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FOCUS DATA REDUCTION

Contents:

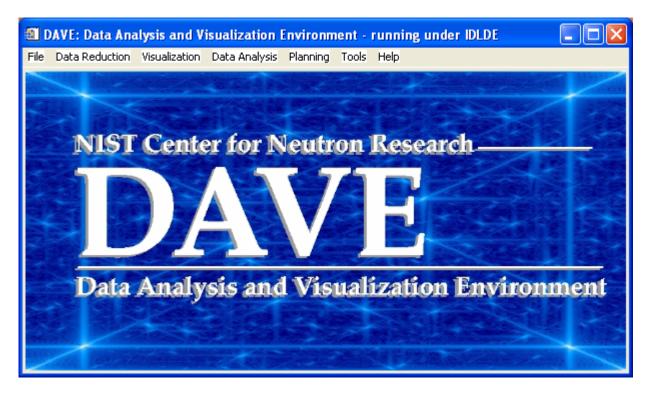
- 1. Introduction
- 2. The Basic Program Interface
- 3. Advanced Program Options
- 4. A Working Example

1. Introduction

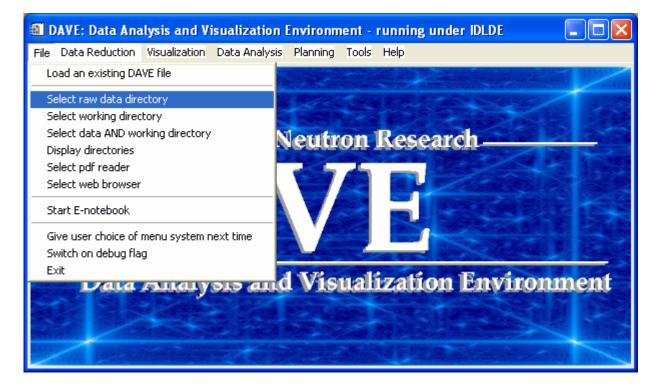
This document describes the program to reduce data collected on the direct time-of flight spectrometer FOCUS, at the SINQ facility located at the Paul Scherrer Institut in Switzerland. The program is known as FDR, which stands for Focus Data Reduction and should be easy to remember by those famililar with American history. FDR was written in IDL 6.1/6.2. and is incorporated within the DAVE program package. DAVE stands for "Data Analysis and Visualisation Environment", and is essentially a series of programs developed by Rob Dimeo and his colleagues at the NIST facility, Washington D.C., for the reduction and analysis of inelastic neutron scattering data. Once the user has reduced a raw FOCUS file, the data analysis programs within the DAVE package will be at his disposal.

2. The Basic Program Interface

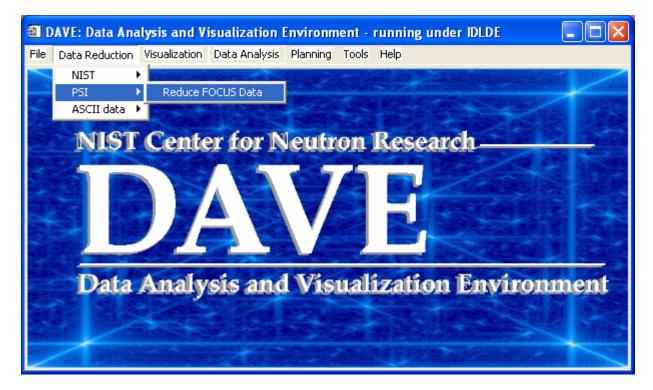
When DAVE is launched in classic view, the user should see a screen like the one shown below:



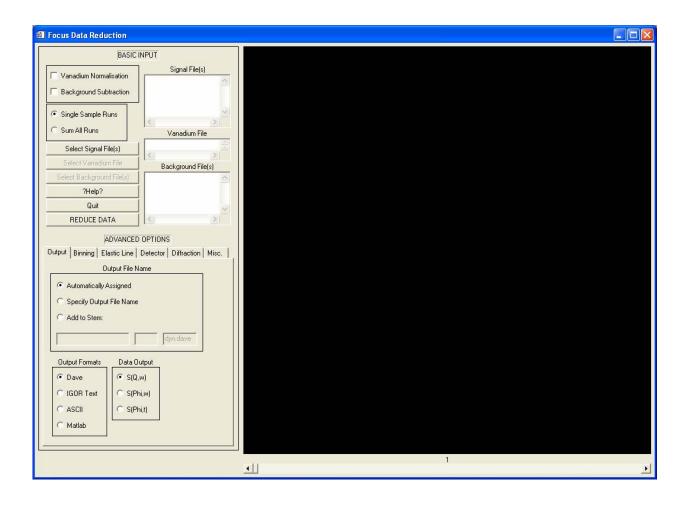
First of all, the program environment needs to be configured, if this has not been done already. This is achieved by clicking on the "File" menu header and selecting the options appropriately. The raw data files should be contained in the raw data directory. All the files that the program outputs will be saved to the working directory.



