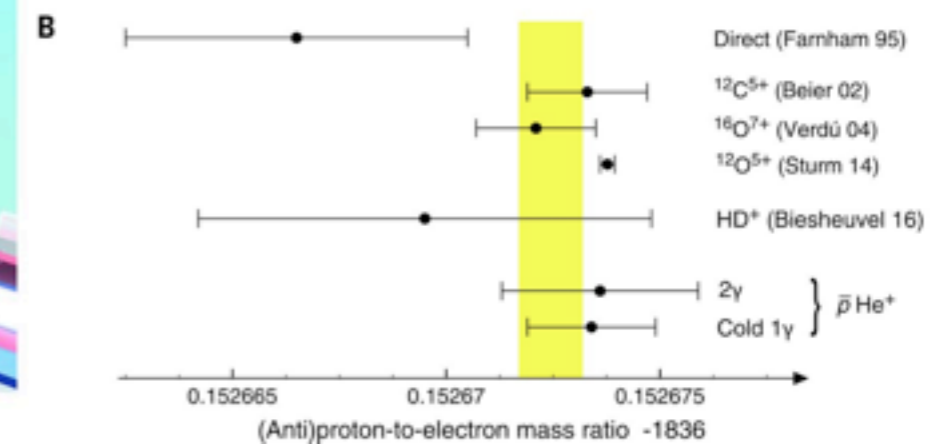
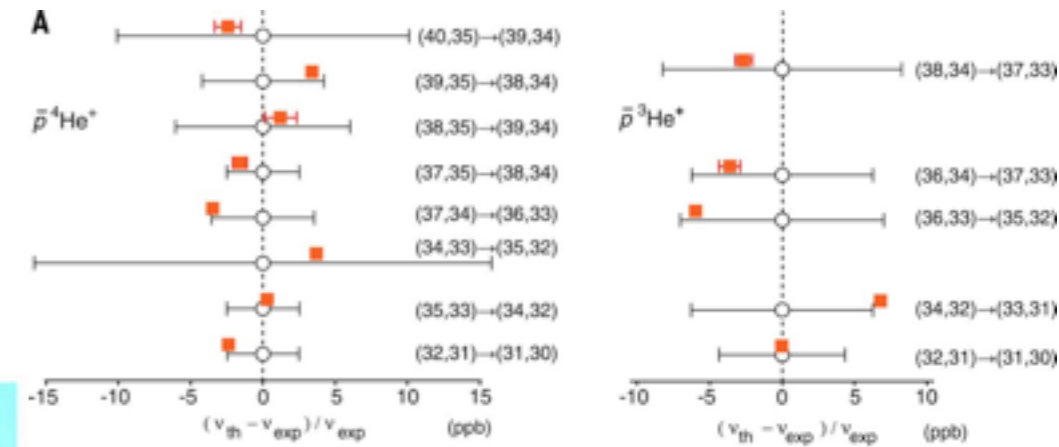
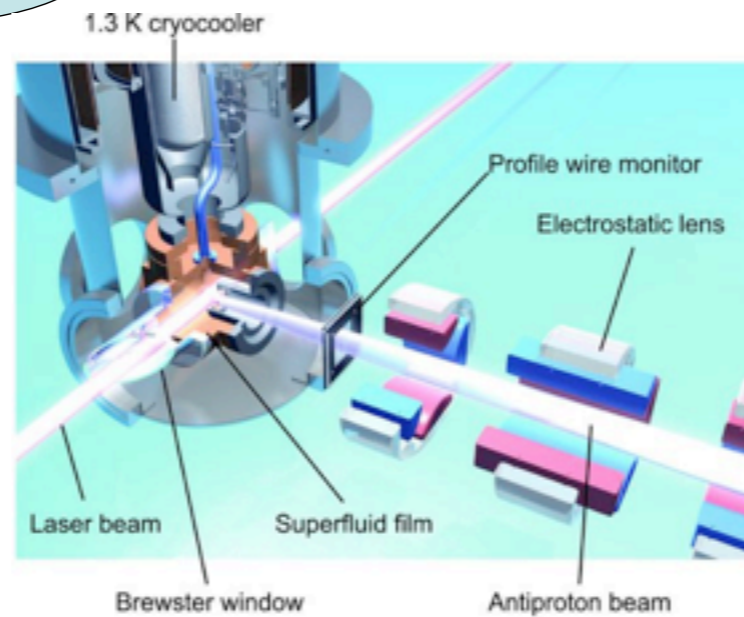
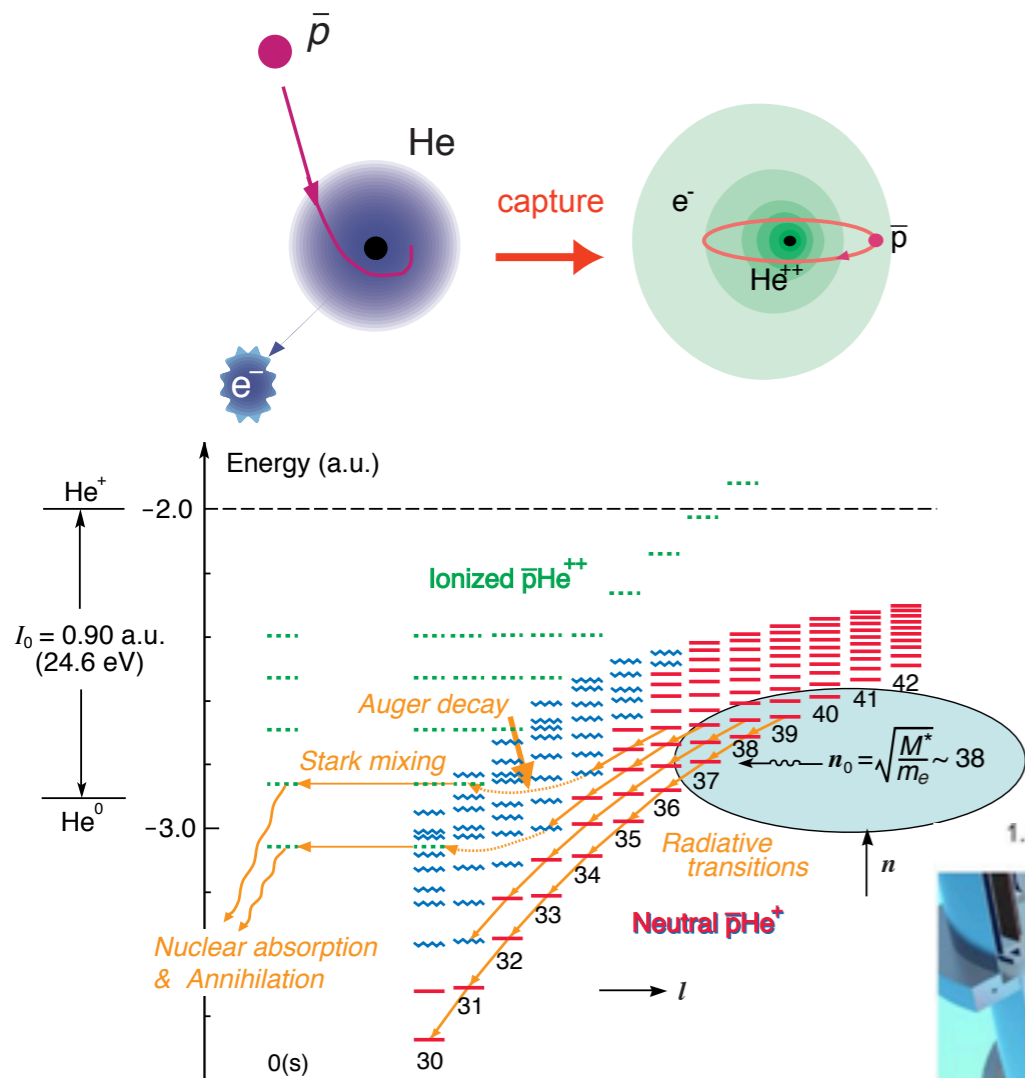


C. Smorra et al., BASE collaboration, Nature 550, 371–374 (19 October 2017)

ASACUSA results ($\bar{p}\text{He}^+$ spectroscopy)

By comparing the calculated and experimental $\bar{p}\text{He}^+$ frequencies, the ratio $M_{\bar{p}}/m_e$ can in principle be determined to a fractional precision of $< 1 \times 10^{-10}$

M. Hori *et al.*,
 Science 04 Nov 2016: Vol. 354, Issue 6312, pp. 610-614
 DOI: 10.1126/science.aaf6702



Combining with ATRAP/BASE:

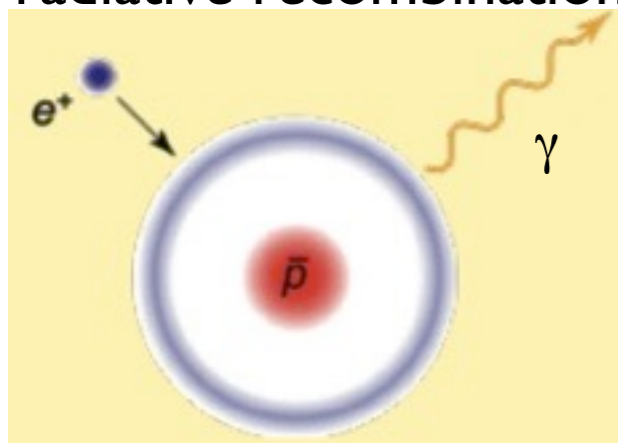
$$\Delta(m_{\bar{p}}, m_p), \Delta(q_{\bar{p}}, q_p) < 5 \times 10^{-10} \text{ (90\% CL)}$$

Antihydrogen production processes

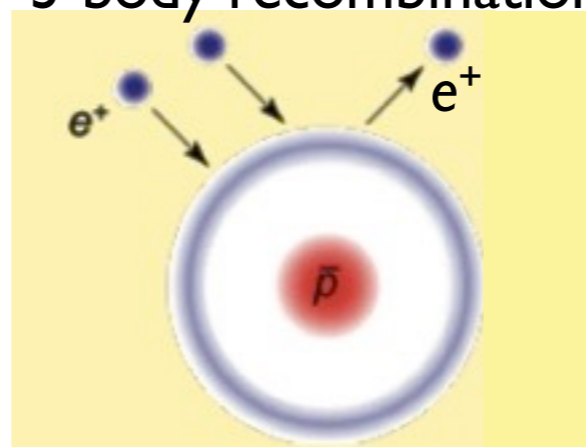
$10^5 \sim 10^7 \bar{p}$

$10^8 \sim 10^{10} e^+$

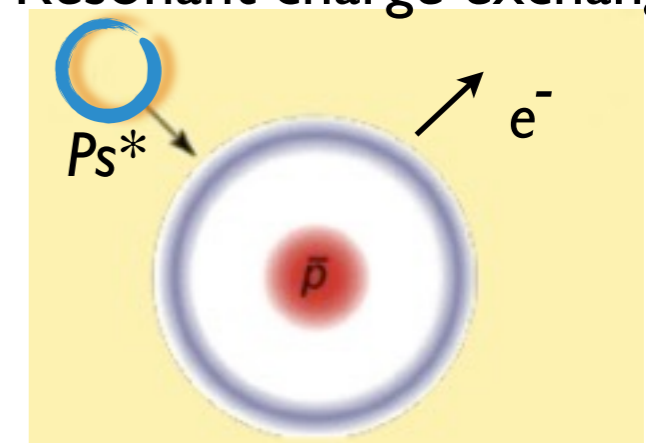
radiative recombination



TBR:
3-body recombination



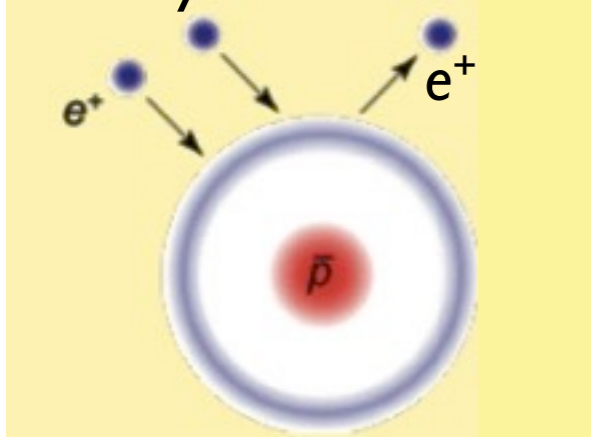
RCE:
Resonant charge exchange



very low rate

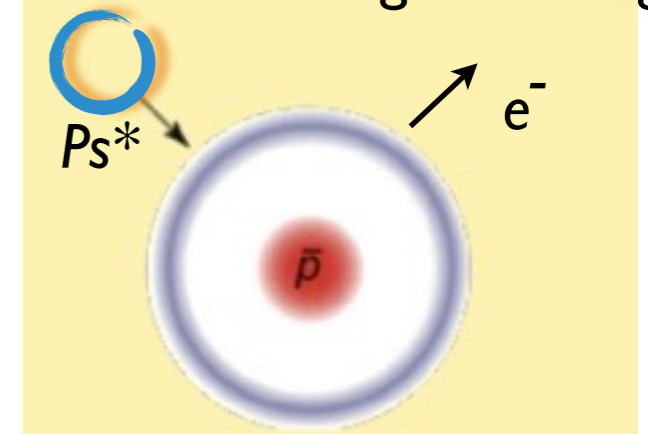
Antihydrogen production processes

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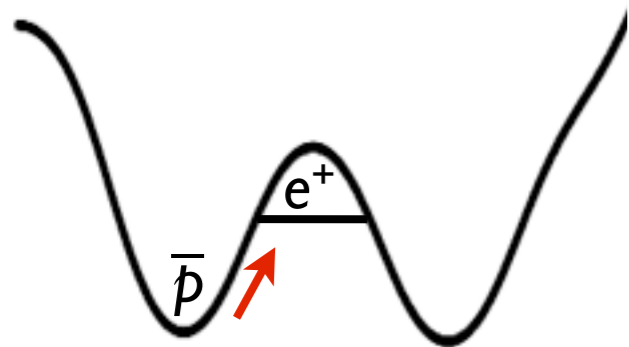


ALPHA
ATRAP
ASACUSA

RCE:
Resonant charge exchange



AEgIS
GBAR



Temperature
(T_{e^+})

Rate \sim Rate (trappable)

n (if trapped)

Temperature

$T_{\bar{p}}$

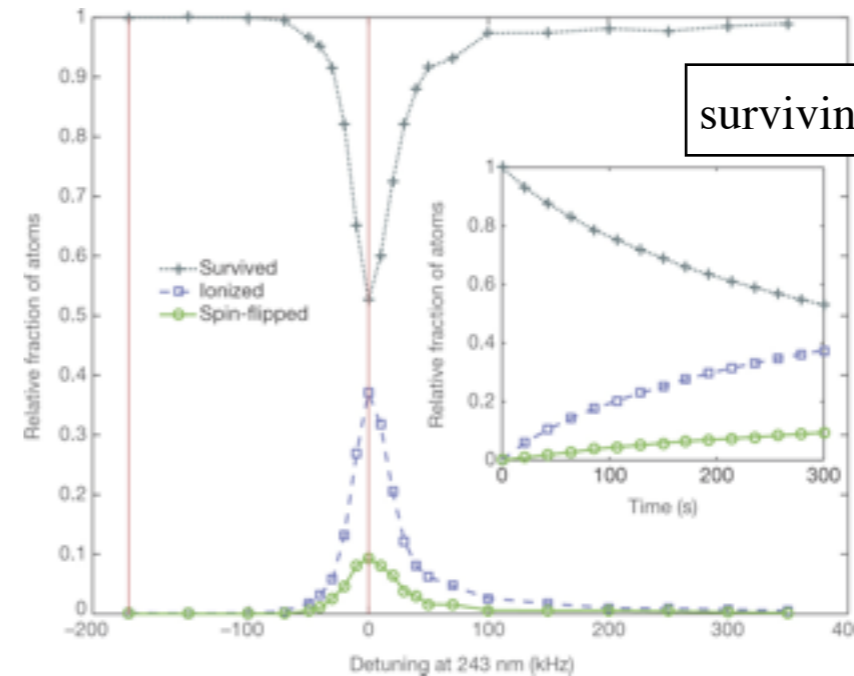
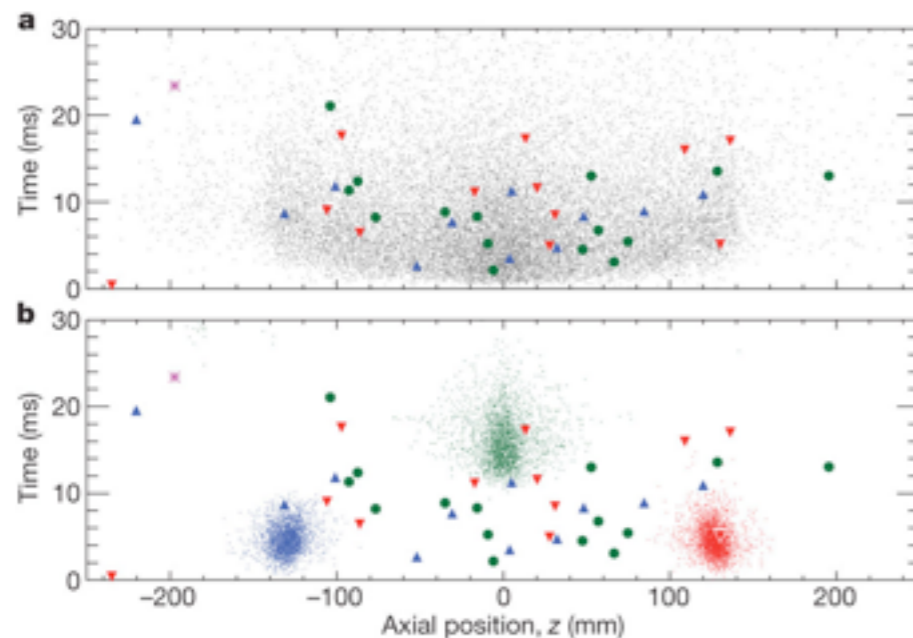
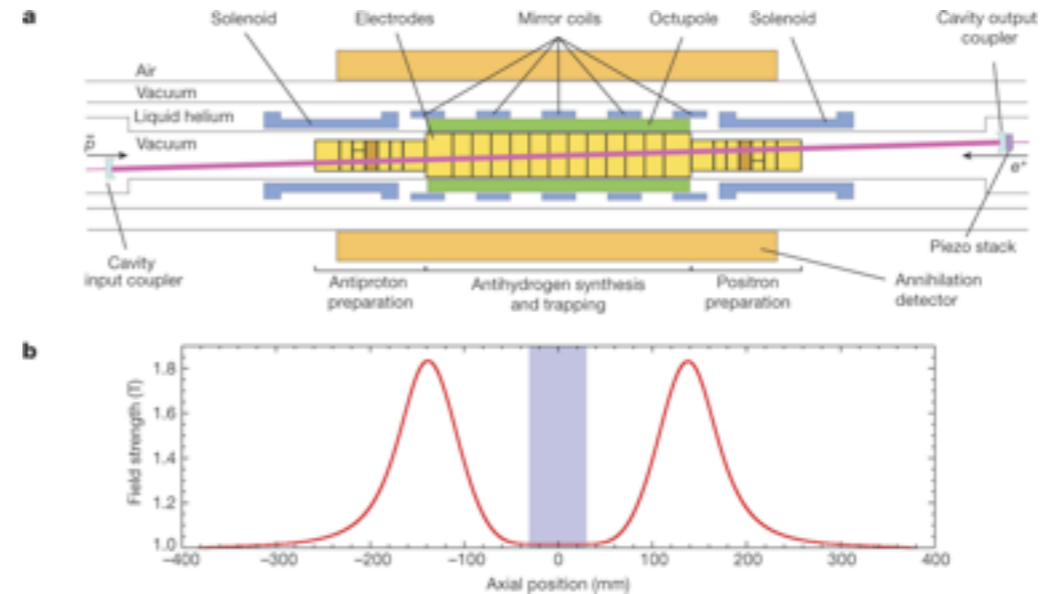
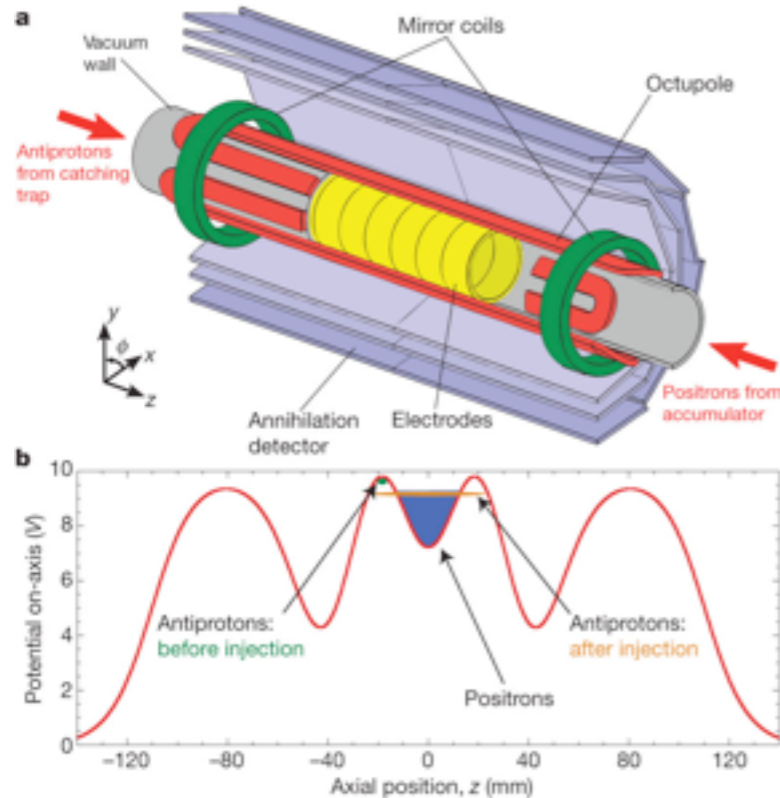
Rate \sim Rate (n_{P_s}, v_{P_s})

n (if trapped or slow)

ALPHA results (trapping, 1s-2s spectroscopy)

G. B. Andresen et al., Nature 468, 673–676 (02 December 2010)

M. Ahmadi et al., Nature 541, 506–510 (26 January 2017)



surviving fraction: $58\% \pm 6\%$

1s-2s to 10^{-10}

further results:

microwave transitions in GS \bar{H}
 $q(\bar{H}) < 0.71 \times 10^{-9} e$

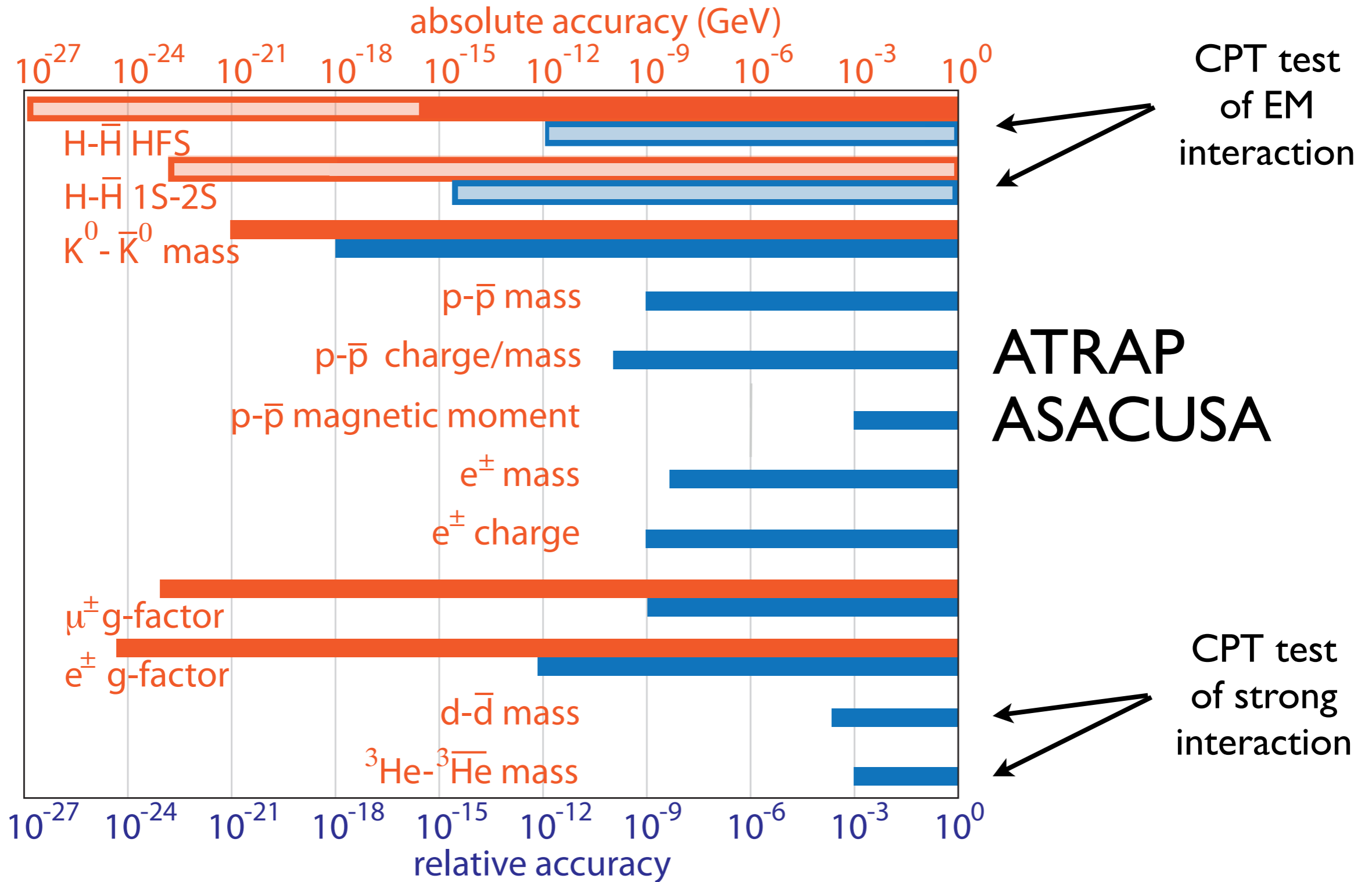
trapping of $\sim 10 \bar{H}$ simultaneously (similar for **ATRAP**)

$HFS_{\bar{H}} = 1,420.4 \pm 0.5 \text{ MHz}$
 M. Ahmadi et al., ALPHA collaboration,
 Nature 548, 66–69 (03 August 2017)

intermediate summary...

2013

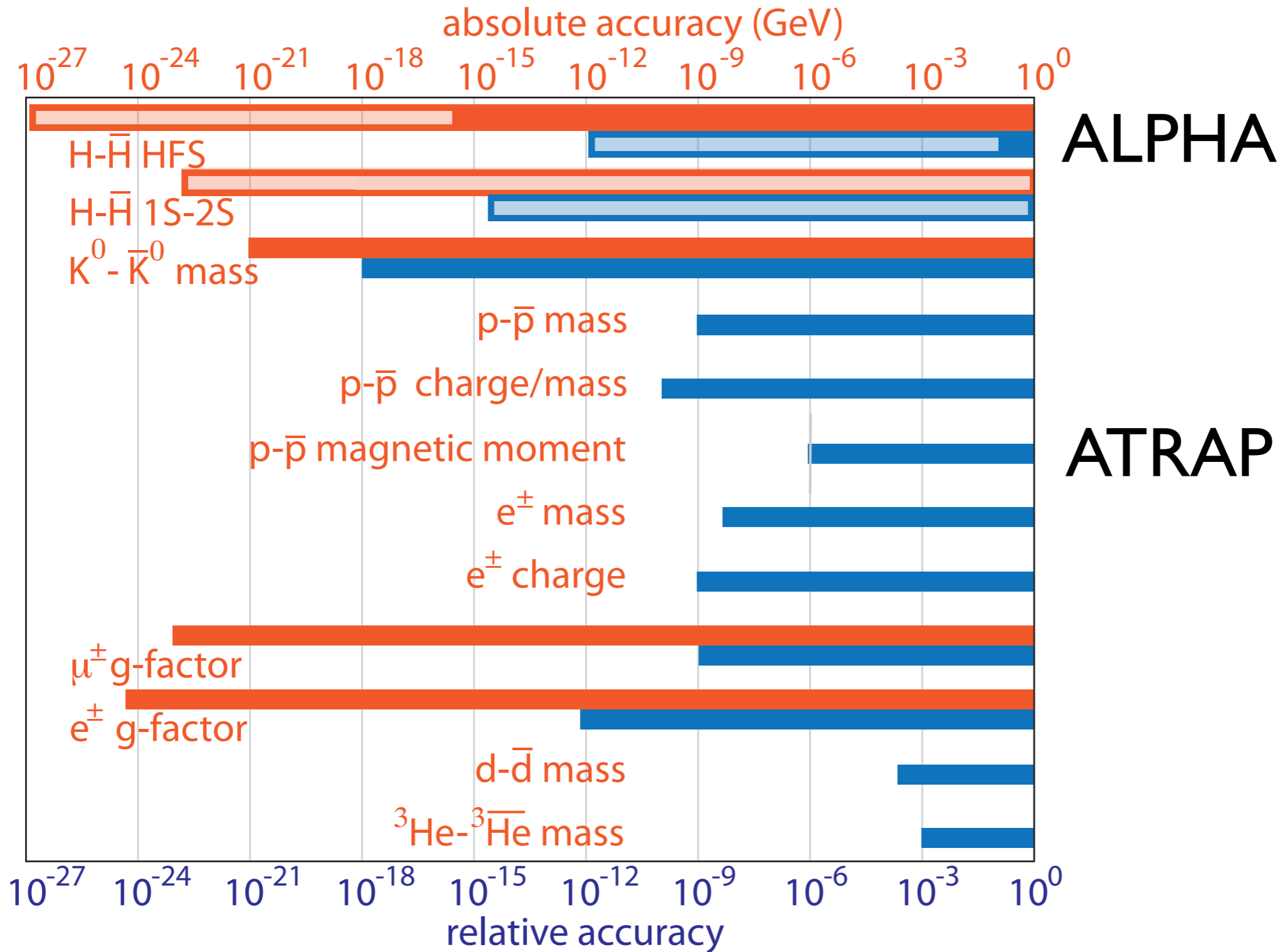
Motivation: CPT



Inconsistent definition of figure of merit: comparison difficult
 Pattern of CPT violation unknown (P: weak interaction; CP: mesons)
 Absolute energy scale: standard model extension (Kostelecky)

2015

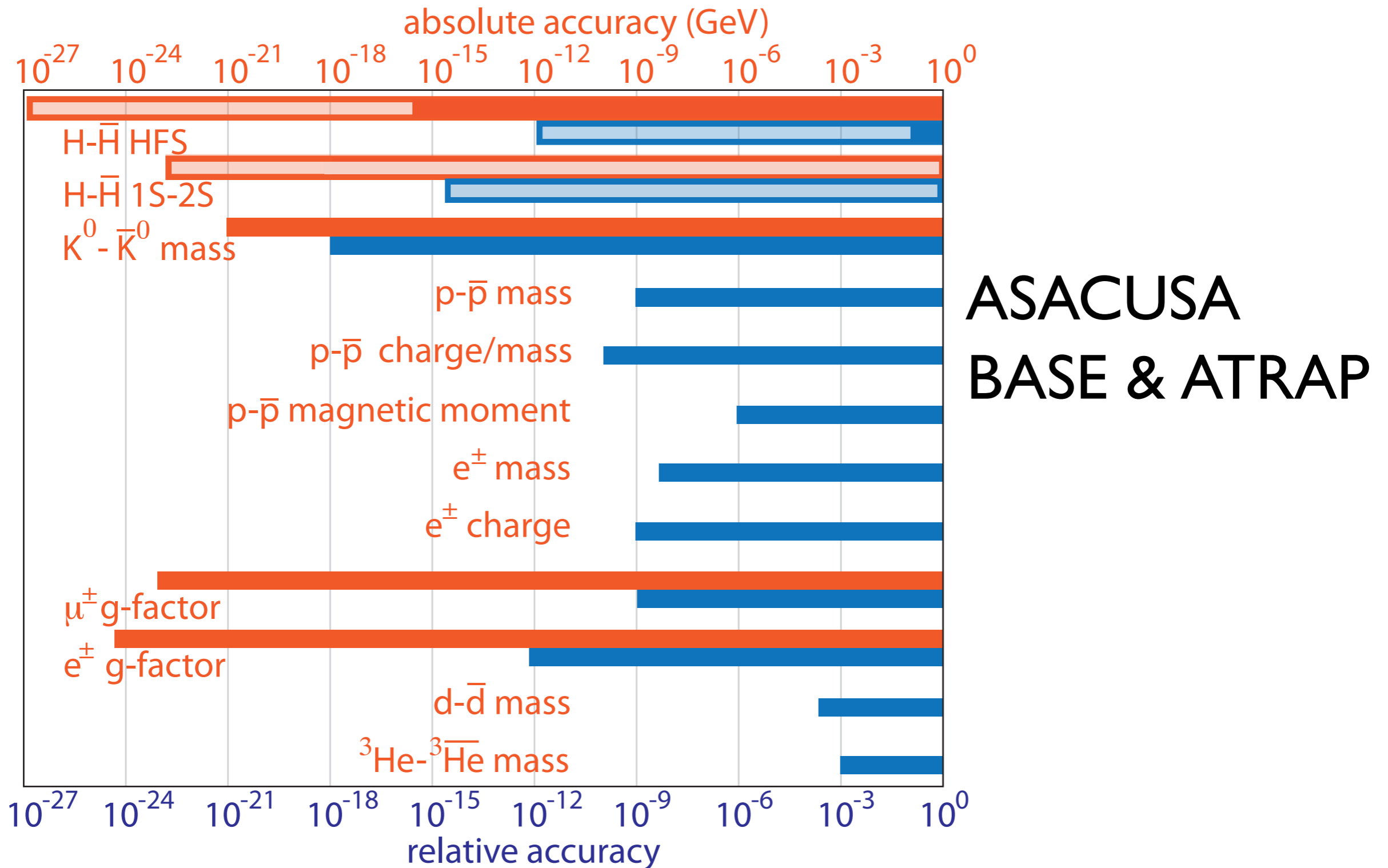
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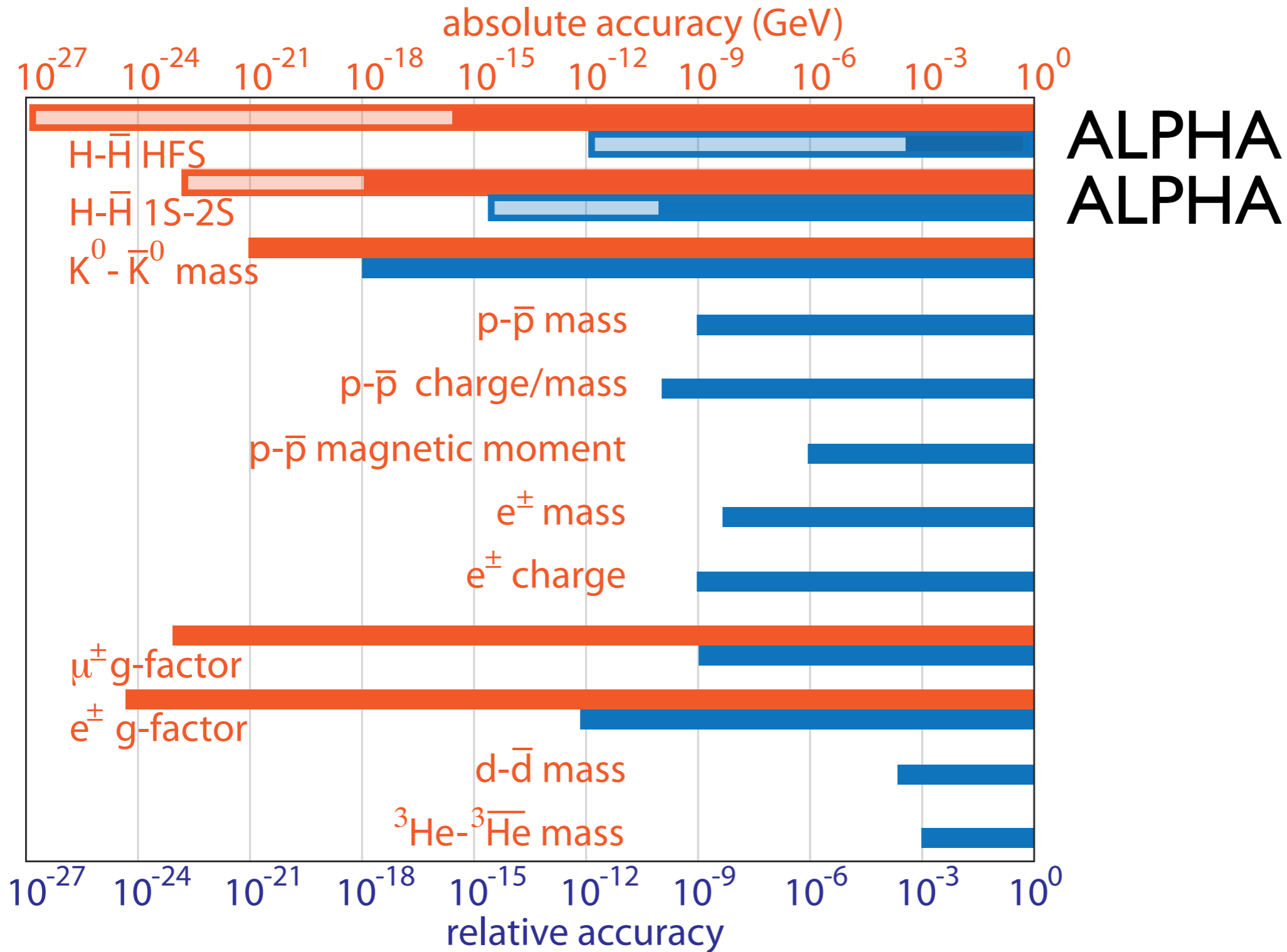
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2016

Motivation: CPT



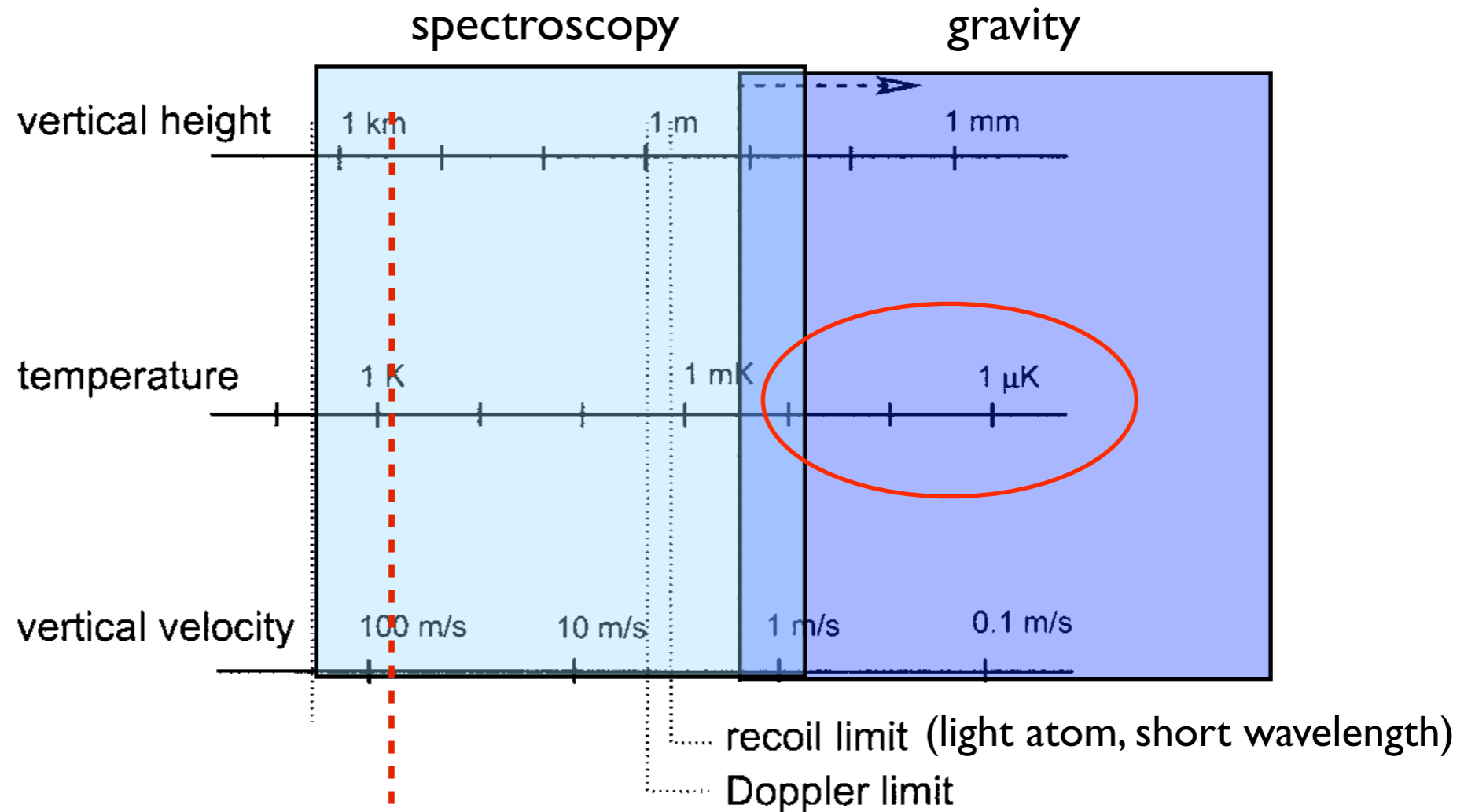
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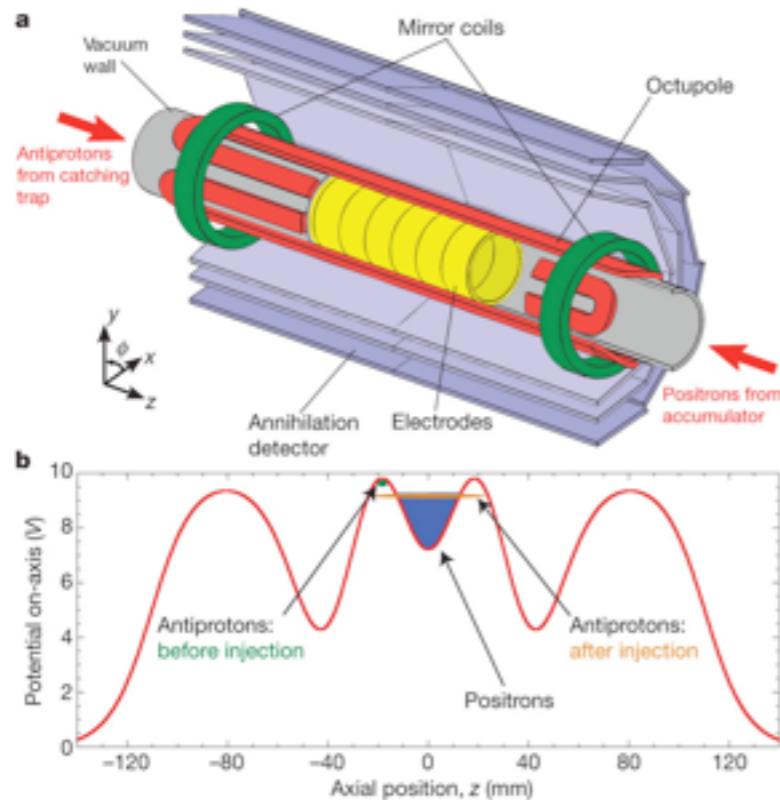
next stop: gravity

the importance of working at low temperature



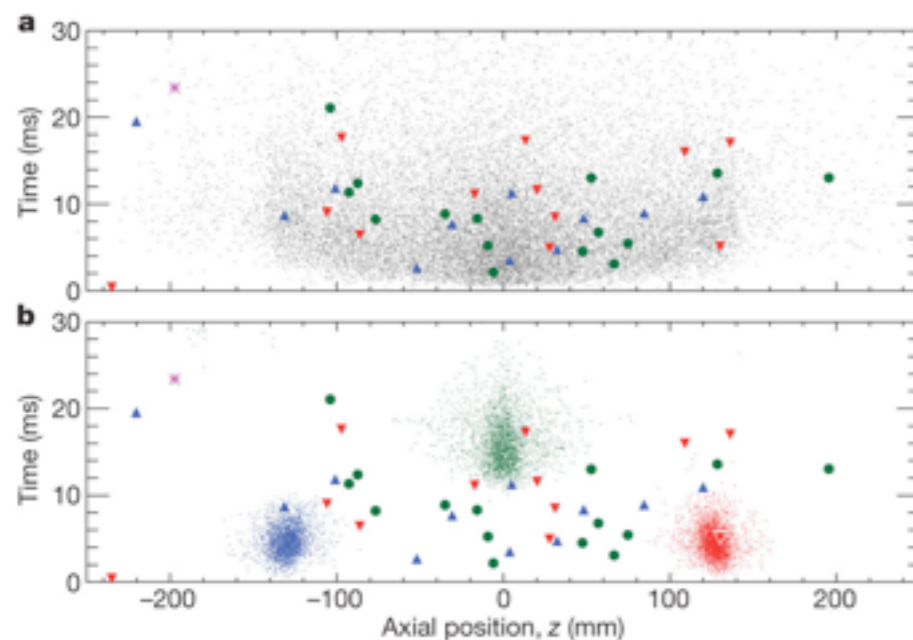
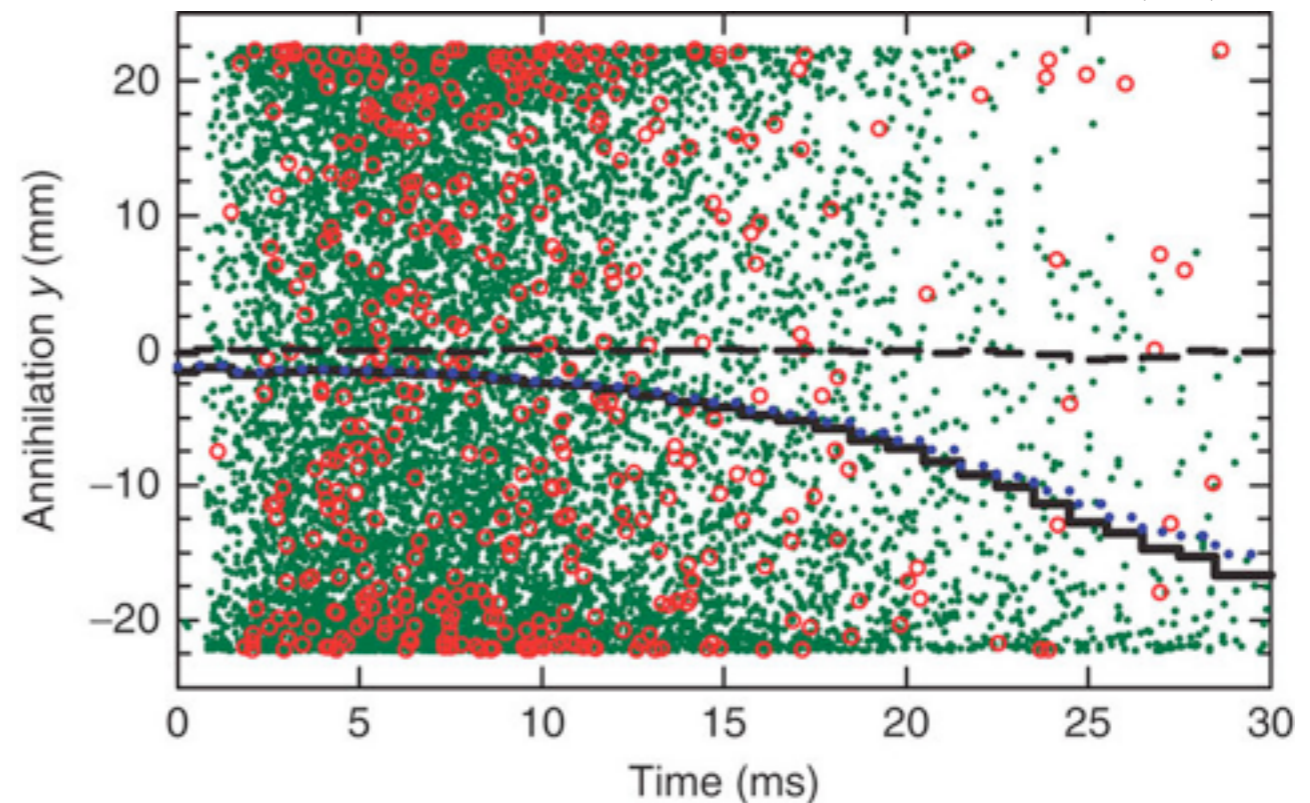
current lowest \bar{H}
temperature (0.5K)

ALPHA results (gravity at 0.5K)



$$F \equiv M_g/M$$

ALPHA collaboration, *Nature Communications* **4**, Article number: 1785 (2013)



$$F_{\bar{H}} < 110$$

“... cooling the anti-atoms, perhaps with lasers, to 30 mK or lower, and by lengthening the magnetic shutdown time constant to 300 ms, we would have the statistical power to measure gravity to the $F=\pm 1$ level ...”