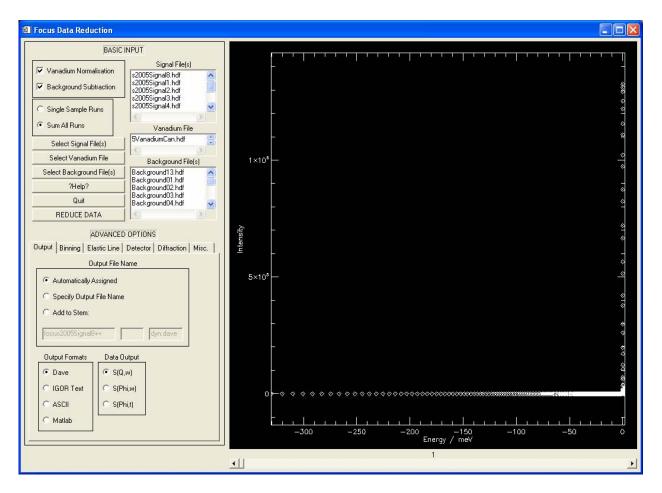
To initiate FDR, click on the Data Reduction Menu and move the mouse to select the program as follows:



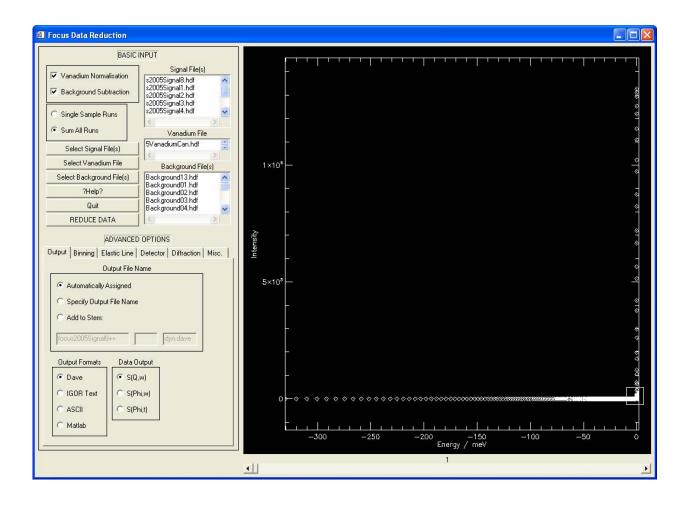
The main menu for the program should then appear:

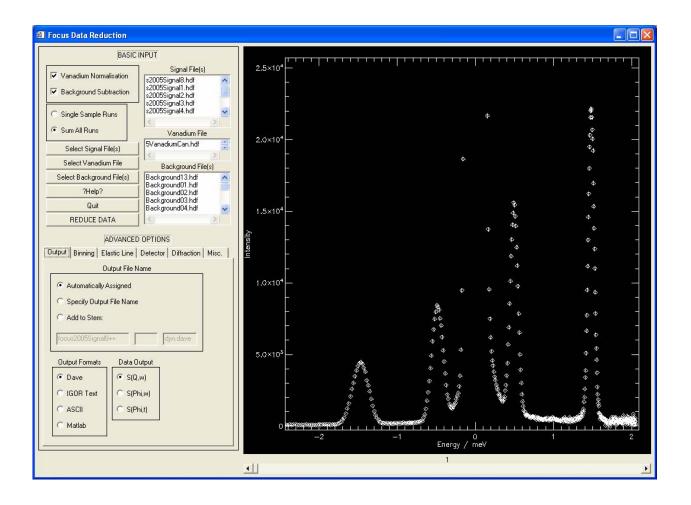


The interface was conceived with the idea of accommodating both the novice and the seasoned user. A data set can be reduced quickly and reliably by entering just the basic input, and allowing the program to choose the remaining settings. The only essential information required by the program is signal files. The basic input menu largely follows that written by Rob Dimeo for his HFBS reduction program, allowing the user to reduce a number of signal files individually (Single Sample Runs) or to sum the files together (Sum All Runs). Background files consist of data from the empty can. The files are automatically summed if more than one is present. The vanadium file consists of data from vanadium metal. When files in all three categories are selected, and the settings of the ADVANCED OPTIONS have not been altered, the action of pressing the "REDUCE DATA" button yields an output that should look something like this:



From inspection of the "Output" menu under "ADVANCED OPTIONS", it may be seen that the program has converted the raw data to $S(Q,\omega)$, and saved the output as a DAVE file, with the stem of the file name taken from the input signal file. The spectrum is plotted on the right-hand-side of the display window. The useful information is typically contained within a few meV of the exciting line. The user can zoom-in on the range of the spectrum of interest by creating a window using the mouse with the left button depressed and then releasing. A click of the right mouse button returns the view window to display the full range.



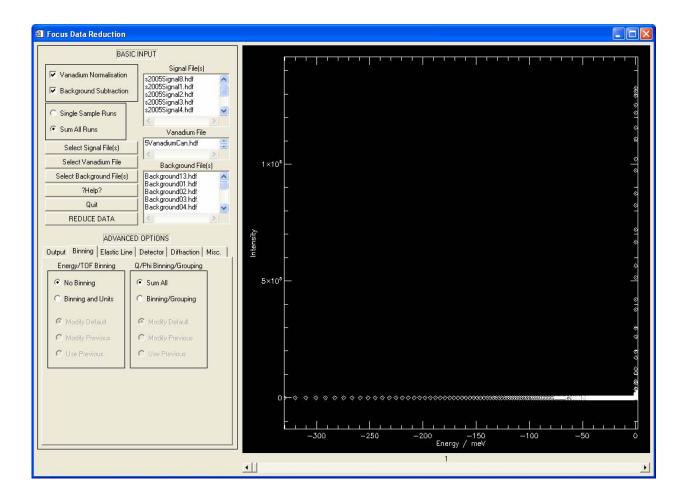


The remaining default settings may be inspected by clicking on the tabs under "ADVANCED OPTIONS".