2016

Temperature of produced \bar{H}
(critical for trapping, gravity measurements)

TBR: fraction trapped out of fraction made $\sim 10^{-4}$

challenge inherent in TBR: e^+ plasma physics \rightarrow
trade-off between # and temperature

possible increase in cold \bar{H} rate by laser-cooling Be^+
to sympathetically cool e^+ but is cooling efficient
enough to counteract heating through \bar{p} injection?

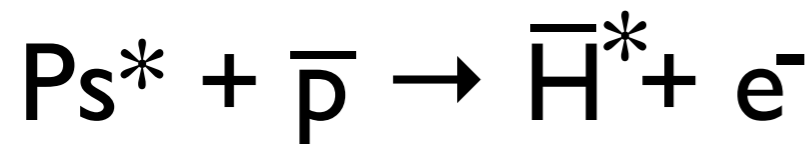
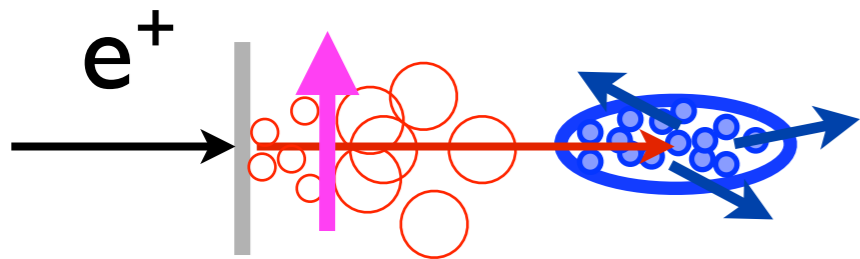
Outlook: 10's \sim 100's of trapped \bar{H} (through stacking)

alternative antihydrogen production method: RCE

$$T_{\bar{H}} \sim T_{\bar{p}}$$

AEgIS

$$T_{Ps} \sim 100 \text{ K}$$

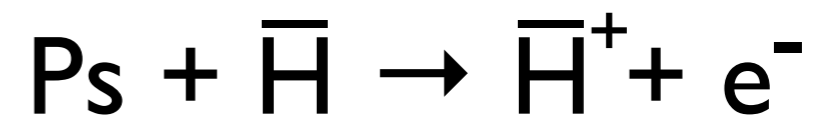
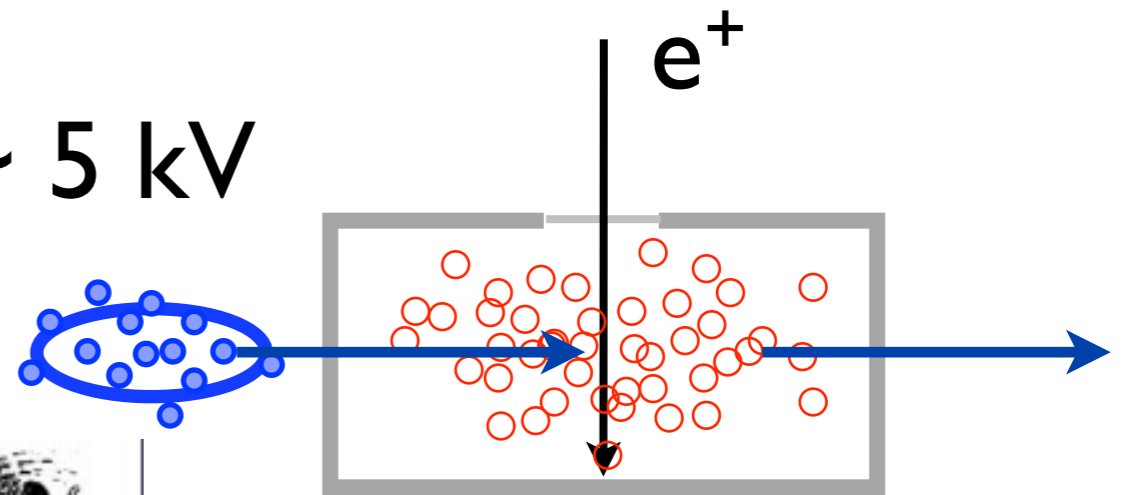


cold \bar{H}^*

but: low rate!

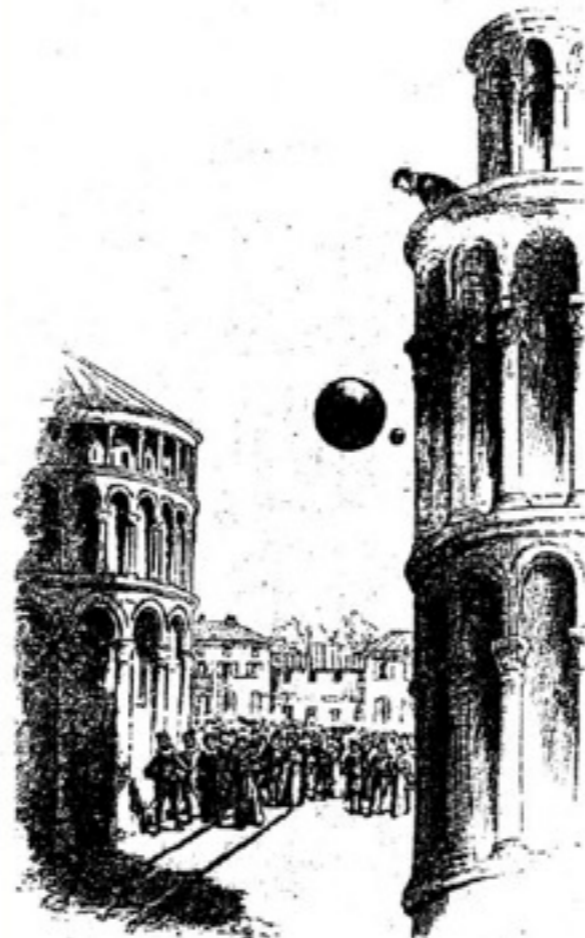
GBAR

$$E_p \sim 5 \text{ kV}$$



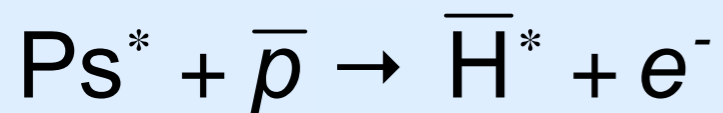
hot \bar{H}^+

but: low rate!

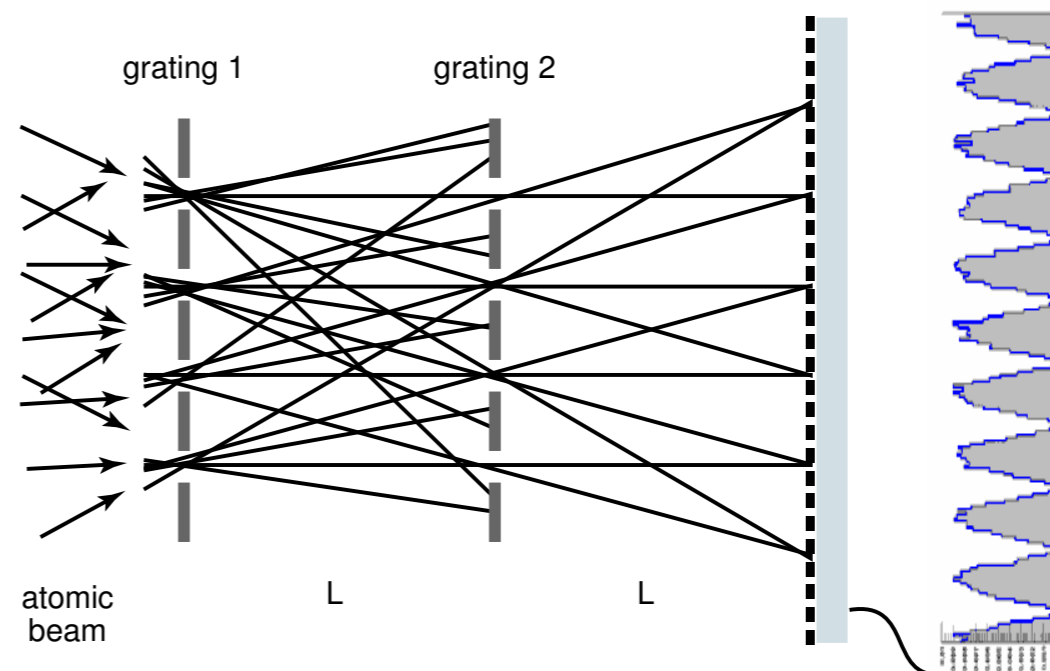
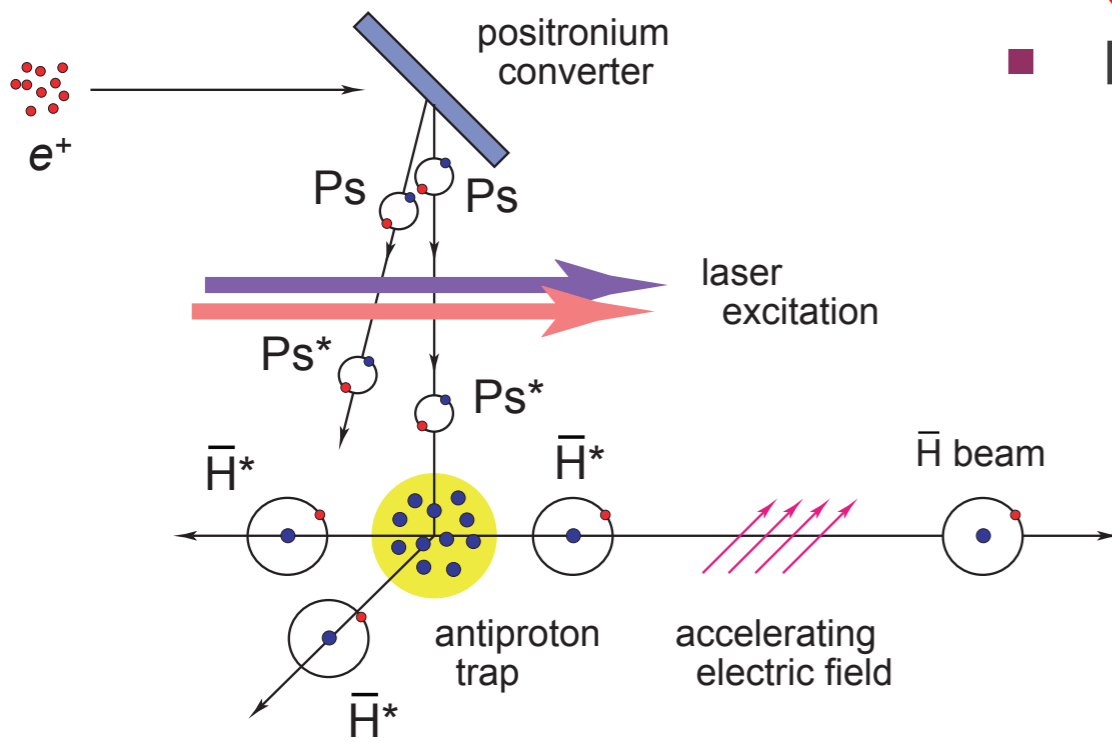


Schematic overview

Physics goals: measurement of the gravitational interaction between matter and antimatter, $\bar{\text{H}}$ spectroscopy, ...

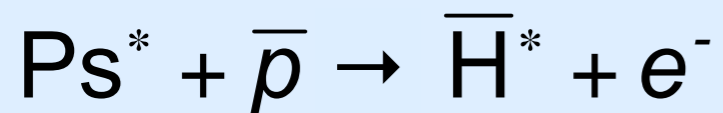


- Anti-hydrogen formation via Charge exchange process with Ps^*
 - o-Ps produced in SiO_2 target close to \bar{p} ; laser-excited to Ps^*
 - $\bar{\text{H}}$ temperature defined by \bar{p} temperature
- Advantages:
 - Pulsed $\bar{\text{H}}$ production (time of flight – Stark acceleration)
 - Narrow and well-defined $\bar{\text{H}}$ n -state distribution
 - Colder production than via mixing process expected
 - Rydberg Ps & $\sigma \approx a_0 n^4 \rightarrow \bar{\text{H}}$ formation enhanced

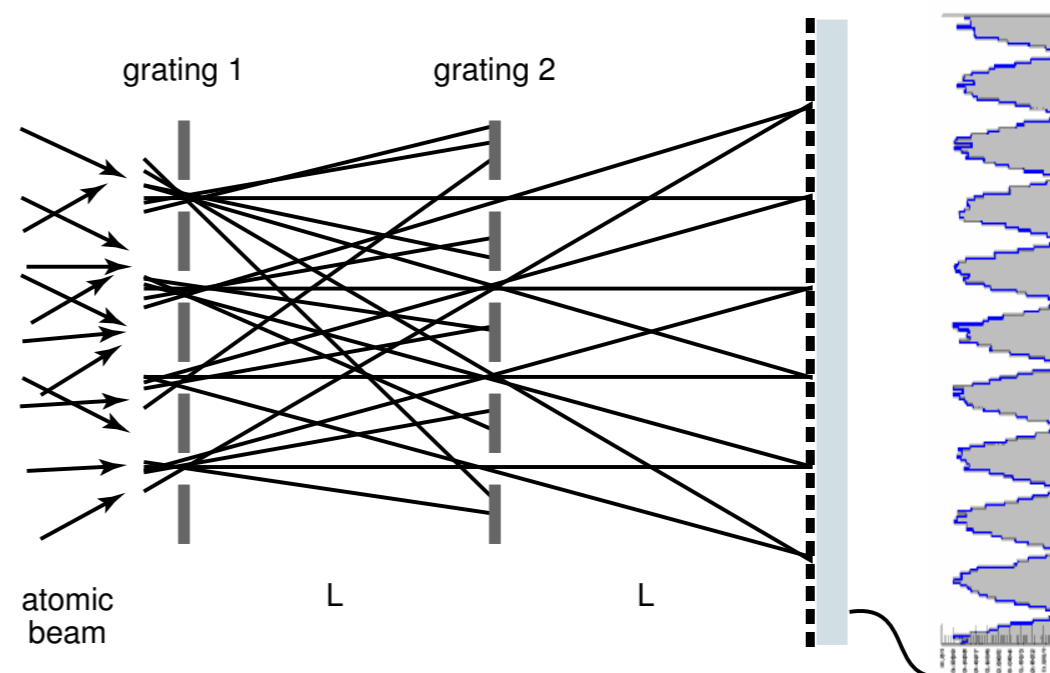
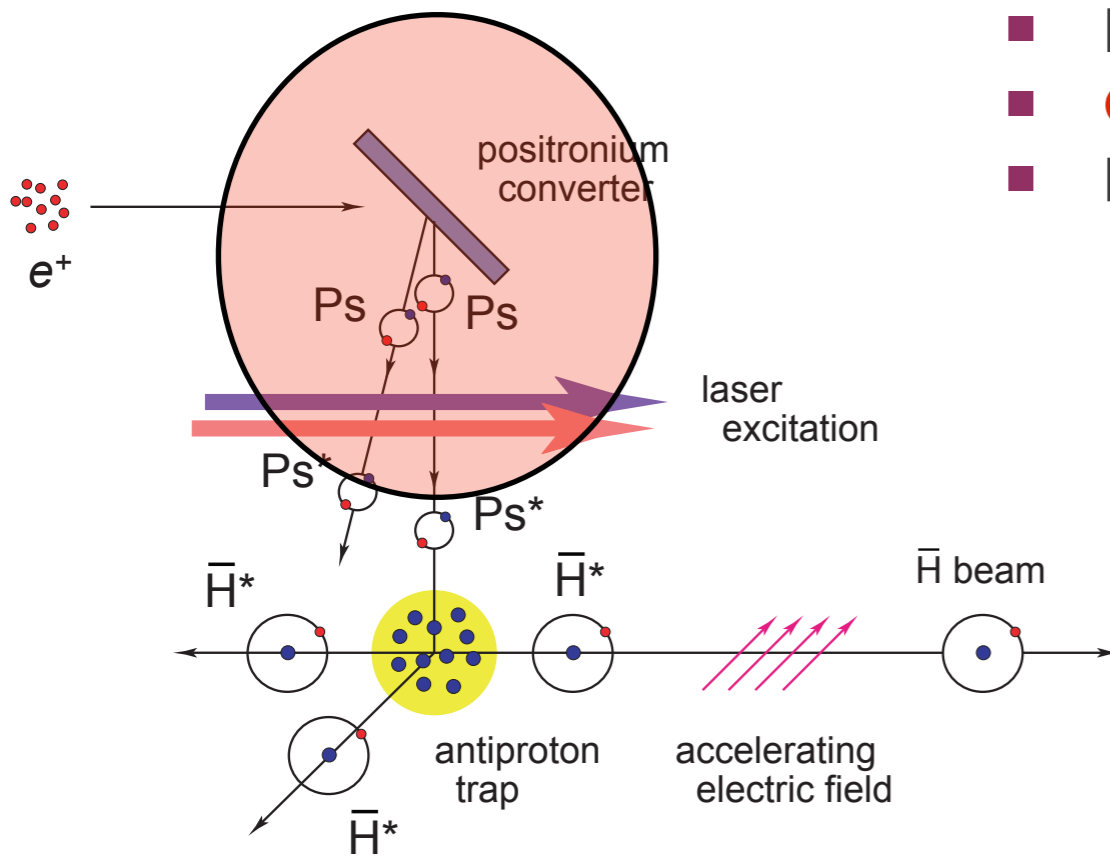


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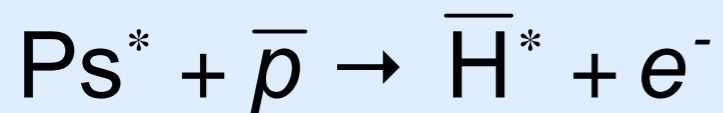


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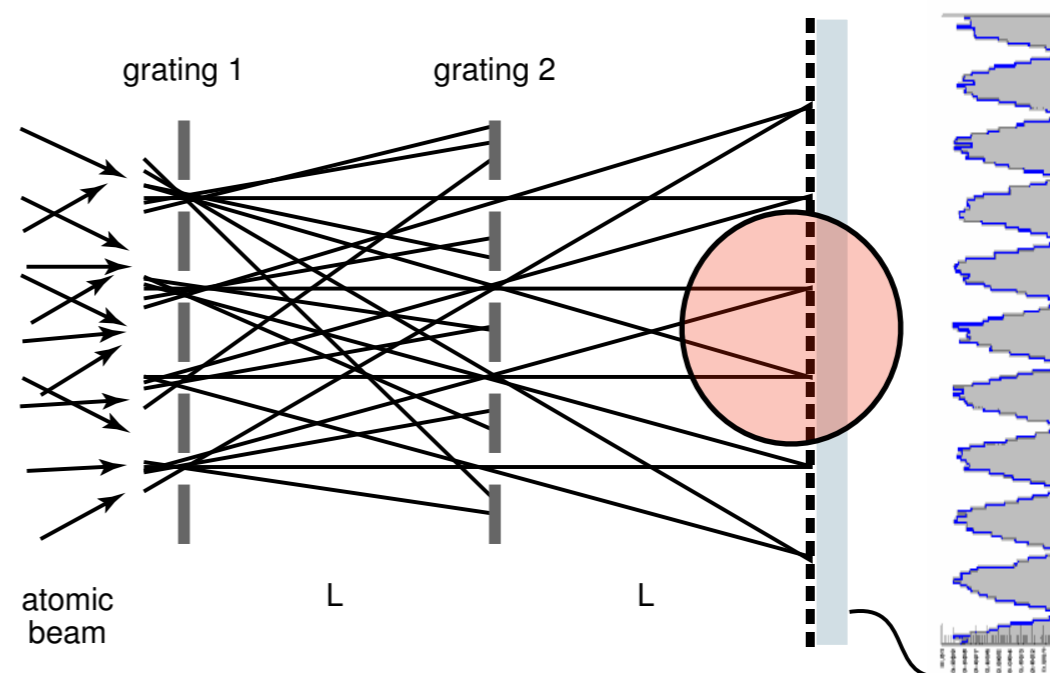
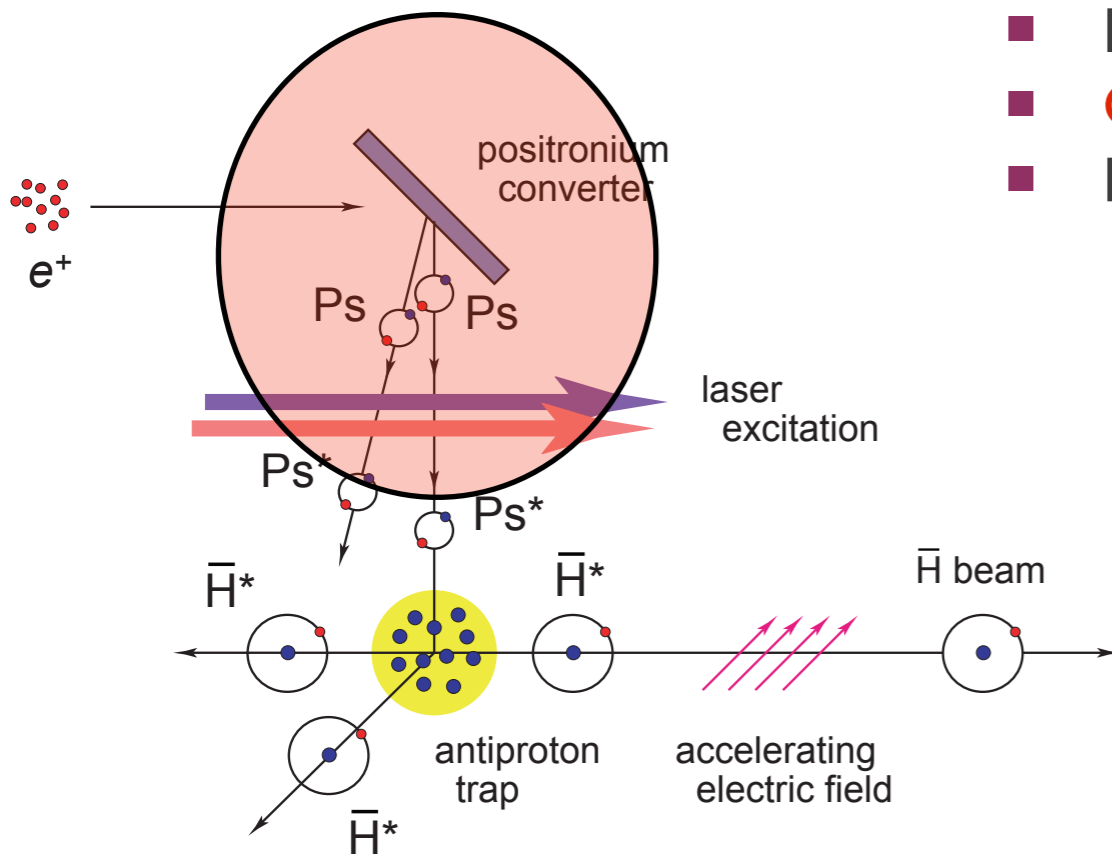


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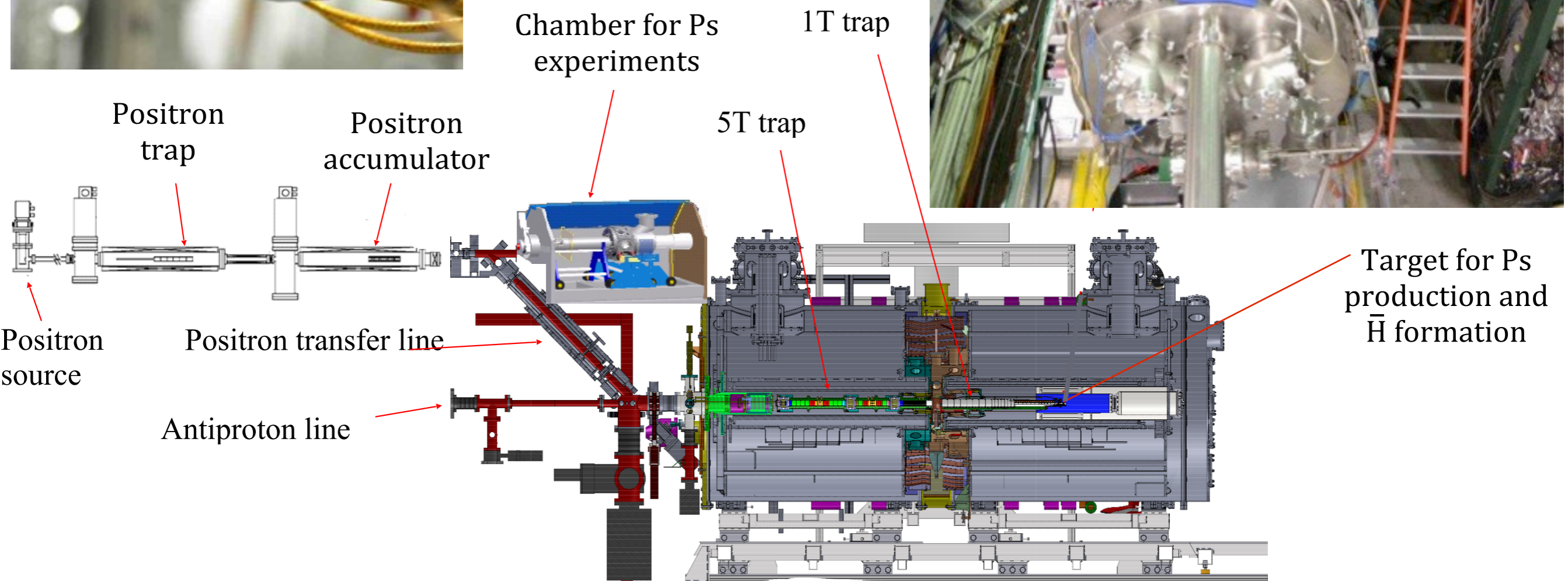
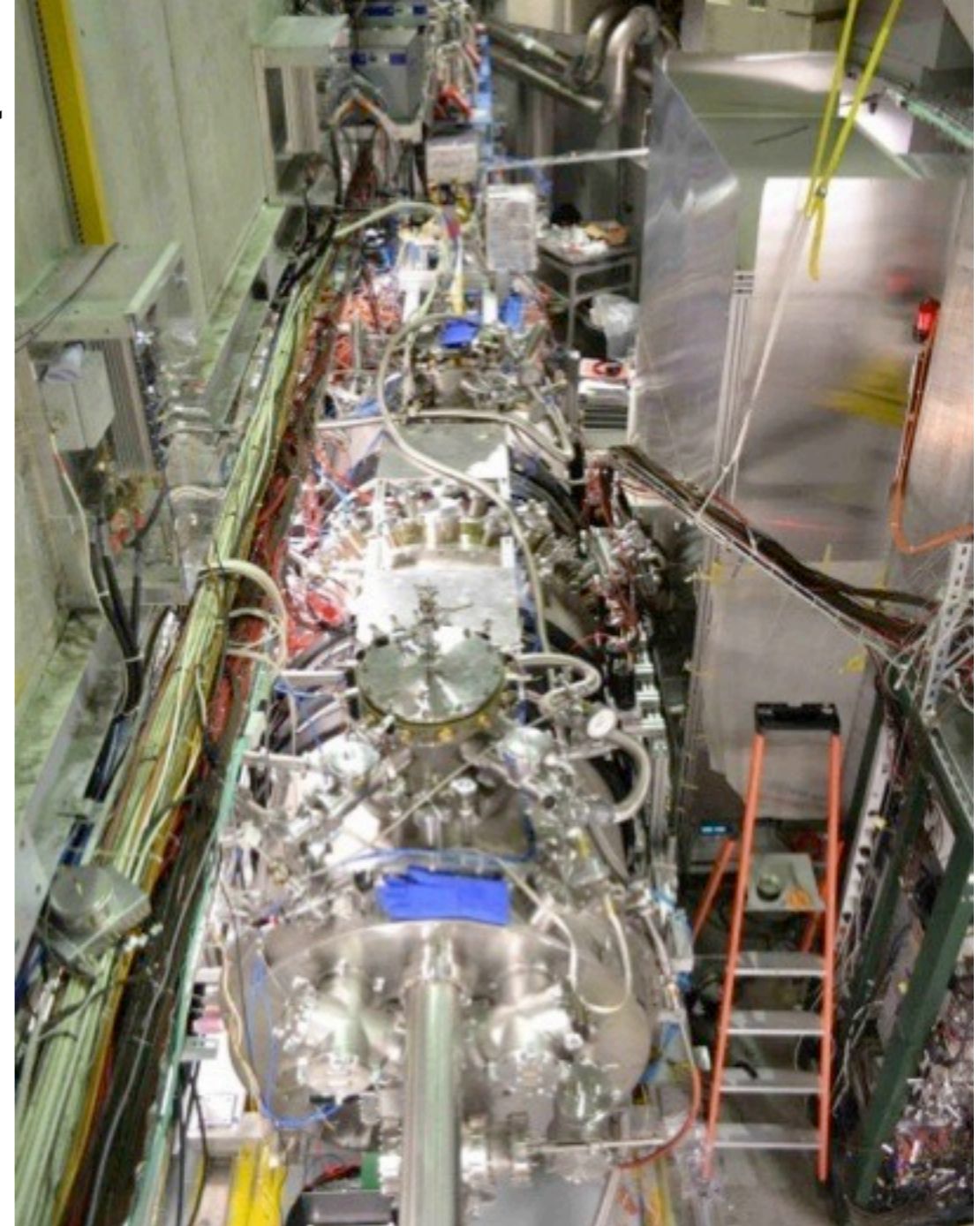
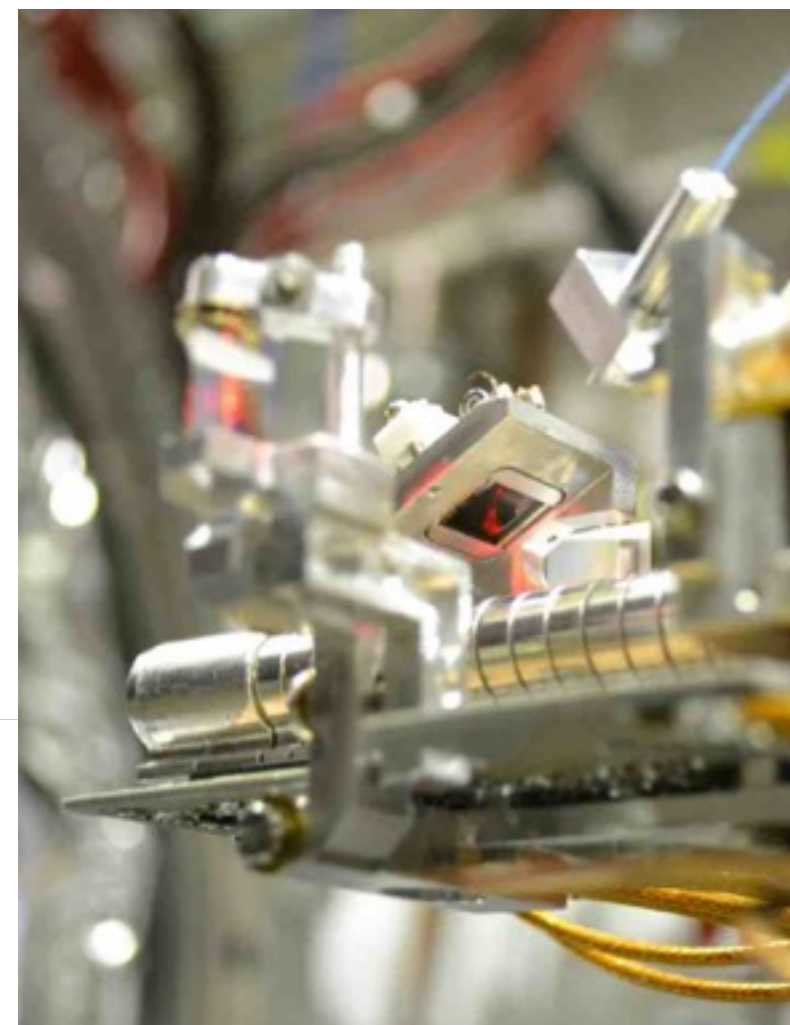
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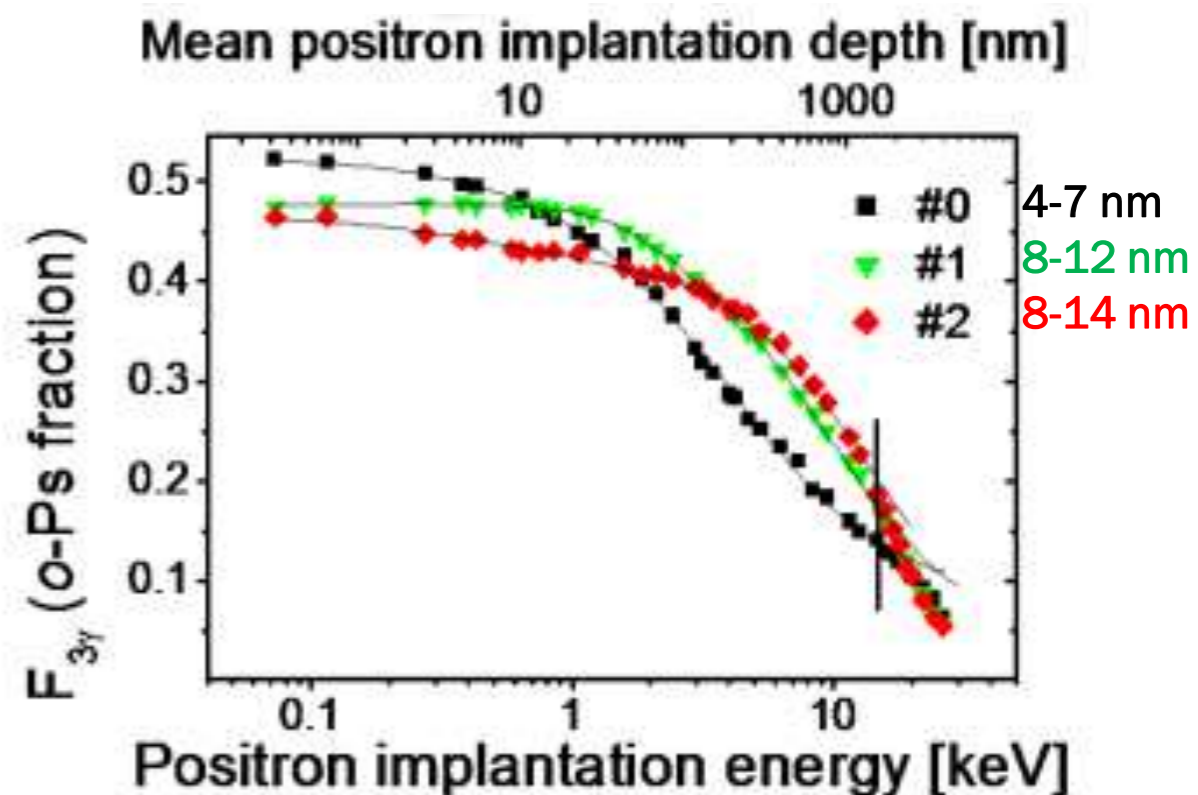
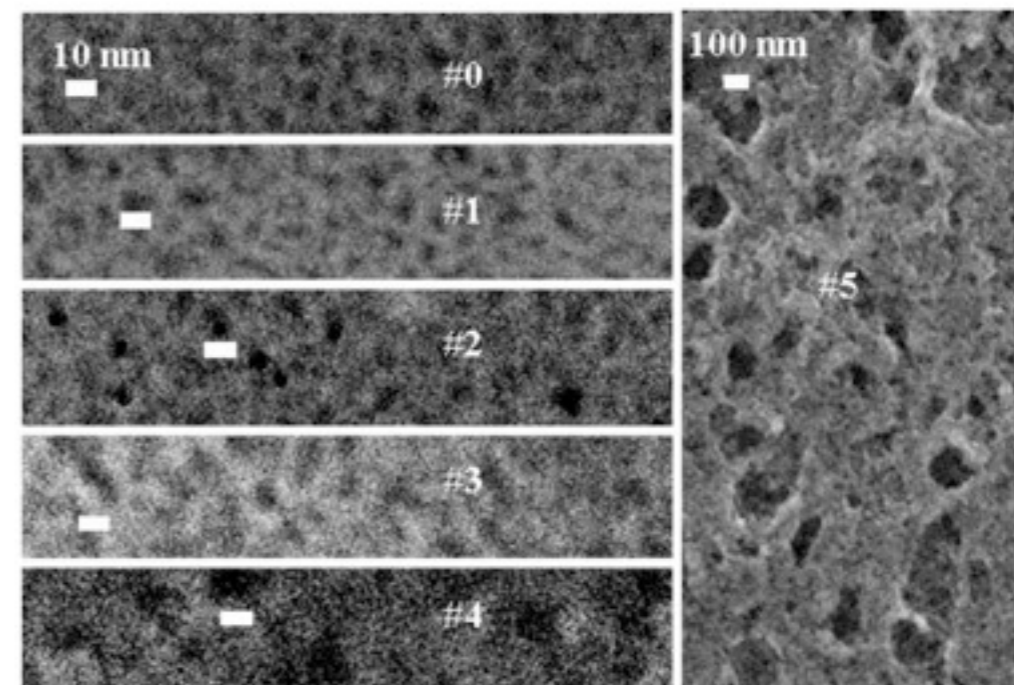
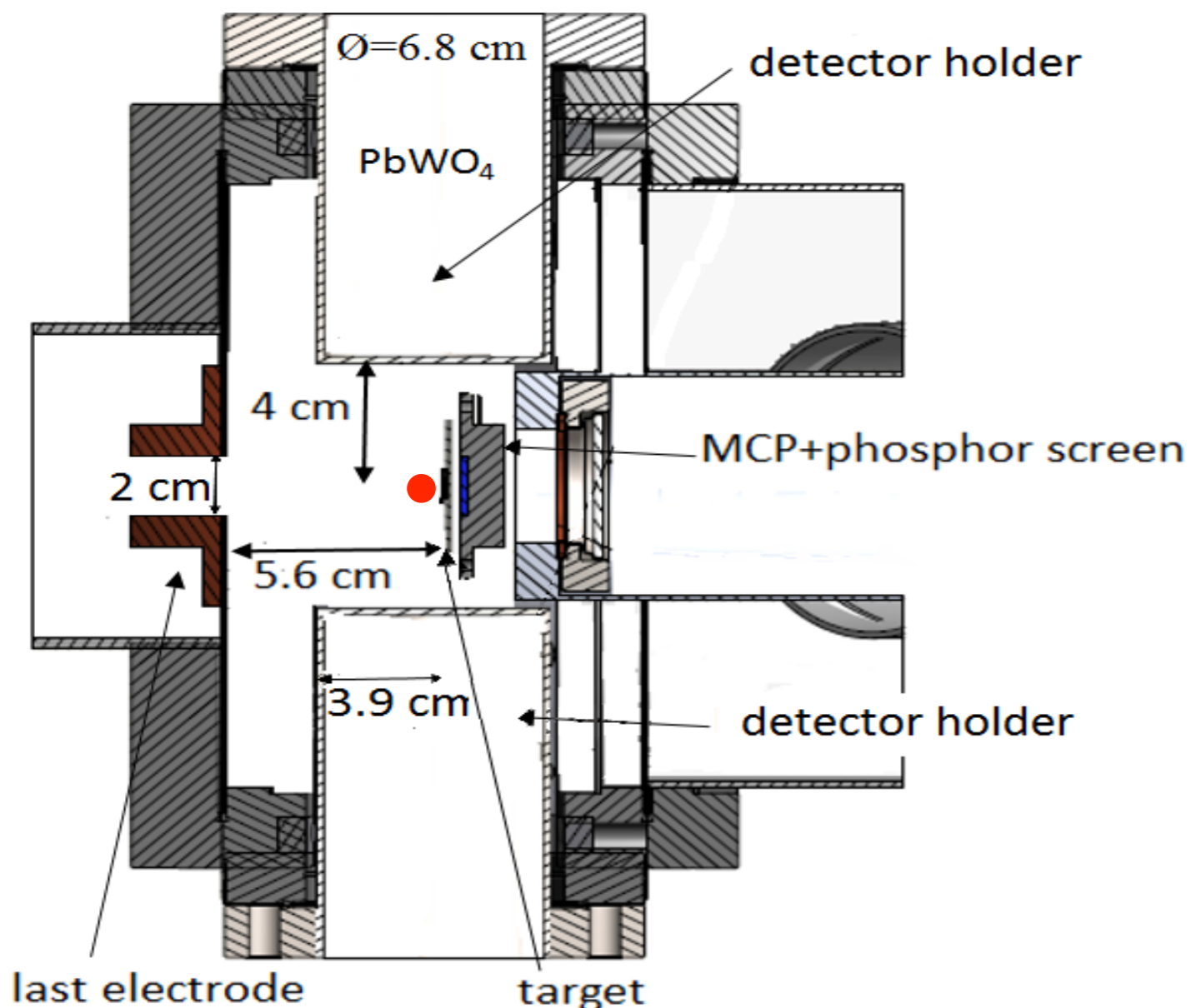