3. Advanced Program Options

In this section, details are given on those Advanced Options that are not self-explanatory.

The default is no energy binning with the data summed over all detectors.



When the "Binning and Units" option is selected under "Energy/TOF Binning", the following menu appears during the reduction process:

| Energy Rebinning Utility | X |
|---|--|
| Range Without Binning # of channels 425 Lower limit -330.459 Upper limit -2.04991 | Binned Range With Interpolation # of channels 187 Lower limit -2.04991 Upper limit 0.562620 |
| Range With Binning # of channels 479 | Range to Include |
| Lower limit -2.04991 | Binned |
| Upper limit 2.04991 | C Binned Without Interpolation |
| Step Size 0.00857704 | Units |
| Accept | ⊙ meV O cm-1 O GHz |
| Cancel | |

The energy range is divided into regions with and without binning. The binned region is further divided into regions with and without interpolation. All fields are automatically updated by changing one of the editable fields and then pressing the carriage return. The total number of output channels cannot exceed the total number of input channels. If binned data are required over a wide energy range, but interpolation is not desired, the step size should be increased, and/or the number of channels decreased, until the interpolation range diminishes to zero.

| ed Range With Interpolation nnels 0 nit -2.0499100 nit -2.0499100 |
|---|
| Range to Include ned ned Without Interpolation Units V O cm-1 O GHz |
| _ |

The default is to sum the data summed over all detectors. If Q binning is desired then "Binning/Grouping" under "Q/Phi Binning/Grouping" should be selected. Before attempting to calculate the Q dependence, the user should know the accessible Q-range for the transition in question. This may be determined using the DAVE neutron calculator facility, which is to be found under the "Planning" menu.

When Q binning is selected the following dialogue appears:

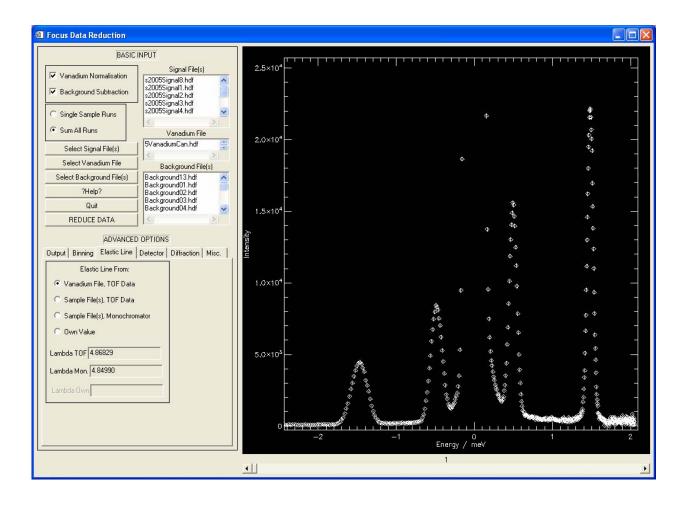
| a QRebinning Utility | × |
|--------------------------------|--------------------------------|
| Current Binning # of bins 5 | Desired Binning # of bins 5 |
| Lower Q limit 0.25106353 | Lower Q limit 0.25106353 |
| Upper Q limit 2.6439979 | Upper Q limit 2.6439979 |
| Step Size 0.47858687 | Step Size 0.47858687 |
| Acc | cept |
| Car | ncel |

Press the carriage return after modifying a given editable field. If the quantity to be calculated is not $S(Q,\omega)$ but $S(phi,\omega)$ then checking the "Binning/Grouping" box under "Q/Phi Binning/Grouping" results in the following dialogue appearing during the reduction process:

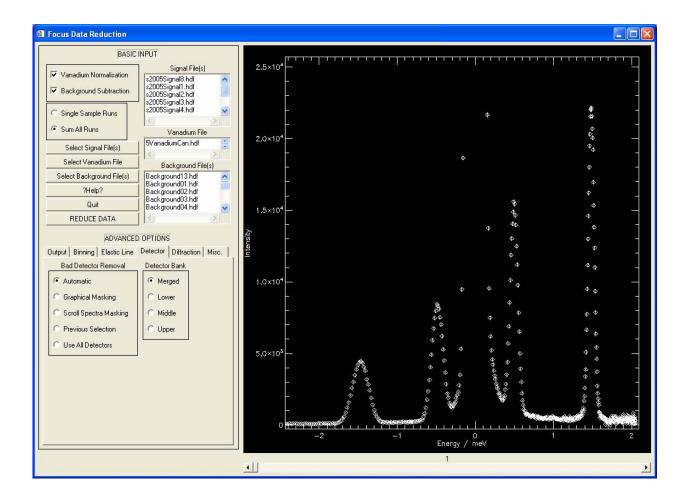
| oup | E very: | 1 | Spe | cify Gro | ouping: | 1-10,1 | 1-18,19 | 9-78 | | | Gro | up all | | | | | | | | | | | | | | | |
|------|----------|---------|-----|----------|----------|----------|---------|------|------|----|-----|--------|-----|------------|-----|-----|----|-----|----|----------|----|-----|----|-----|----|-----|----|
| imbe | er of Gr | oups: [| 266 | | | | | | | | | | | | | | | | | | | | | | | | |
| tecl | tor Num | iber 1 | | · Thet | a Value: | 1 | 1.217 | | | | | | | | | | | | | | | | | | | | |
| | | 1 | | | | | | | | | | | Deb | ector | | | | | | | | | | | | | |
| , | 1 | | 2 | | 3 Г | - | 4 [| 1 | 5 🗆 | 6 | Г | 7 | | 22 | 8 | (| | 10 | | 11 | Г | 12 | Г | 13 | Г | 14 | |
| | | Г | 16 | F | 17 | | 18 | | 19 Г | - | 20 | Γ. | 21 | | 22 | Г | 23 | Г | 24 | F | 25 | Г | 26 | Г | 27 | E | 28 |
| - | 29 | Г | 30 | - | 31 | F | 32 | | 33 Г | - | 34 | Г | 35 | Г | 36 | Г | 37 | Г | 38 | Г | 39 | Г | 40 | Г | 41 | Г | 42 |
| | 43 | Г | 44 | Г | 45 | Г | 46 | | 47 T | | 48 | Г | 49 | Г | 50 | Г | 51 | Г | 52 | Г | 53 | Г | 54 | Г | 55 | Г | 56 |
| - | 57 | Г | 58 | Г | 59 | Г | 60 | Г | 61 F | | 62 | | 63 | Г | 64 | Г | 65 | Г | 66 | Г | 67 | Г | 68 | Г | 69 | Г | 70 |
| - | | Г | 72 | Г | 73 | Г | 74 | Г | 75 F | - | 76 | Г | 77 | Г | 78 | Г | 79 | Г | 80 | Г | 81 | Г | 82 | Г | 83 | Г | 84 |
| - | 85 | Г | 86 | | 87 | Г | 88 | Г | 89 F | í. | 90 | Г | 91 | Г | 92 | Г | 93 | Г | 94 | Г | 95 | Г | 96 | Г | 97 | | |
| | 98 | Г | 99 | Г | 100 | Г | 101 | Г | 102 | Г | 103 | Г | 10- | 4 Г | - 1 | 05 | Г | 106 | Г | 107 | Г | 108 | Г | 109 | Г | 110 | |
| 1 | 111 | | 112 | Г | 113 | Г | 114 | Г | 115 | Г | 11 | 6 🗖 | 1 | 17 | Г | 118 | Г | 119 | Г | 120 | Г | 121 | | 122 | Г | 123 | |
| | 124 | Г | 125 | Г | 126 | Г | 127 | Г | 128 | Г | 12 | :9 F | 1: | 30 | Г | 131 | Г | 132 | | 133 | Г | 134 | Г | 135 | Г | 136 | |
| | 137 | Г | 138 | Г | 139 | Г | 140 | Г | 141 | Г | 14 | 2 [| 1 | 43 | Г | 144 | Г | 145 | Γ | 146 | Г | 147 | Г | 148 | Г | 149 | |
| | 150 | Г | 151 | Г | 152 | Γ | 153 | Г | 154 | Г | 15 | 5 🗆 | 1 | 56 | Г | 157 | Г | 158 | Г | 159 | Г | 160 | | 161 | Г | 162 | |
| | 163 | Γ | 164 | Г | 165 | Г | 166 | Г | 167 | Γ | 16 | 8 🗆 | 1 | 59 | Г | 170 | Γ | 171 | Γ | 172 | Γ | 173 | | 174 | Γ | 175 | |
| | 176 | Г | 177 | Г | 178 | Г | 179 | Г | 180 | Г | 18 | n L | i 1 | 32 | Г | 183 | Г | 184 | Г | 185 | Г | 186 | Г | 187 | Г | 188 | |
| | 189 | | 190 | Г | 191 | Г | 192 | Г | 193 | Γ | 19 | 14 🗆 | 1 1 | 95 | Г | 196 | | 197 | | 198 | Г | 199 | Г | 200 | Г | 201 | |
| | 202 | | 203 | Г | 204 | Γ | 205 | Г | 206 | Γ | 20 | | | 08 | Г | 209 | Г | 210 | | 211 | Г | 212 | Г | 213 | Г | 214 | |
| | 215 | - | 216 | Γ | 217 | Γ | 218 | Г | 219 | Γ | 22 | | 2. | | Γ | 222 | | 223 | | 224 | | 225 | Г | 226 | Г | 227 | |
| | 228 | | 229 | | 230 | | 231 | Г | 232 | | 23 | | 2 | | | 235 | | 236 | | 237 | | 238 | | 239 | | 240 | |
| | 241 | | 242 | | 243 | | 244 | Г | 245 | | 24 | | | | | 248 | E | 249 | | 250 | | 251 | Г | 252 | | 253 | |
| | 254 | | 255 | | 256 | | 257 | | 258 | | 25 | i9 | 21 | 50 | | 261 | Γ | 262 | | 263 | | 264 | | 265 | Г | 266 | |

The detectors may be grouped in three different ways: By entering a value in the "Group Every" field and pressing the carriage return; by entering a string in the "Specify Grouping" field, paying attention to the format in the example; and finally by checking the boxes for a given group. The detectors selected in a given group may be viewed by moving the slider bar appropriately.