AEgIS experiment



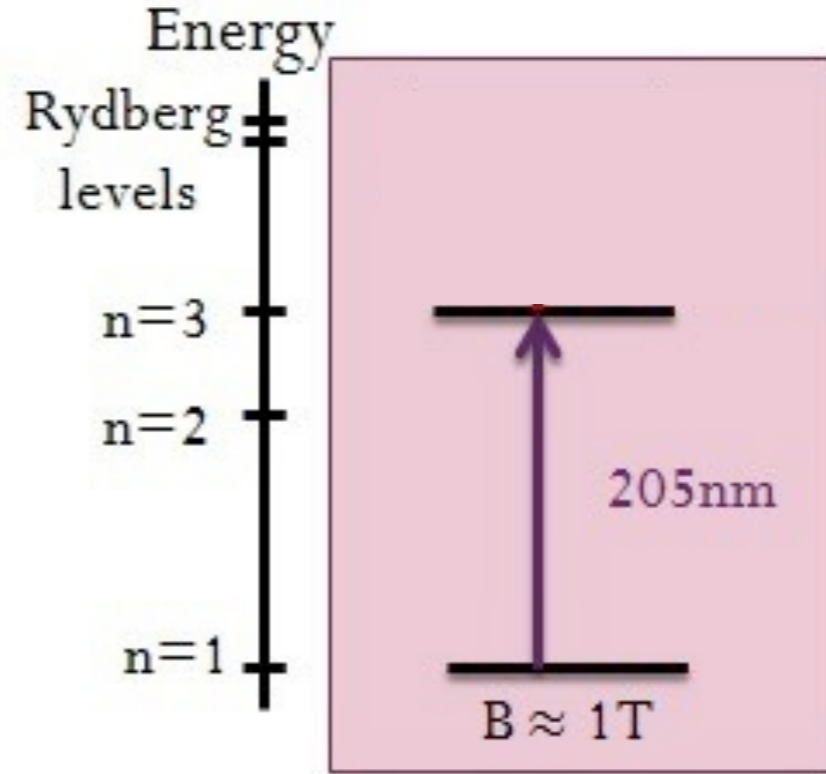
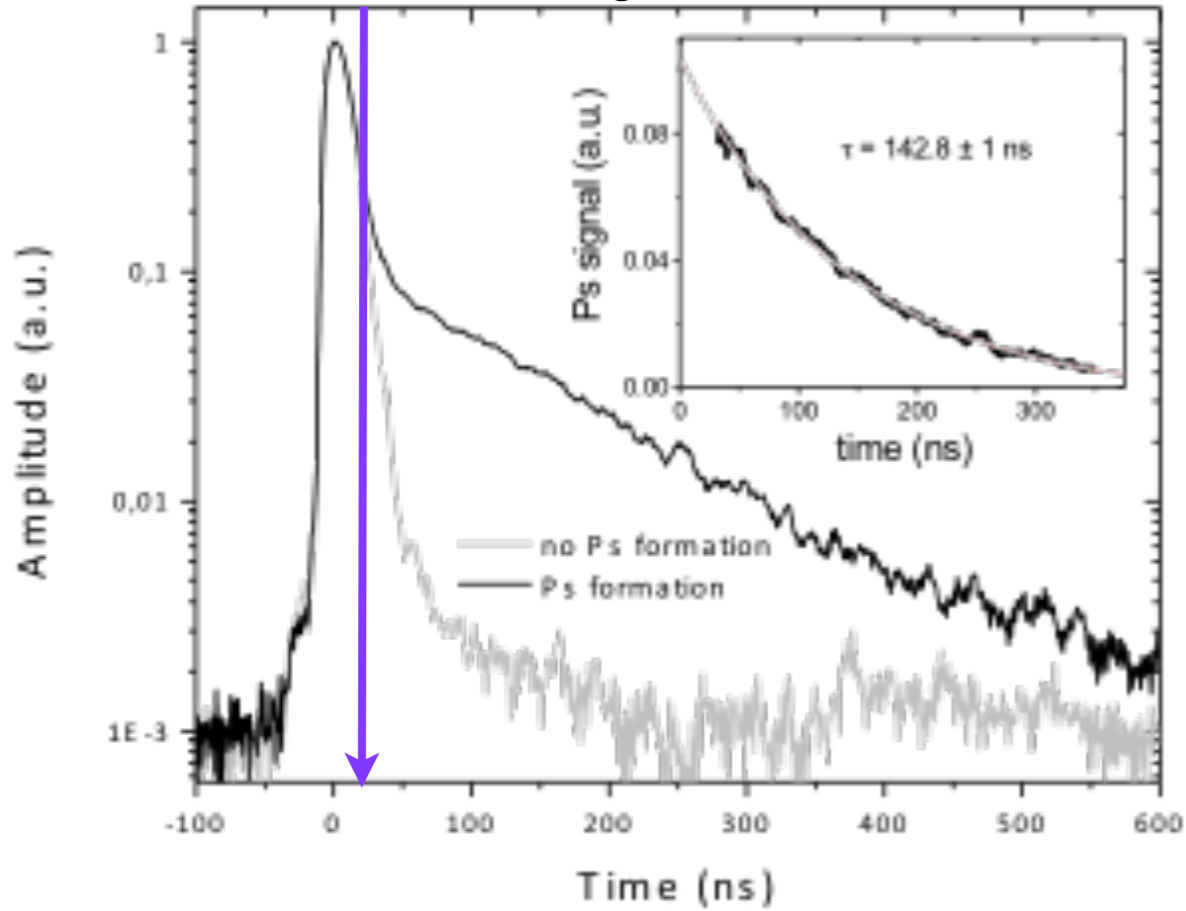
Study of Ps formation, Ps laser excitation, Ps physics: SSPALS

Silica-based nano-porous target (SEM image)

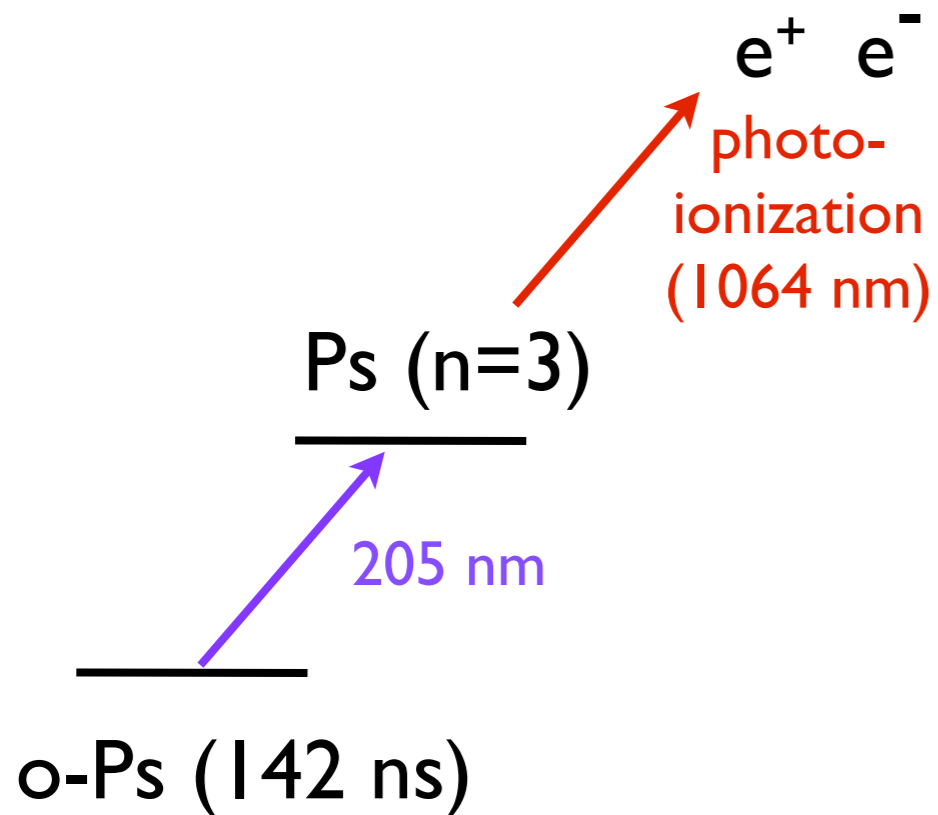


Ps excitation into n=3

average of 15 shots



Energy 60μJ, pulse 2ns, waist 6mm x 8mm, $\sigma \sim 110\text{GHz}$

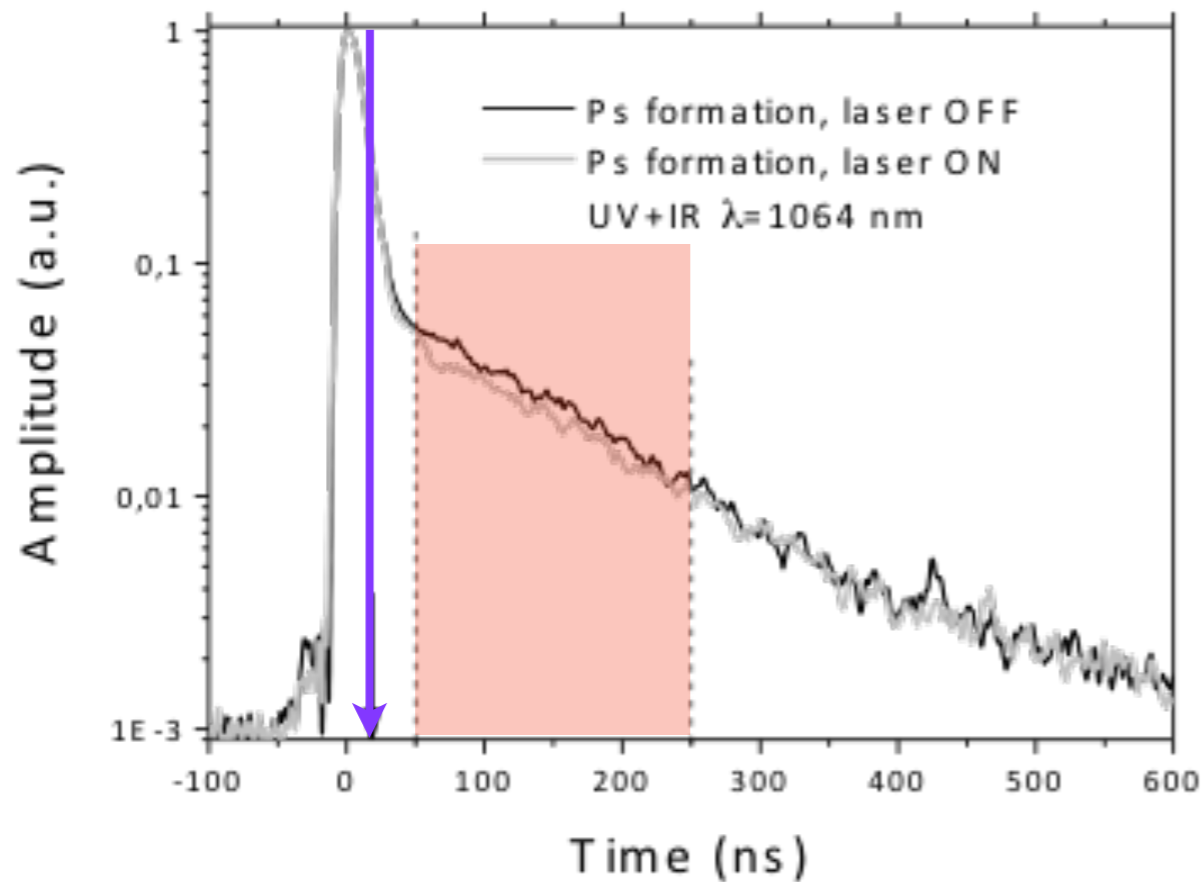


expect *decrease* of o-Ps population on resonance

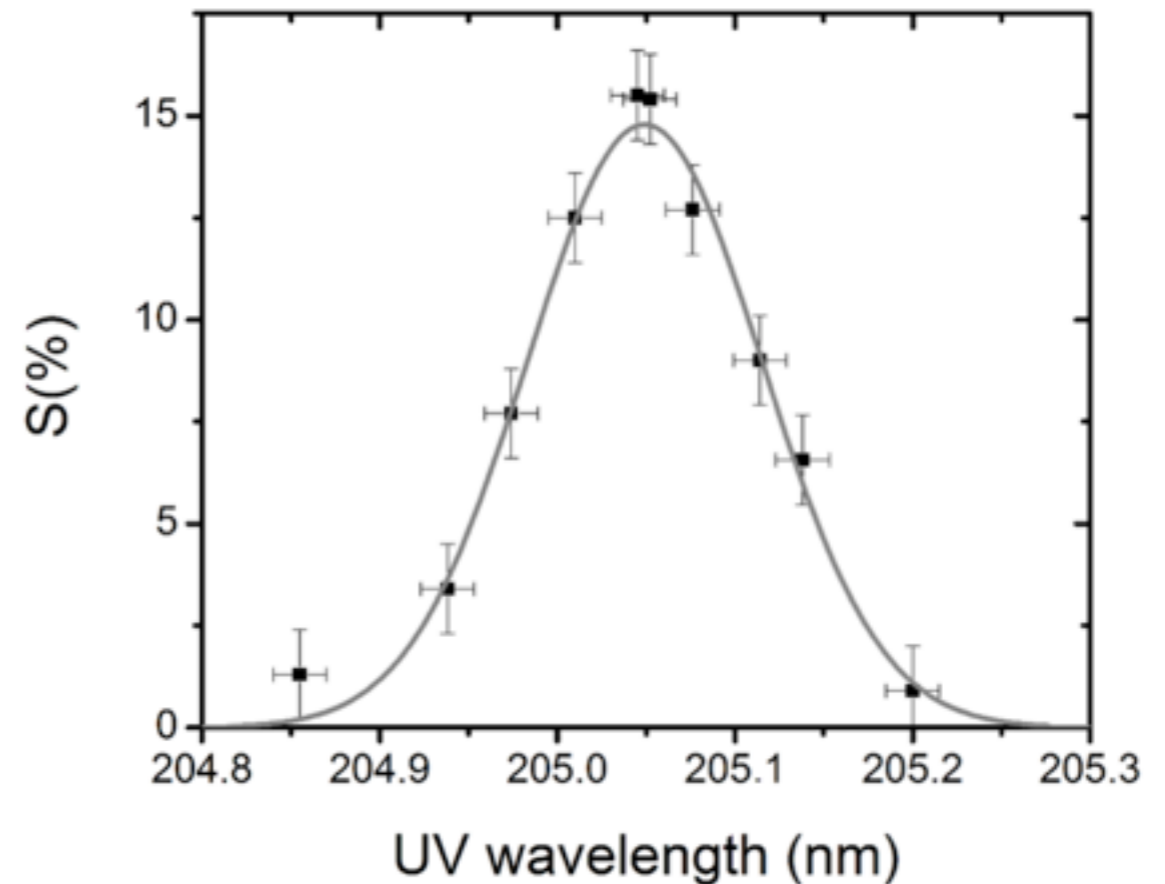
→ *decrease* in (delayed) annihilation rate

Measurement of Ps decay signal, alternating UV on/off, and scanning over UV wavelength

average of 15 shots



3P excitation line centered at 205.05 ± 0.02 nm



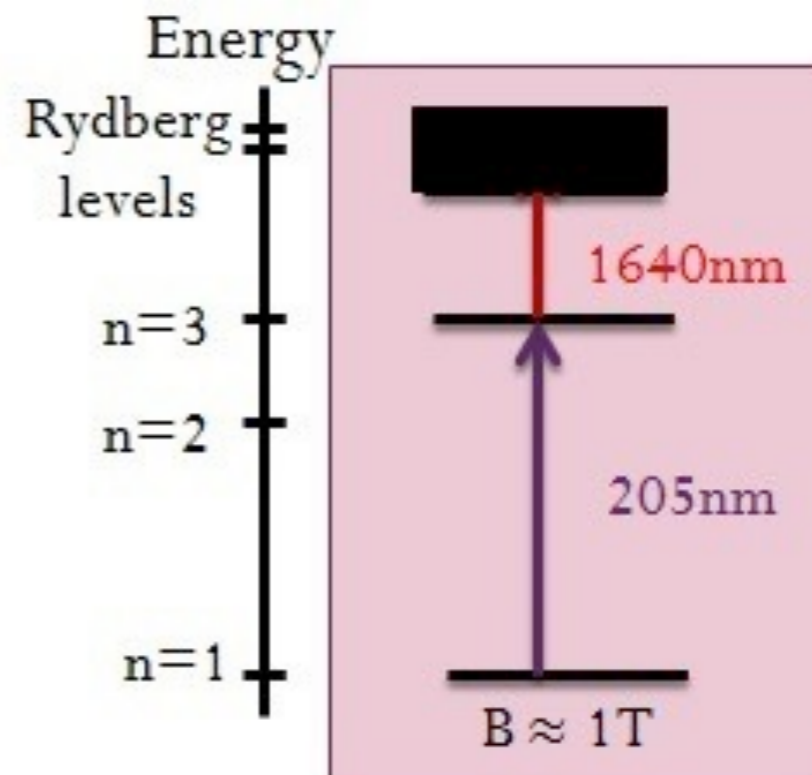
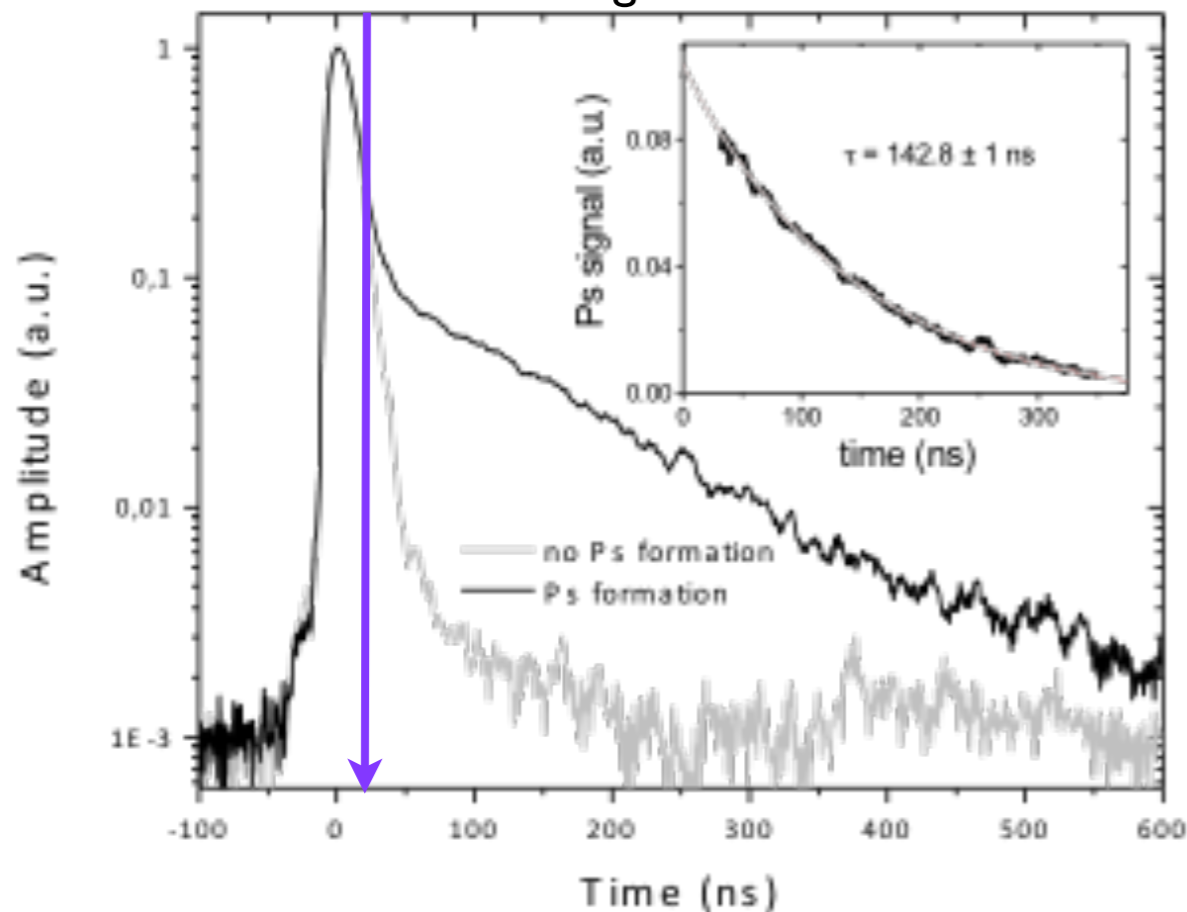
$$S(\%) = \frac{(\text{Area laser OFF} - \text{Area laser ON})}{\text{Area laser OFF}}$$

Excitation + photoionization efficiency $\sim 15\%$ (limited by laser linewidth)

From this measurement, extract an average temperature of the excited o-Ps : $T \sim 1300$ K (Doppler broadening)

Ps excitation from n=3 into n~15

average of 15 shots



Ps* (n=15, $\tau \sim \mu\text{s}$)

$\nearrow \sim 1700 \text{ nm}$

Ps (n=3)

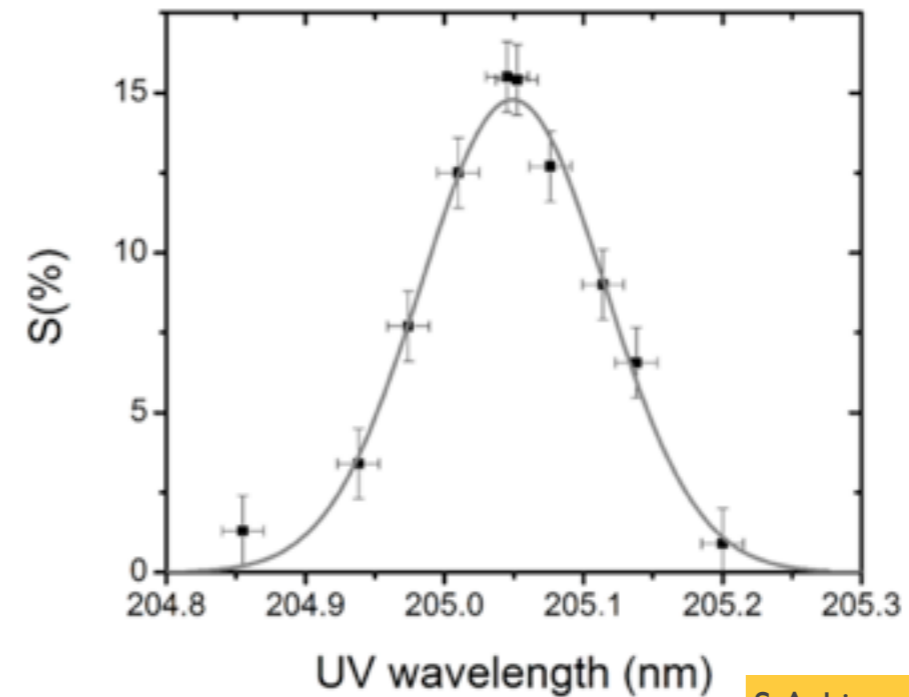
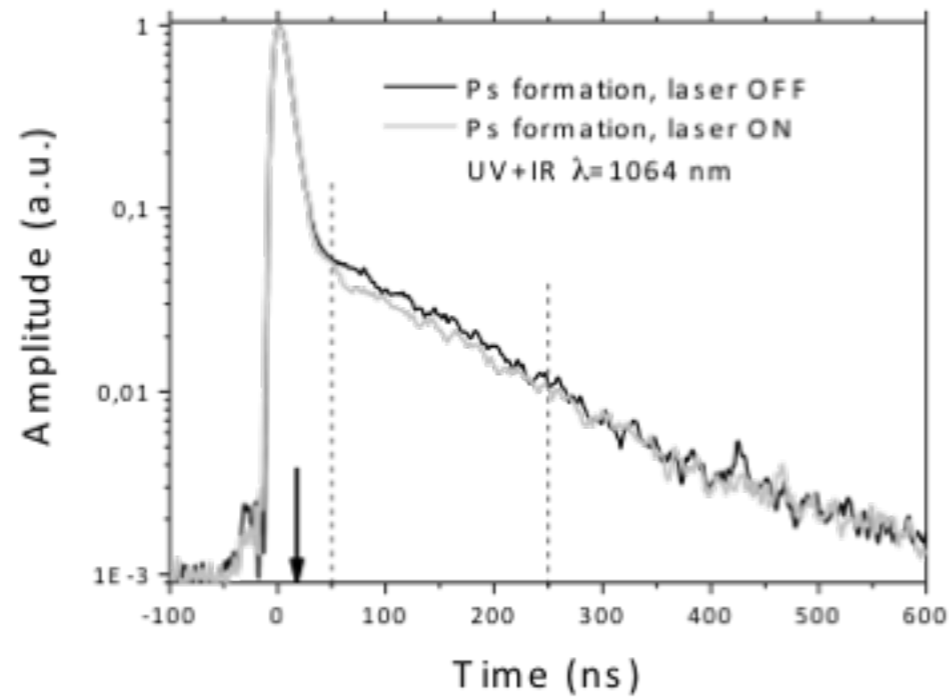
$\nearrow 205 \text{ nm}$

o-Ps (142 ns)

expect *decrease* of o-Ps population on resonance and *appearance* of long-lived Ps*