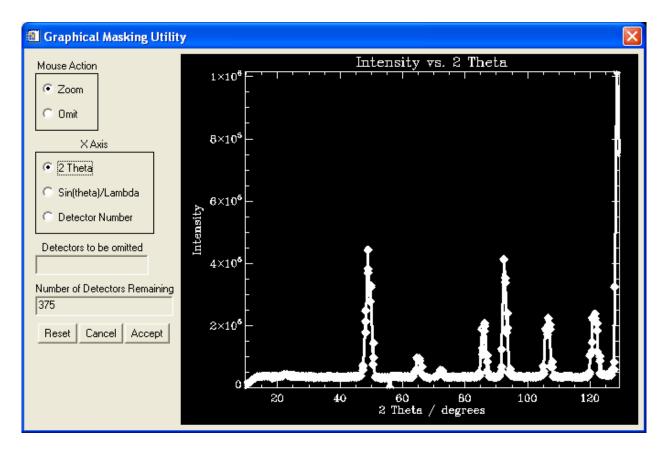
The elastic line will be calculated from the vanadium file unless otherwise specified. If no vanadium file is selected then the default is to calculate the elastic line from the signal data. During the reduction process, the value of lambda, calculated from time-of-flight data will be displayed. This should be similar to the value calculated from the monochromator position.



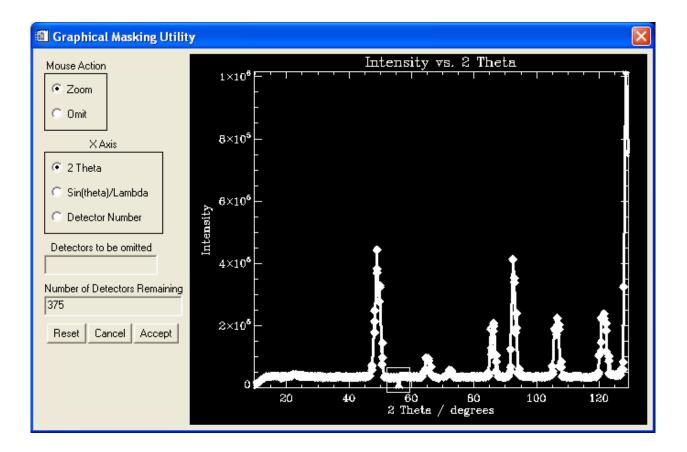
The automatic bad detector removal option removes all detectors with a signal less than a 20% percentage of the average. It is especially important to remove dead detectors, as the program will crash if these detectors are not removed and the program tries to correct for variable detector efficiency. There are three detector banks: upper, middle and lower. The default "merged" option is self-explanatory.

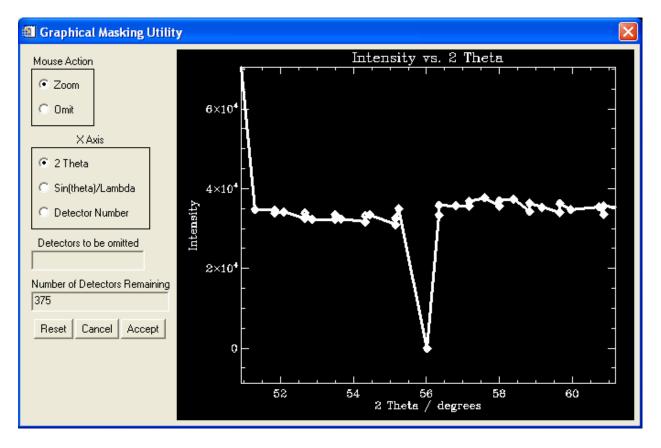


The Graphical Masking and Scroll Spectra Masking tools allow the user to remove certain detectors manually during the reduction process. If the Graphical Masking option is chosen, the following dialogue appears during the reduction process:

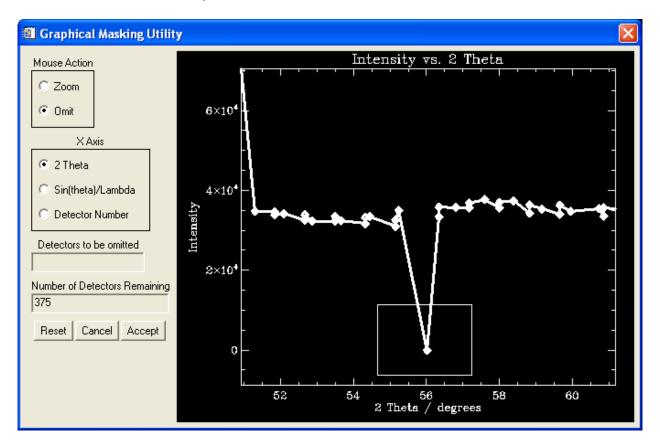


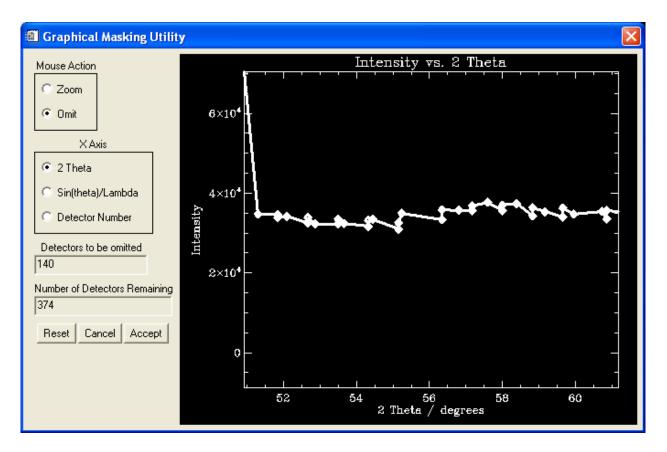
The data may be plotted as a function of 2 Theta, Sin(theta)/lambda or detector number. Be aware that some detectors will have the same value of theta, when housed in different detector banks. Zoom into the area of interest using the mouse.





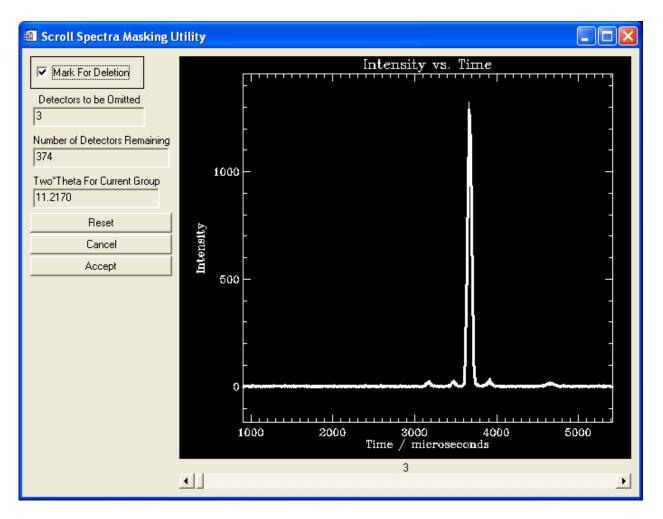
Next, change the mouse action to omit. Detectors, whose intensity and x value are enclosed within the window defined by the mouse, will now be removed:





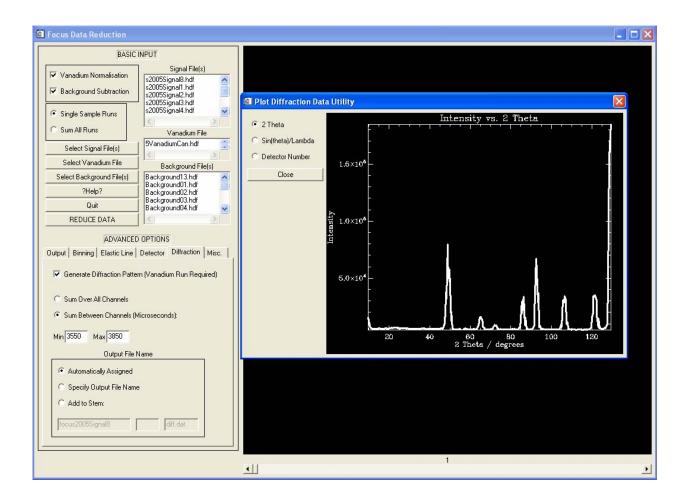
Press "Accept" when completed.

By contrast, the Scroll Spectra Masking option allows the user to view the raw data for each detector.



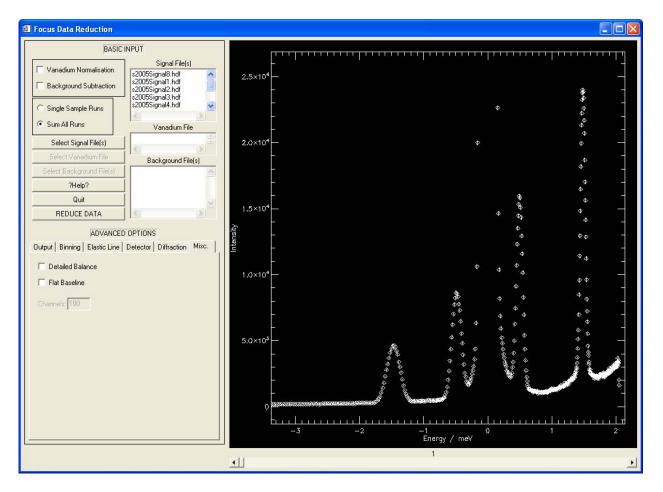
The data from a given detector may be viewed by moving the slider bar appropriately. Those detectors marked for deletion will be removed when the Accept button is pressed.

A low-resolution diffraction pattern may be calculated from data collected on the FOCUS instrument. In order for the intensities to be meaningful, the detectors must be calibrated using data from a vanadium run. The intensities may be calculated either from the sum over all channels, or from channels in a given time window. The latter option is preferable when inelastic transitions are strong. The diffraction pattern is plotted during the reduction process and a file containing the data is created.