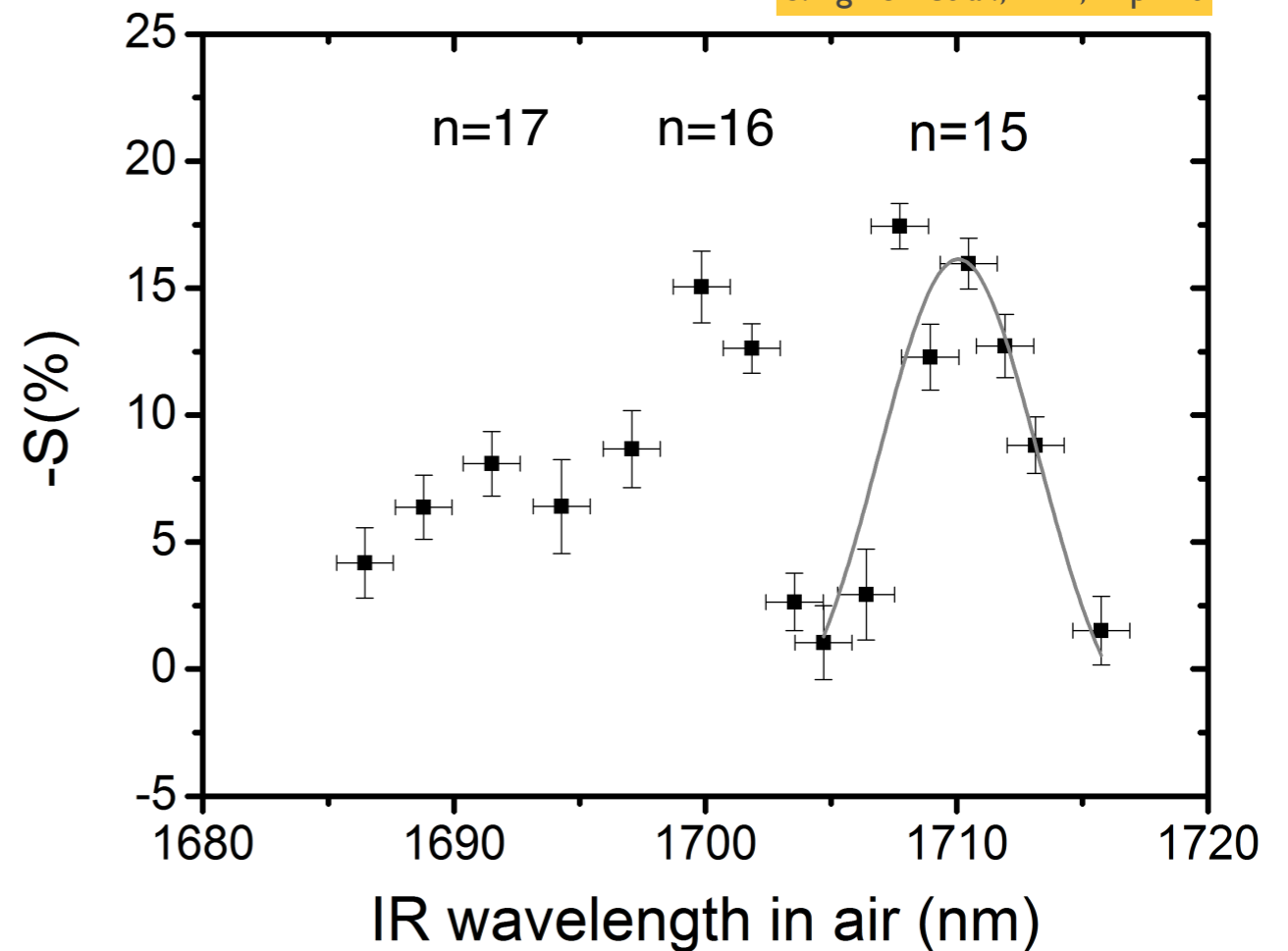
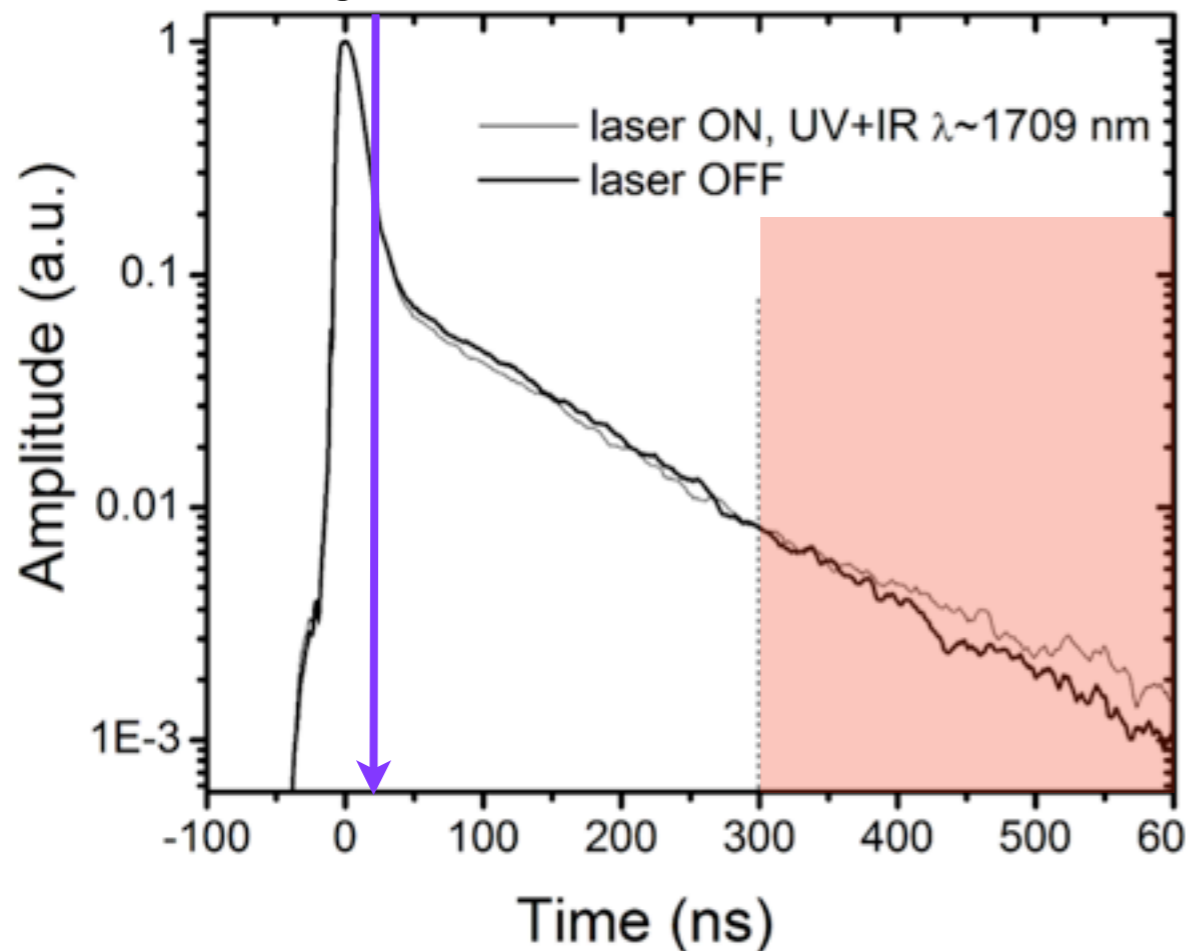
\rightarrow *increase* in (very delayed) annihilation rate

Measurement of Ps decay signal, alternating UV+IR on/off, and scanning over IR wavelength

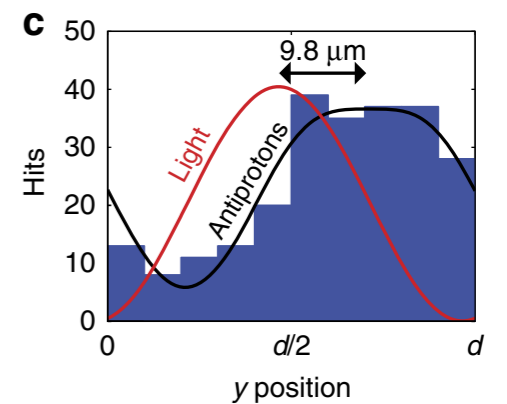
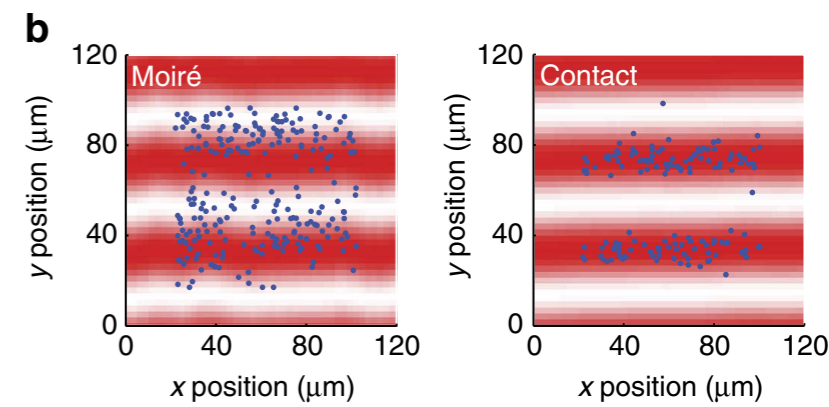
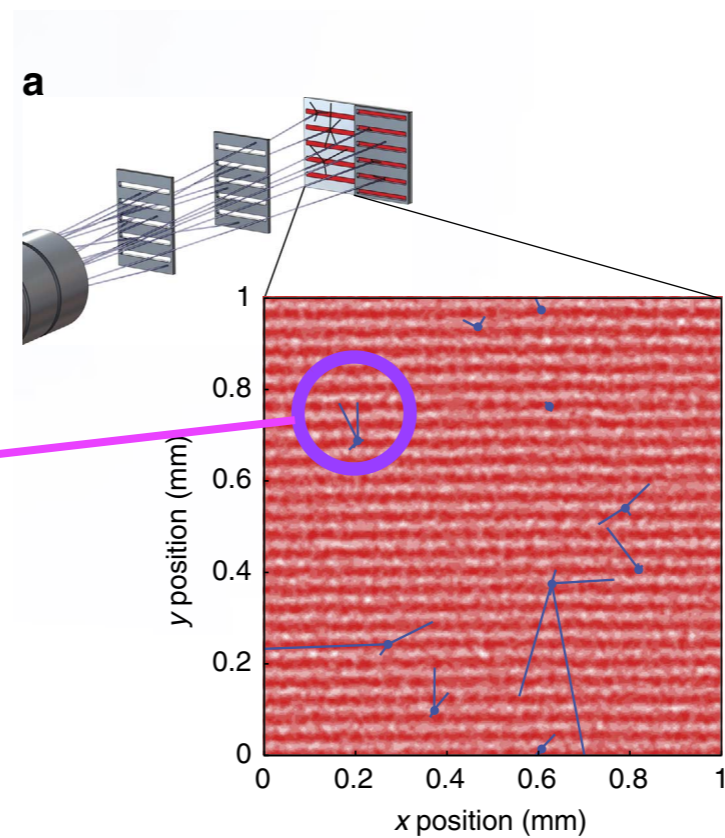
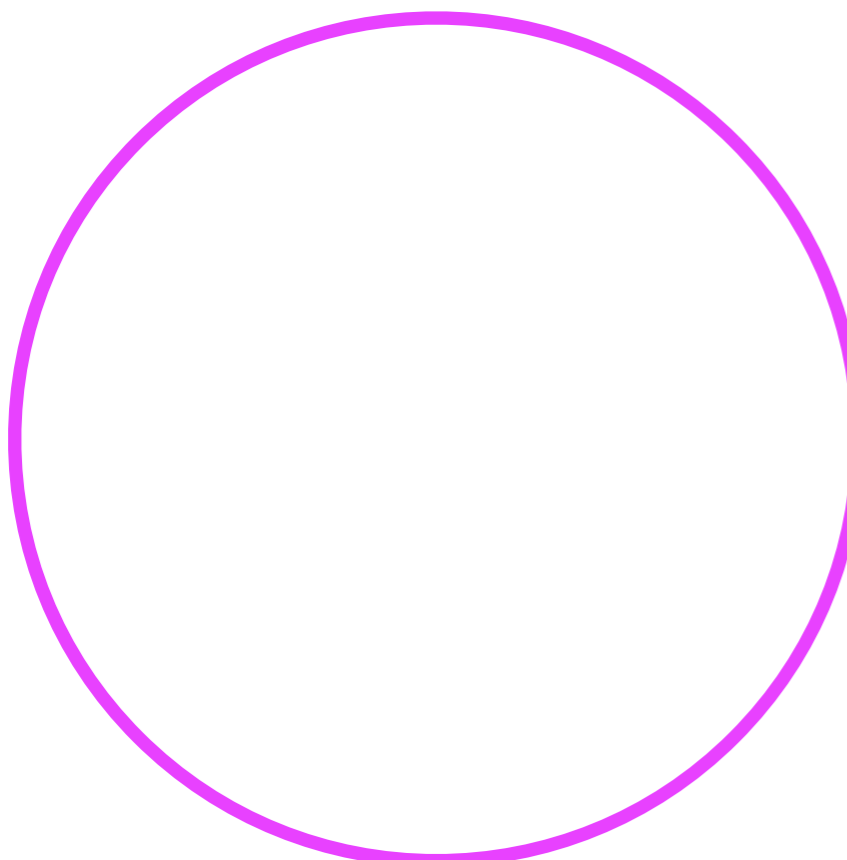
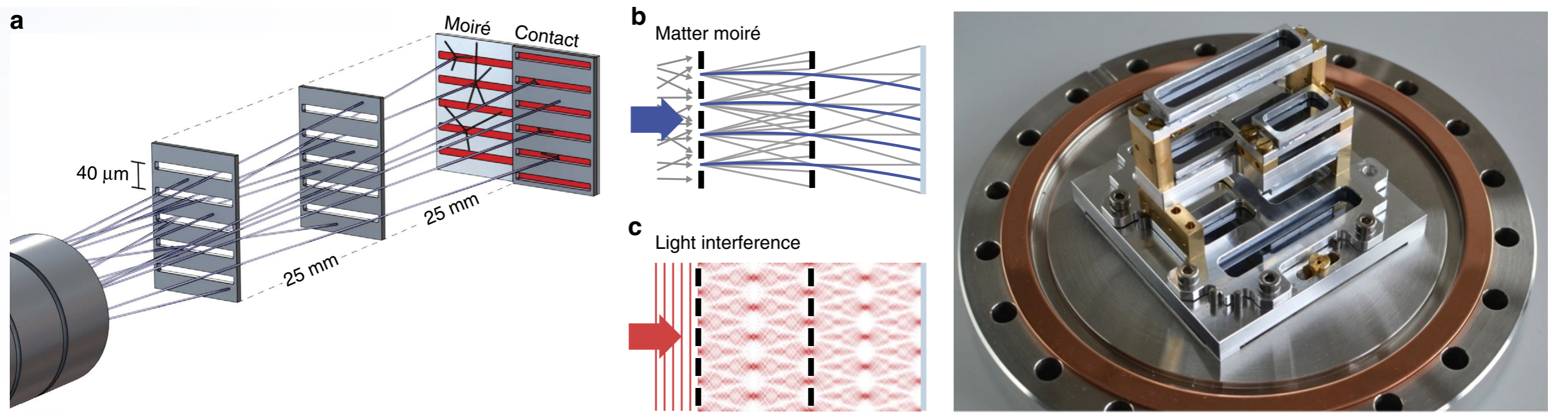


S.Aghion et al., PRA, in print

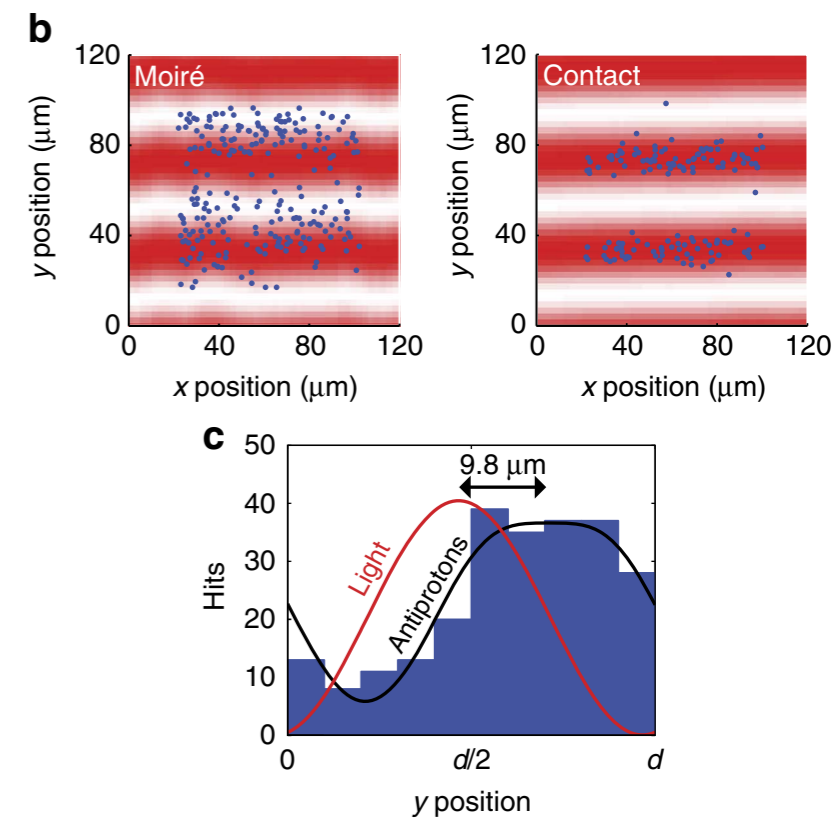
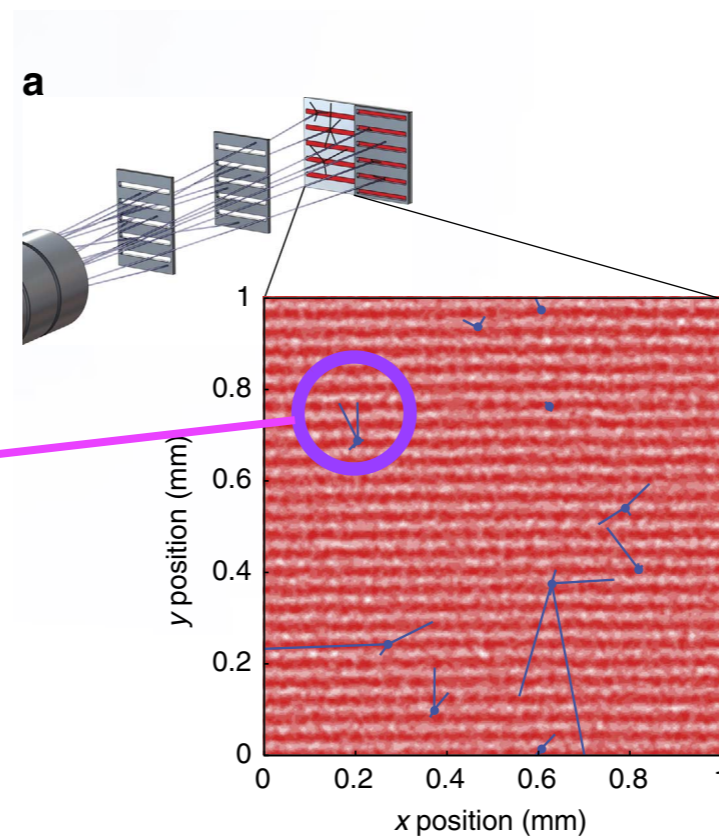
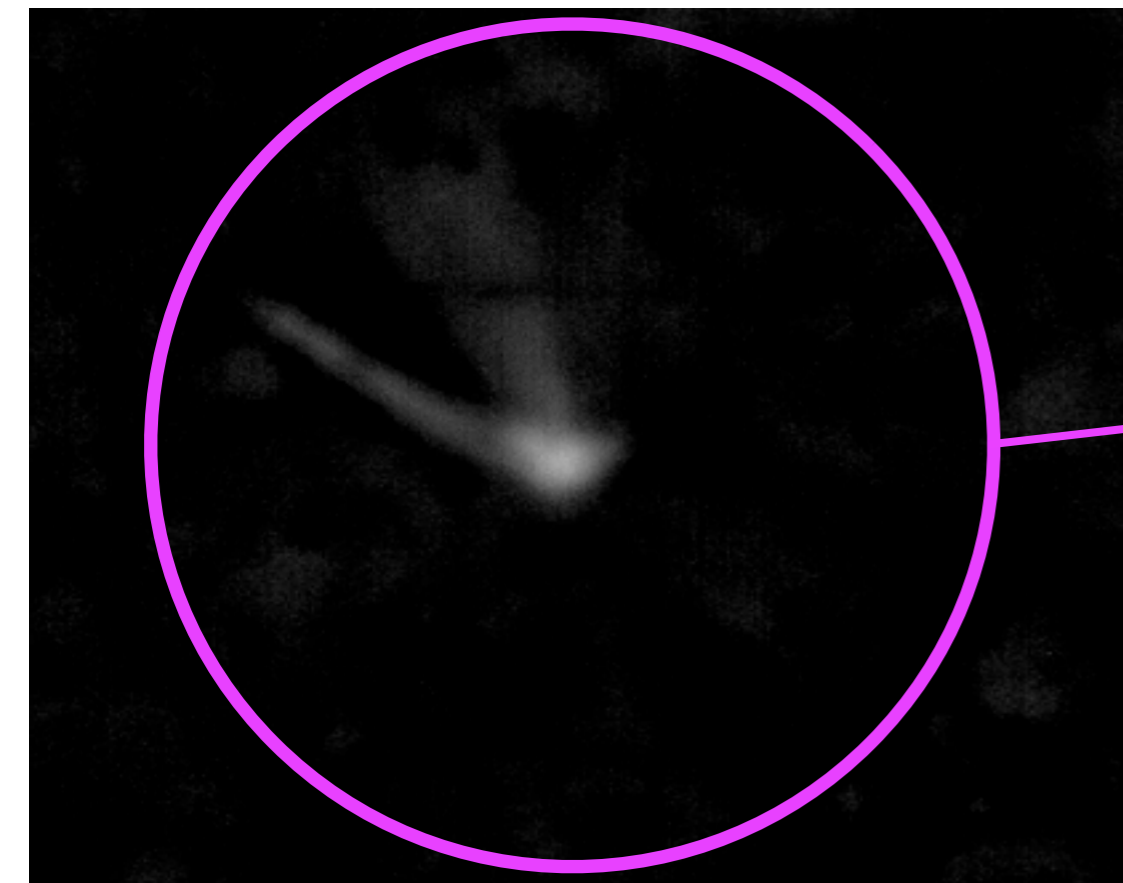
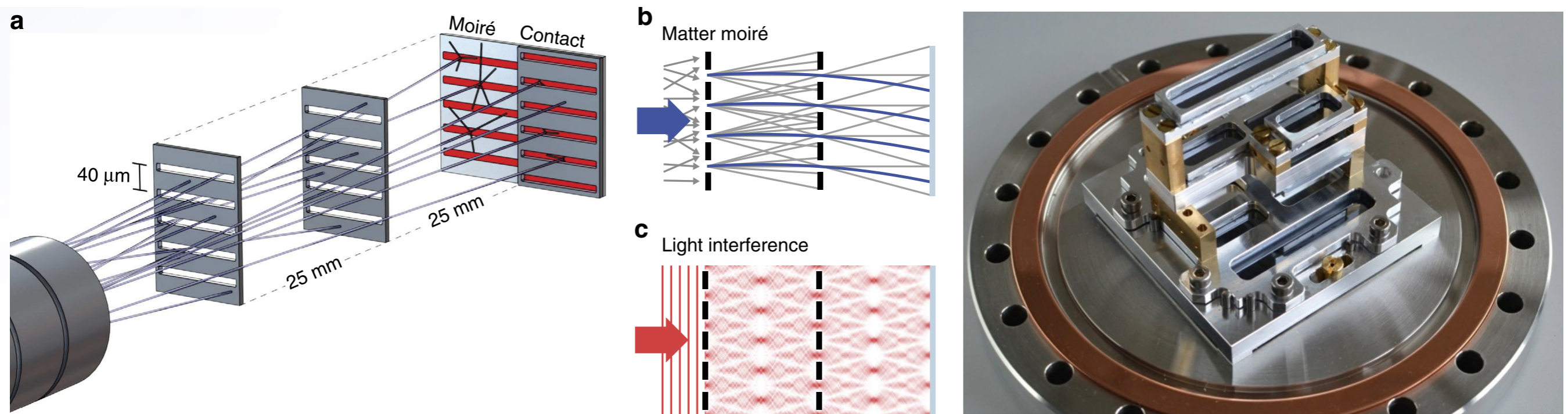
average of 40 shots



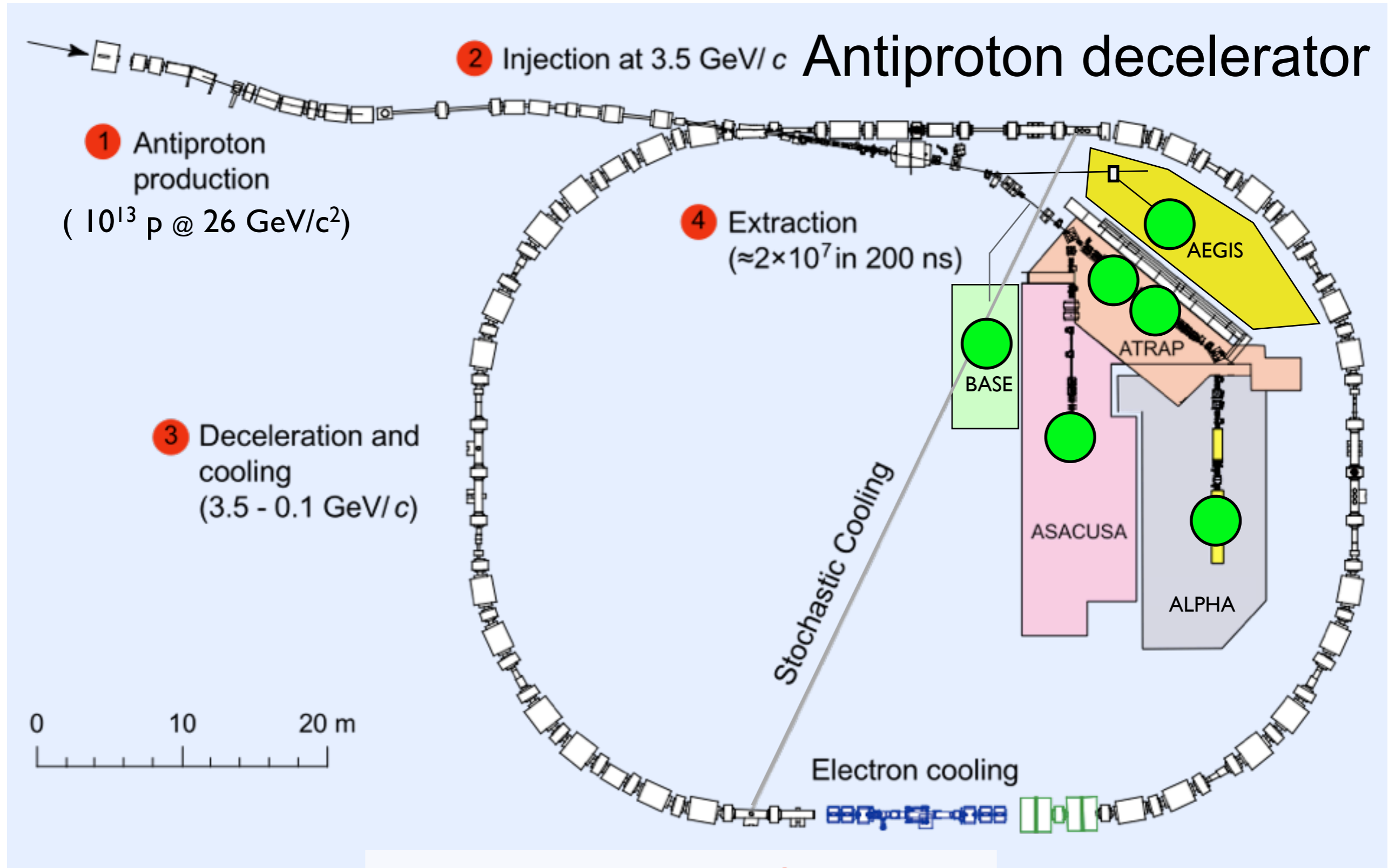
Deflectometer test with antiprotons



Deflectometer test with antiprotons



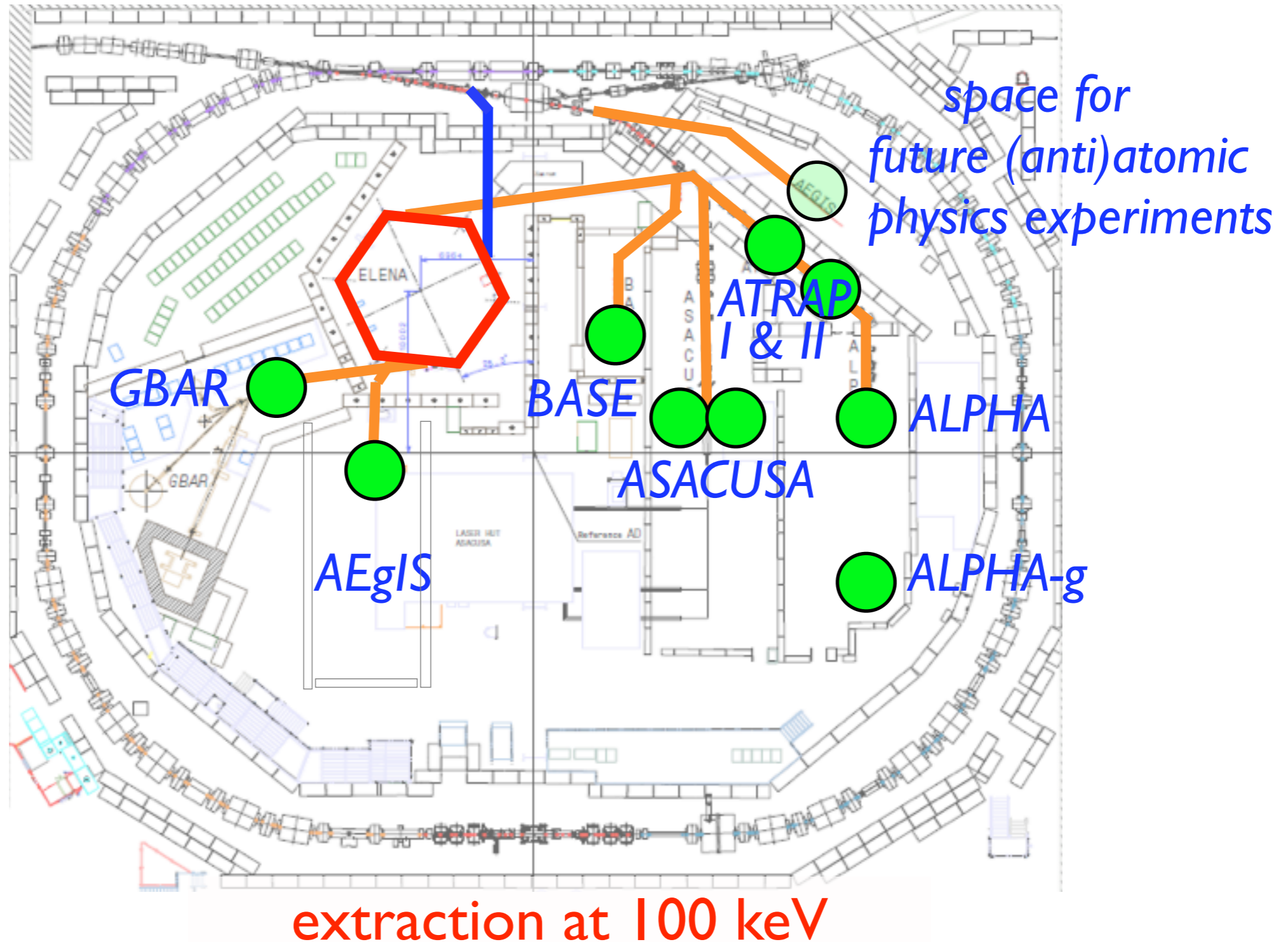
Two main challenges: more / colder antiprotons



extraction at 5.3 MeV

Two main challenges: more / colder antiprotons
current methods for trapping them are quite inefficient

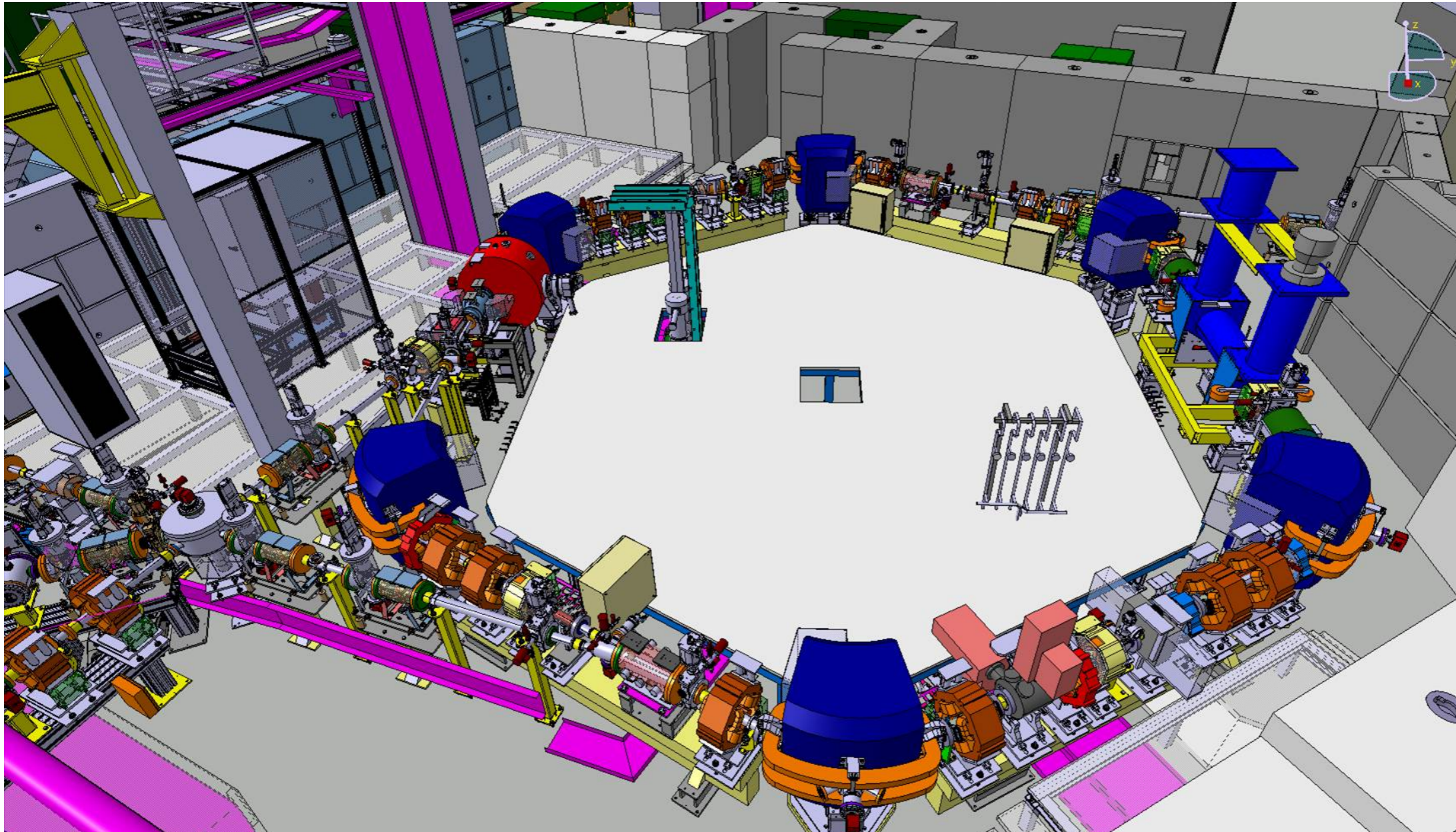
ELENA to the rescue



ELENA is a tiny new decelerator that:

- dramatically slows down the antiprotons from the AD
- increases the *antiproton* trapping efficiency x 100
- allows 4 experiments to run in parallel
- allows new experiments to be considered

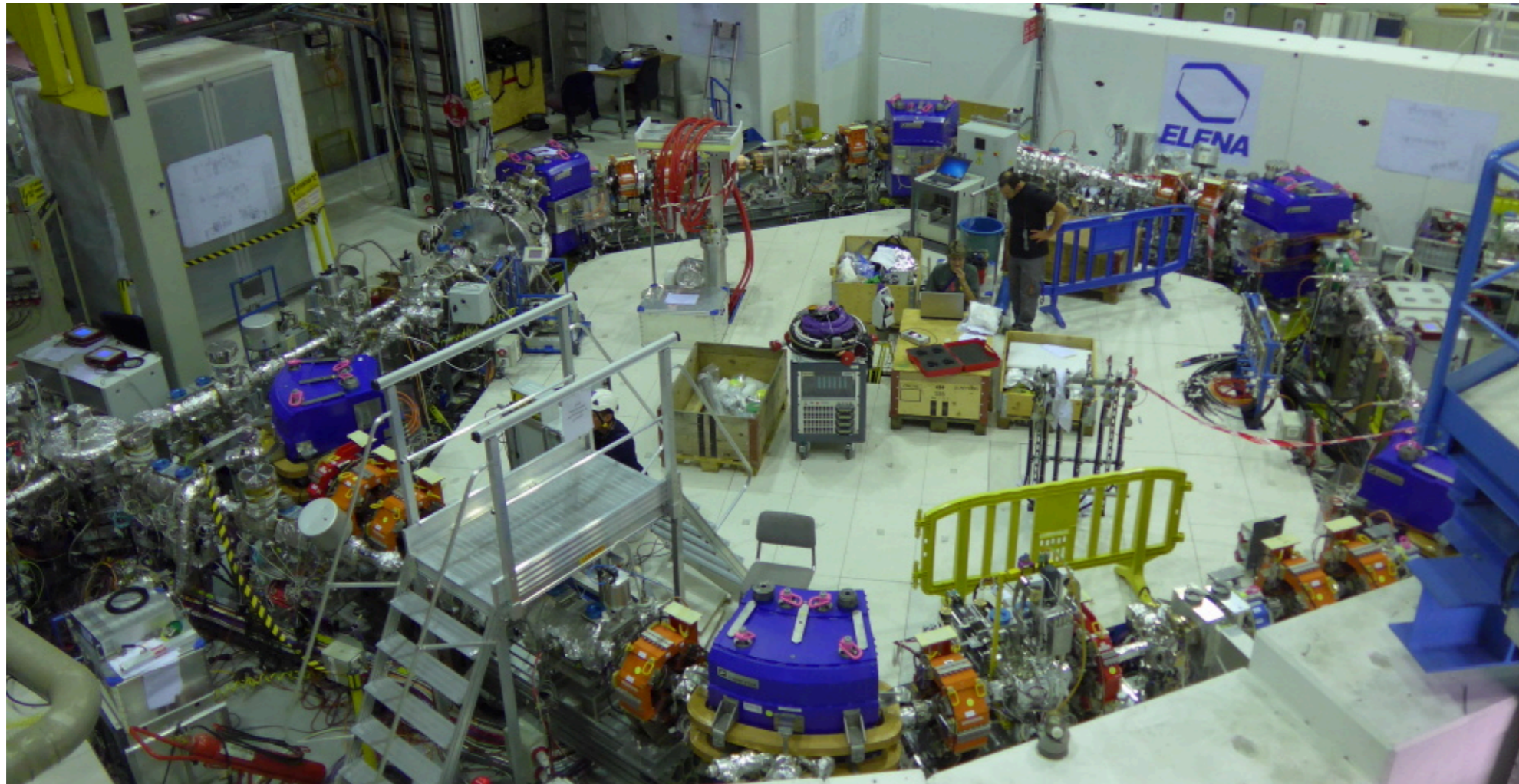
commissioning
ongoing in 2017



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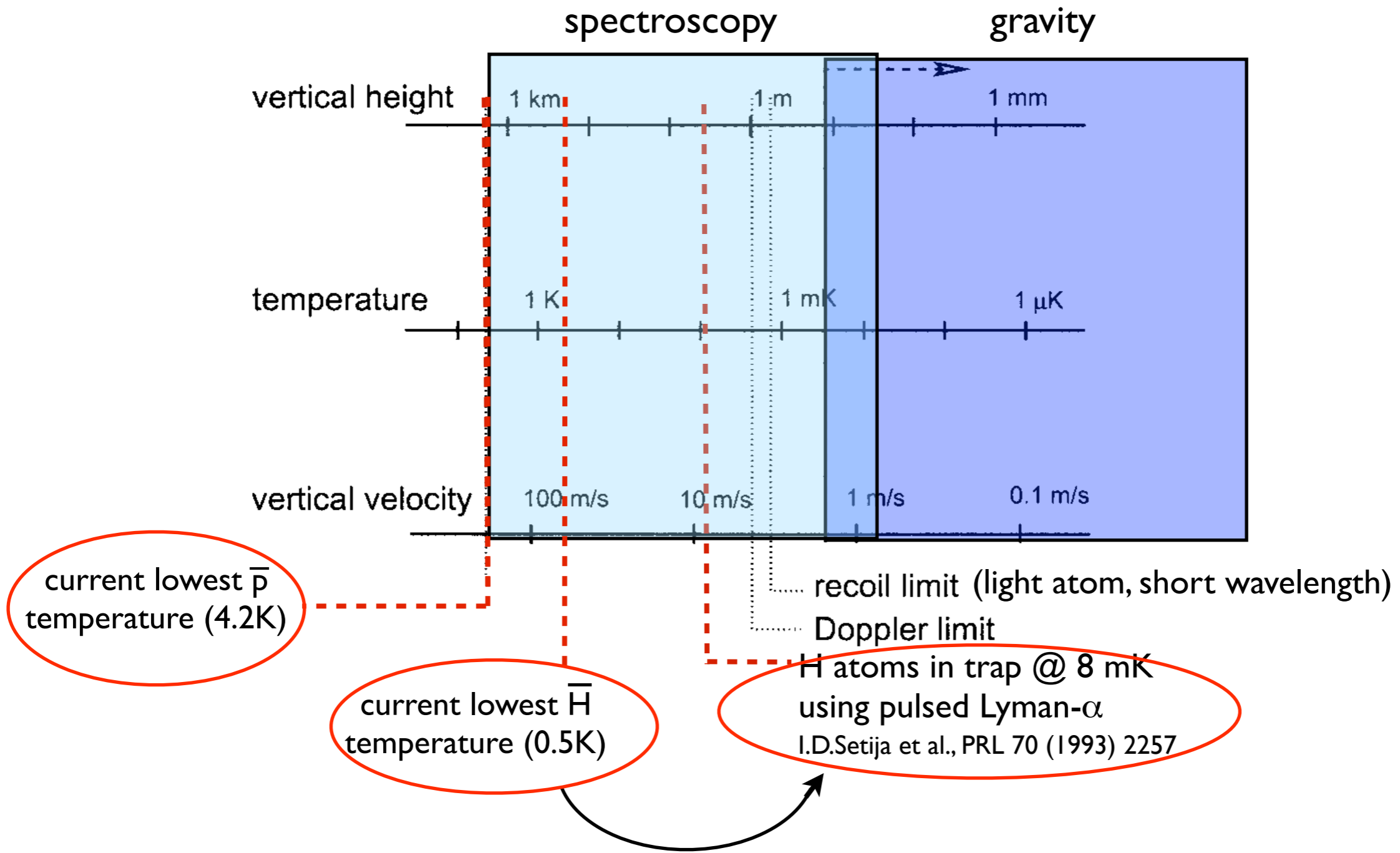
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ongoing in 2017



Two main challenges: more / colder antiprotons

“Ultra-cold” ($\sim 1 \mu\text{K}$) Antihydrogen



IS \rightarrow 2P laser cooling: cw Lyman- α source
Eikema, Walz, Hänsch, PRL 86 (2001) 5679

very long-term goals: gravity, spectroscopy in sub-mK traps
sympathetic cooling to the rescue

GBAR experiment

cooling of \bar{H}^+

J. Walz and T. Hänsch, Gen. Rel. and Grav. 36 (2004) 561

formation of \bar{H}^+ (binding energy = 0.754 eV)

how? perhaps through $Ps(2p) + \bar{H}(1s) \rightarrow \bar{H}^+ + e^-$

Roy & Sinha, EPJD 47 (2008) 327

sympathetic cooling of \bar{H}^+

e.g. In^+ \rightarrow 20 μ K

photodetachment at $\sim 6083 \text{ cm}^{-1}$

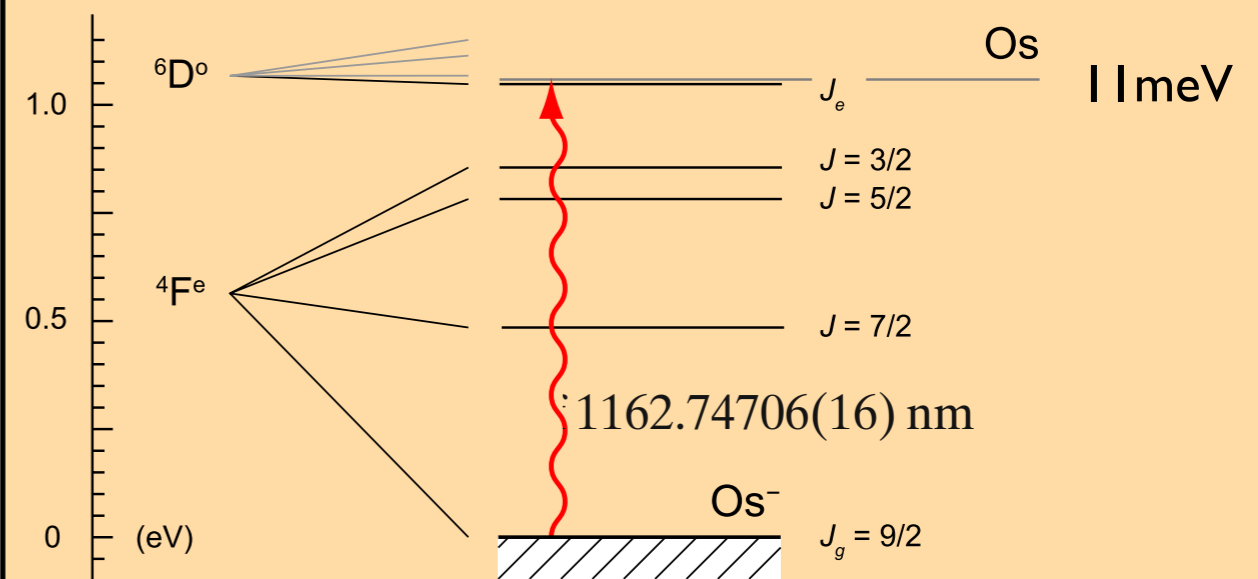
gravity measurement via "TOD"