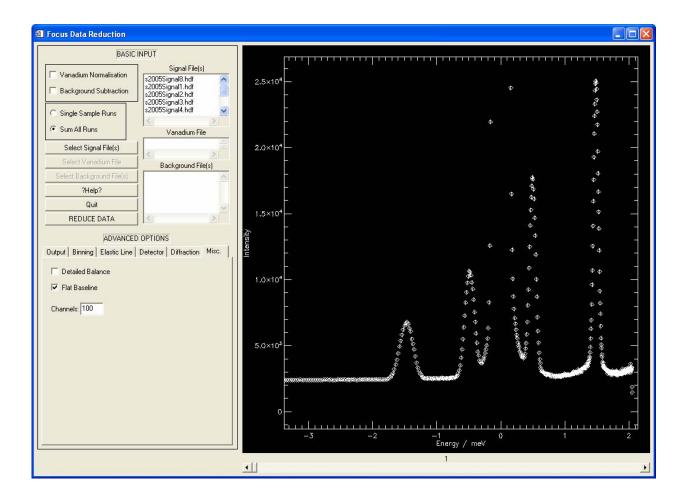


If the Detailed Balance option is selected under the Misc. tab, the data are multiplied by the factor $e^{-E/2kT}$. The "correction" is appropriate only for excitations obeying Bose statistics, phonons for example.

If data of the empty can were not collected, then the baseline can be fiddled using the flat baseline option, also to be found under the Misc. tab. With no background subtraction, the baseline is seen to rise steeply on the neutron energy loss side of the spectrum:

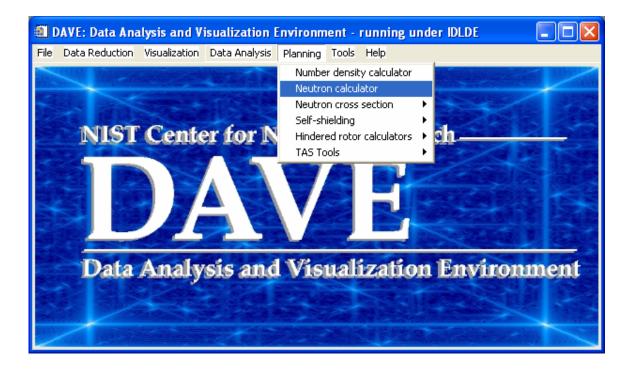


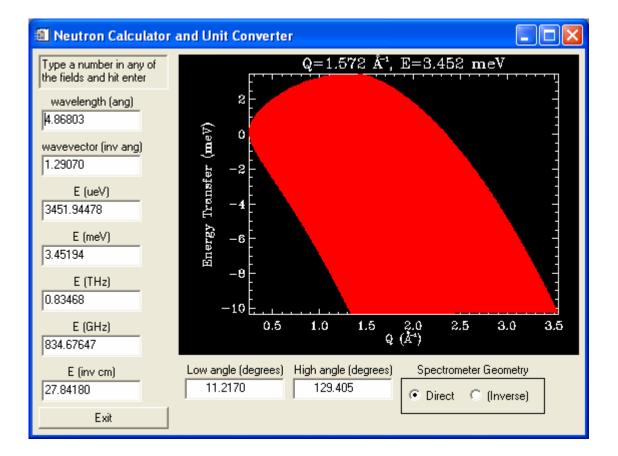
This is because the conversion from time to energy incorporates an intensity correction of t_f^4 , where t_f is the time-of-flight from the sample to the detector. The flat baseline option subtracts a constant from all the raw data before conversion to energy. This constant is calculated as follows: The mean number of counts is calculated successively over a given number of channels. In the example shown below, 100 channels. If there are 904 channels, then 10 mean values are calculated. The lowest mean value is the constant to be subtracted from all the data. By removing this offset, a flat baseline can be obtained



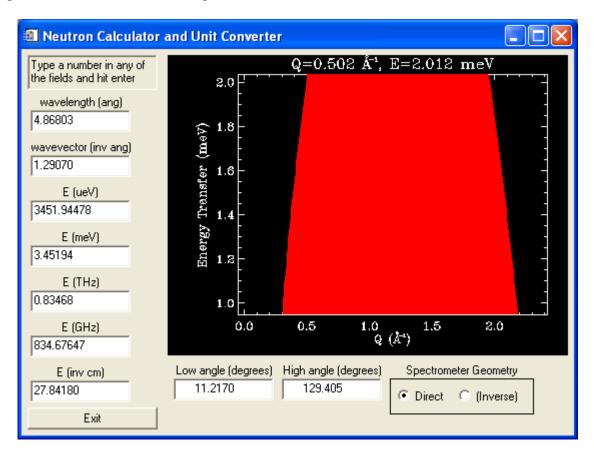
4. A Working Example

In this section, the steps that the user might take to reduce his data set, and calculate the Q-dependence of a given transition, are described. Let us calculate the Q-dependence of the intensity of the high-energy transition in the spectrum above. The first step is to know the accessible Q-range for the transition in question. This is determined by the energy of the incoming neutrons and the range of detector angles. Start the DAVE neutron calculator as shown, and enter the approproate values.

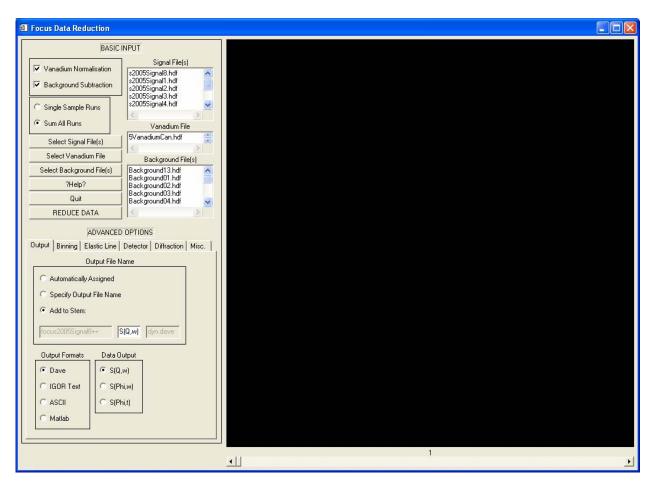




The red section shows the accessible Q-range. For a transition falling in the 1 to 2 meV energy range, the detectors cover a Q range from ca. 0.5 to 2.0 Å^{-1} .