Anion cooling for AEGIS: Os^-, La^-, C_2^-

cooling of \bar{p}

Warring et al, PRL 102 (2009) 043001

Fischer et al, PRL 104 (2010) 073004



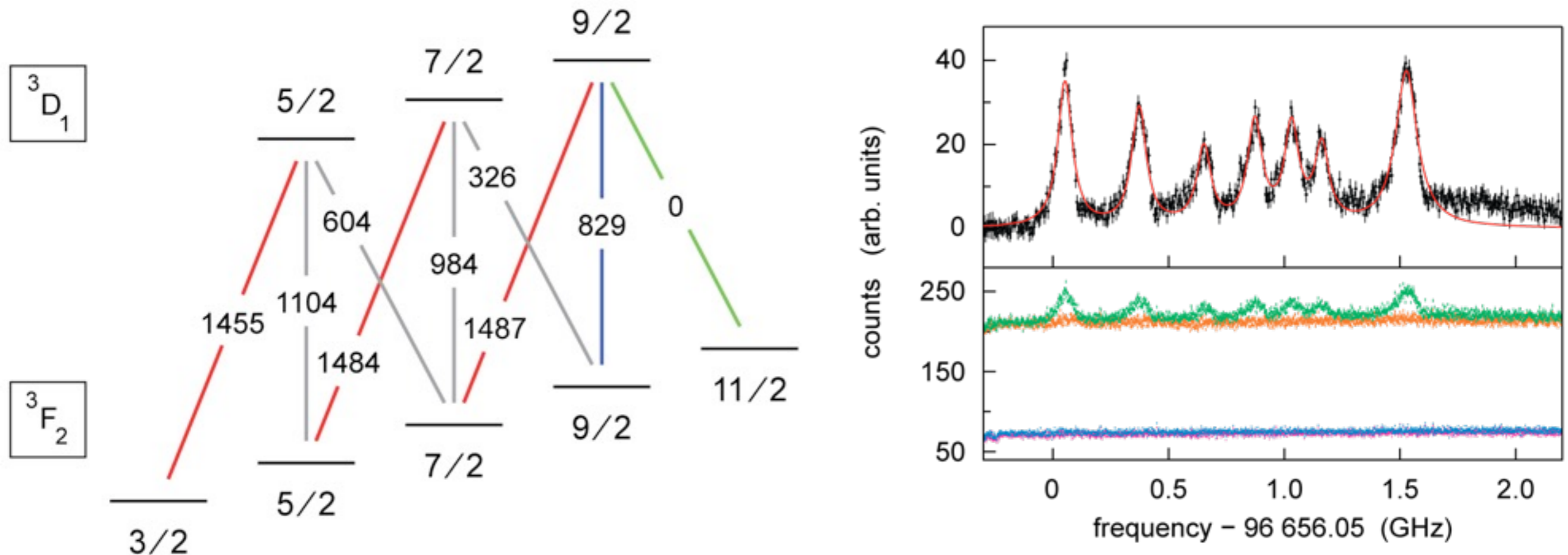
very weak cooling

\rightarrow best to start at $\sim 4\text{K}$ and cool to Doppler limit ($T_D \approx 0.24 \mu\text{K}$)

should allow reaching same precision on g as with atoms (10^{-6} or better)

laser-cooling of anions (→ sympathetic cooling of antiprotons)

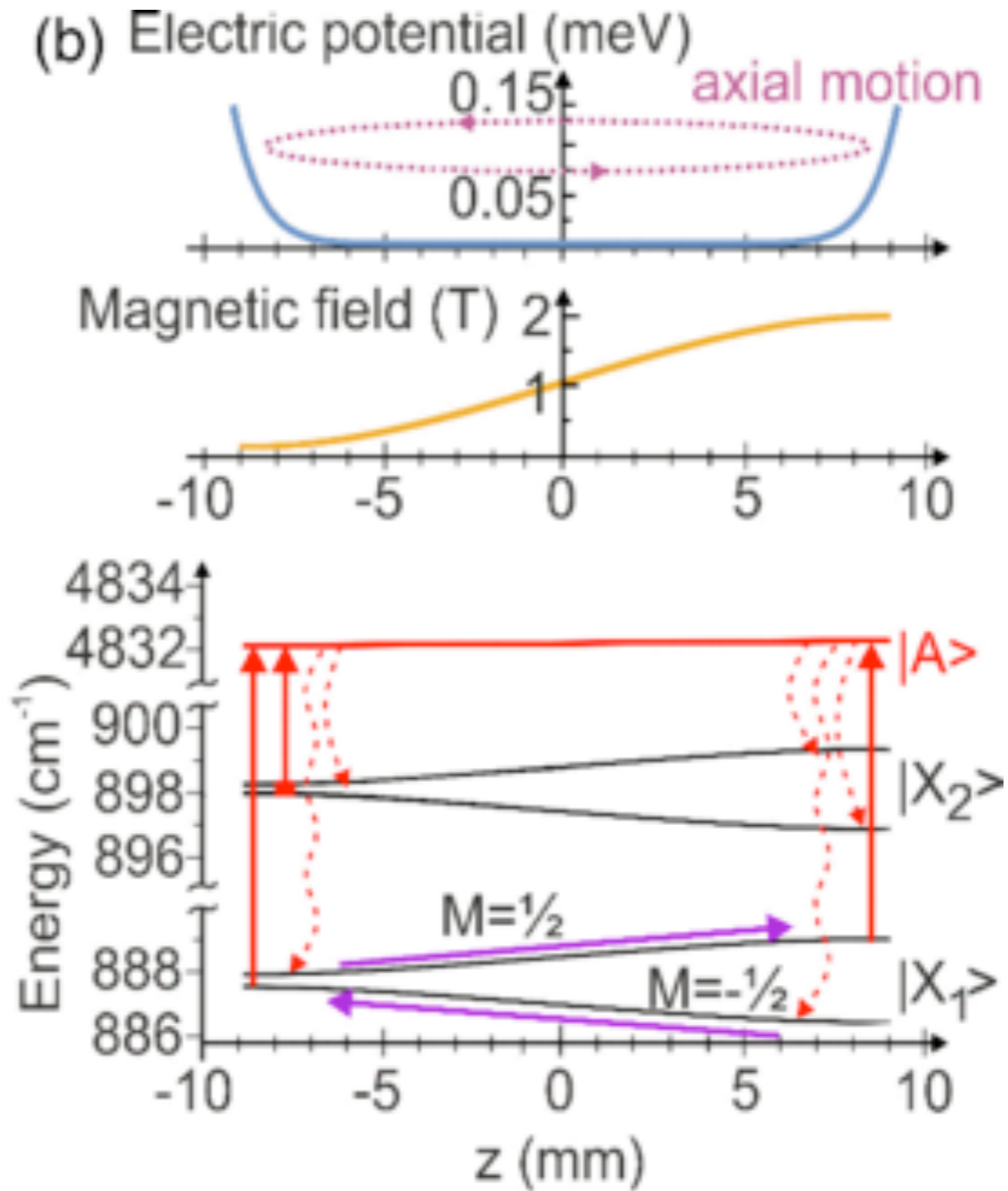
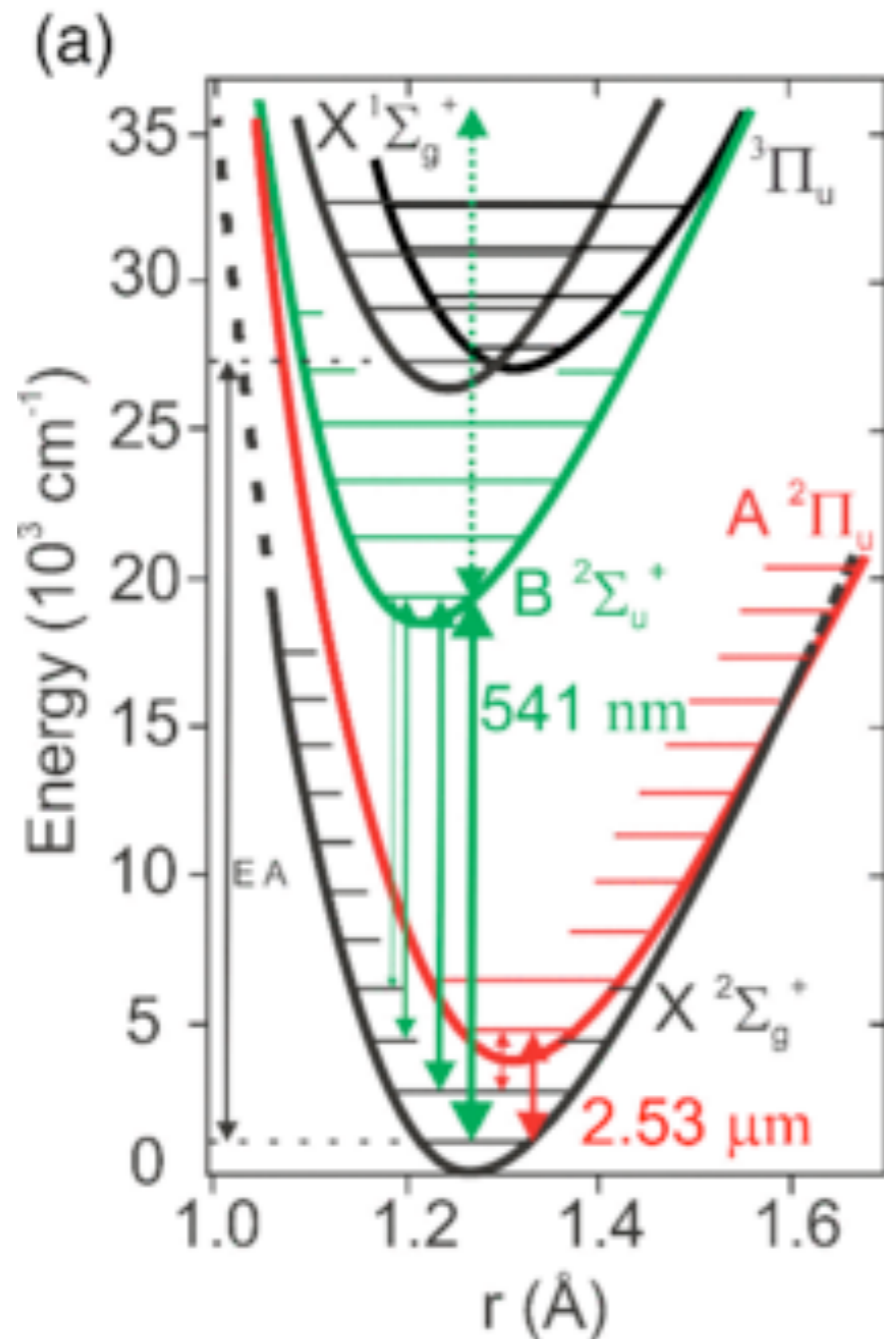
ongoing work in Heidelberg with La^- : HF transitions fully characterized
transition (cooling) rate of several kHz
(only) 3 laser wavelengths required for cooling



- next step: trapping, cooling of La^-

Anion cooling for AEGIS: C_2^-

Sisyphus cooling



Electronic and vibrational levels of C_2^-

Arrow width \sim Franck-Condon transition strength

other measurements with
antihydrogen-like atoms & ions...

\bar{H} : charge neutrality ...

Ps , muonium: gravity (lepton sensitivity)

$\mu\bar{p}$: gravity (2nd generation), antiproton charge radius

$\bar{p}p$, $\bar{p}d$: gravity (baryon sensitivity), spectroscopy, ...

ions: \bar{H}^+ gravity, CPT (ultra-cold \bar{H})

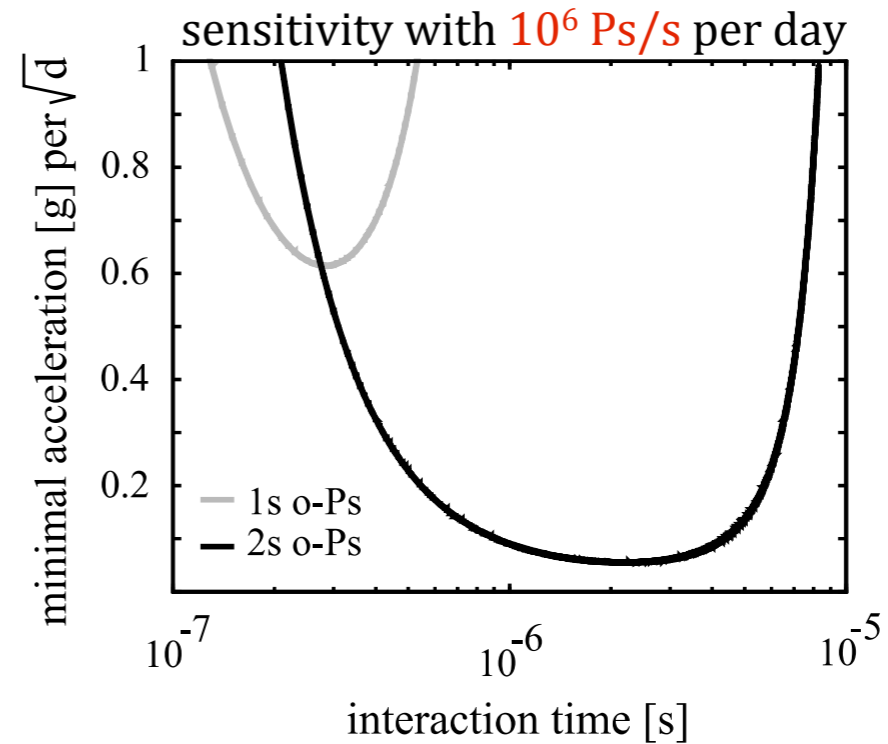
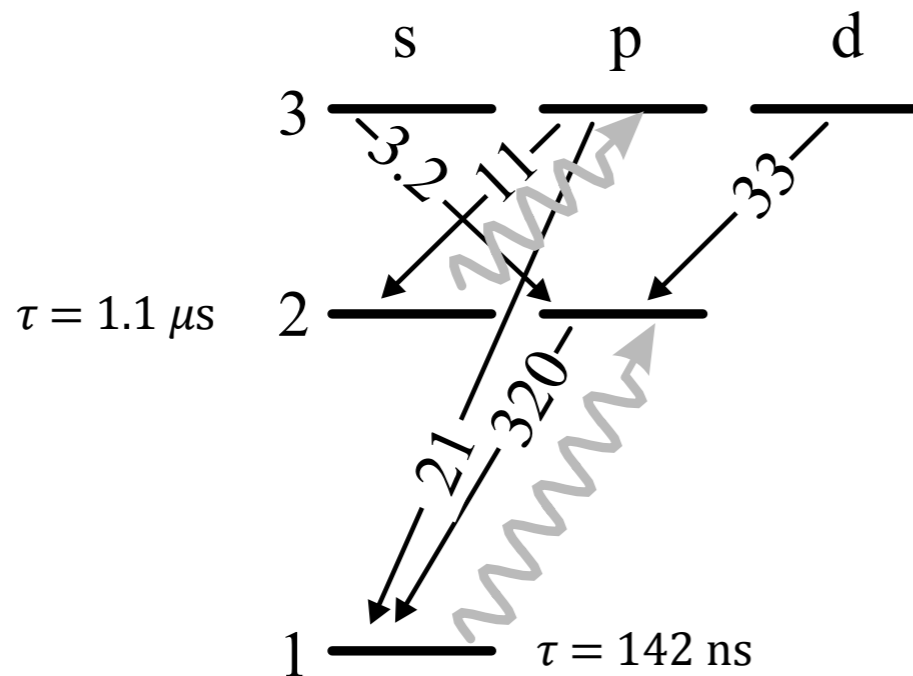
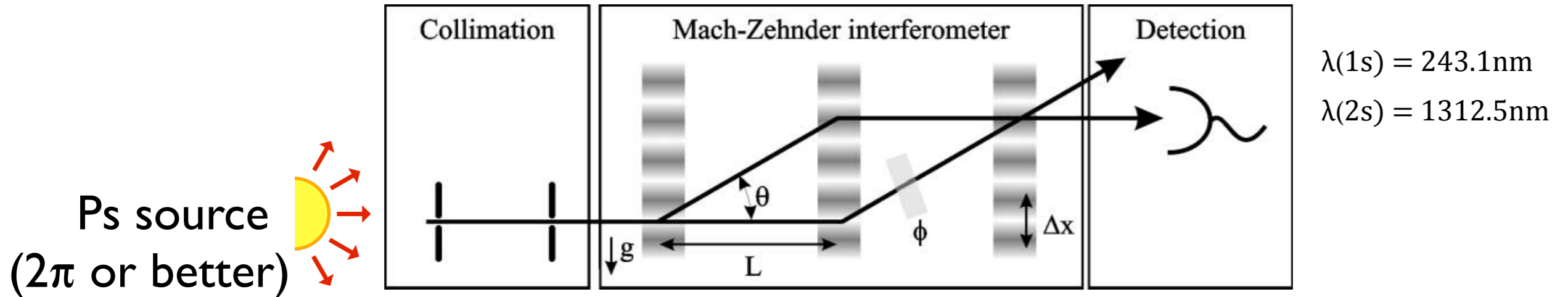
ions: H_2^+ , resp. \bar{H}_2^- proton-electron mass ratio μ

$\bar{p}N$: trapped \bar{p} (AD) + radioisotopes (ISOLDE) = PUMA

positronium...

physics interest: QED atomic spectrum, **gravity (via matter wave interferometry)**

M. Oberthaler, [Volume 192, Issues 1-2](#), (2002) 129



$v_{\text{Ps}} \sim 100\ \text{km/s} \rightarrow$ interaction time of $1\ \mu\text{s} \sim 10\ \text{cm}$

protonium...

physics interest: QCD-induced shift, broadening of QED atomic spectrum

“traditionally” formed by injecting \bar{p} into liquid hydrogen

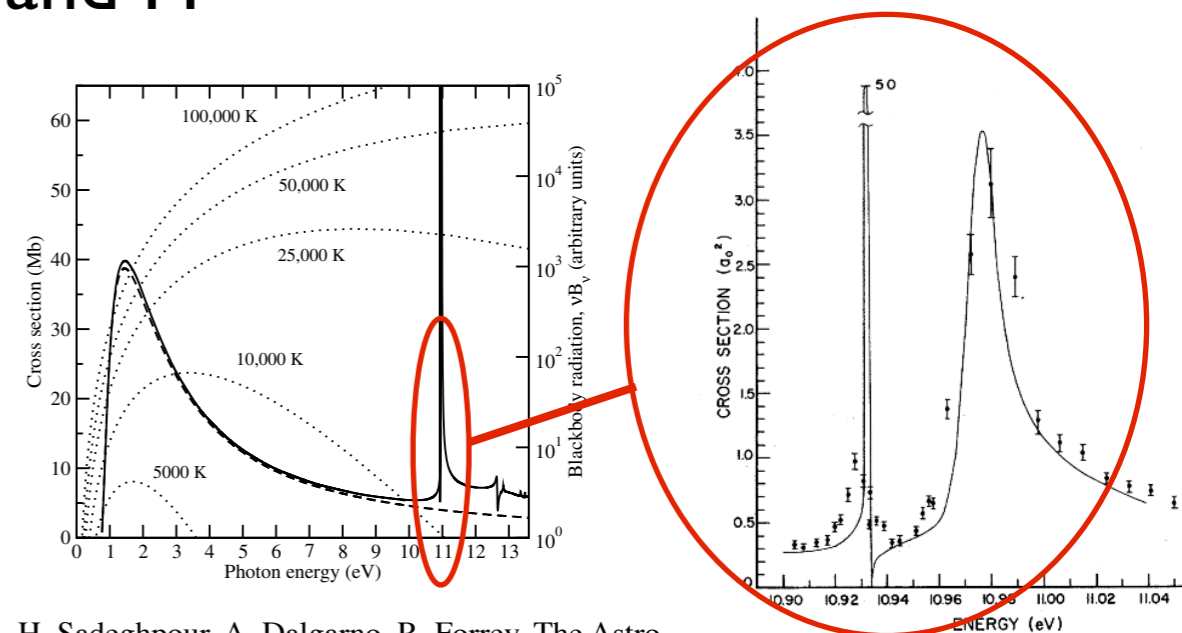
spontaneous formation in $n \sim 40$, Stark mixing, rapid annihilation

spectroscopy resolution determined by fluorescence detector resolution

alternative: **pulsed** formation via co-trapped \bar{p} and H^-

- photo-ionize $H^- \rightarrow H + e^-$
- charge exchange $H + \bar{p} \rightarrow p\bar{p}(40) + e^-$

pulsed formation \rightarrow **laser spectroscopy on $p\bar{p}$** ;
resolution determined by **laser resolution**



H. Sadeghpour, A. Dalgarno, R. Forrey, The Astrophysical Journal Letters, 709:L168–L171, 2010

H. C. Bryant et al., PRL 38 (1977) 228

improvements: **formation rate** increased if $n(H) \gg 1$

improvements: **life time** increased if $n(H) \gg 1$

\rightarrow long-lived cold Rydberg protonium \rightarrow trap/beam \rightarrow **gravity measurement**
precision spectroscopy

longer-term outlook

- advances on spectroscopy with \bar{H} and $\bar{p}\text{He}^+$, as well as in precision measurements with \bar{p} have been impressive in the last few years...
- in these systems, CPT tests now reach $\sim 10^{-10}$ and have the potential to improve sensitivity by several orders of magnitude in the coming years
- **tests of the WEP are becoming feasible**, with precisions that can be expected to initially reach % or ‰ level

work towards **ultra-cold \bar{H}** will open up additional experimental techniques and should lead not only to improved precision tests of CPT, but also of the gravitational interaction: atomic fountains, & laser-interferometric techniques, benefitting from the past and ongoing progress in the fields of atomic physics, quantum optics, molecular physics, ...

Further antihydrogen-like systems like $\bar{p}\mu^+$, Ps , $\bar{p}\text{p}$, \bar{H}^+ , \bar{H}_2^- (and much patience and ingenuity) offer additional opportunities for intriguing tests (**gravity**, high sensitivity measurements of antiproton/positron mass ratio, **gravity tests in purely baryonic or leptonic systems**, ...)