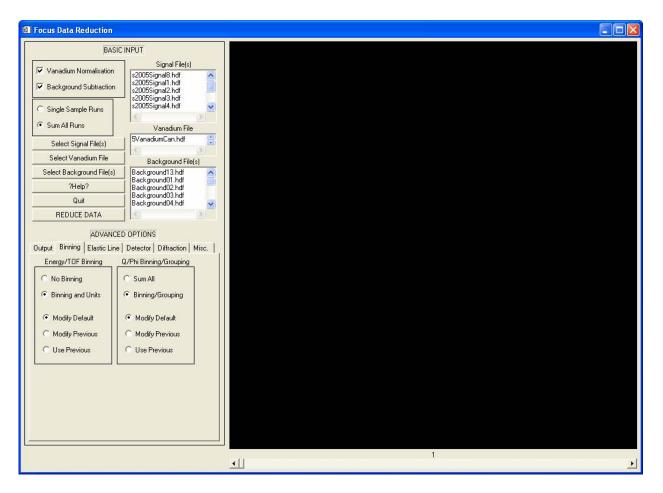


Next, we can reduce the data using the FOCUS Reduction program. A total of eight signal files were collected at 6 K for the sample of interest. We want to sum these files, subtract the background spectrum and calibrate the detectors with a vanadium file.

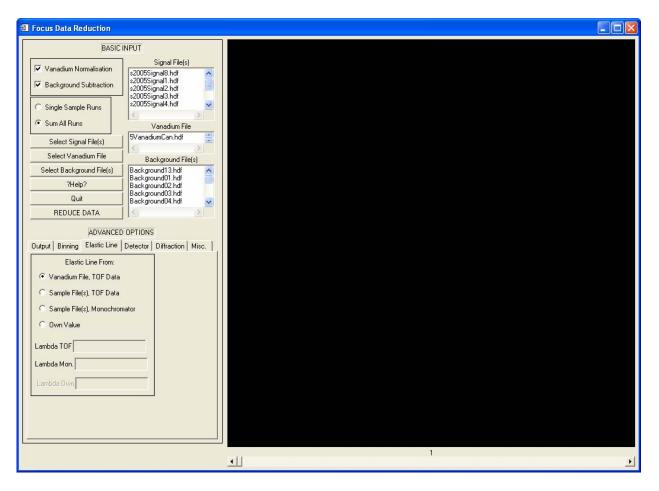


Returning to FDR, we have added S(Q,w) to the stem of the file name, to remind us of the data output we have specified.

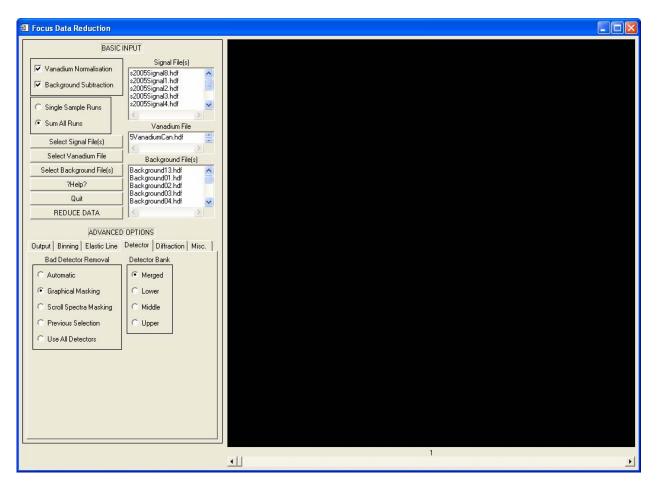
Both energy and Q binning are required:



The elastic line shall be calculated from the vanadium run:



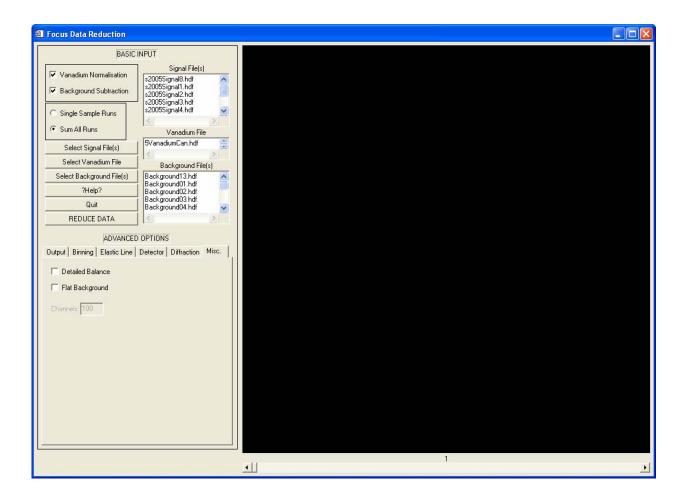
We shall incorporate detectors from all three banks, and remove the bad ones using the Graphical Masking Utility:



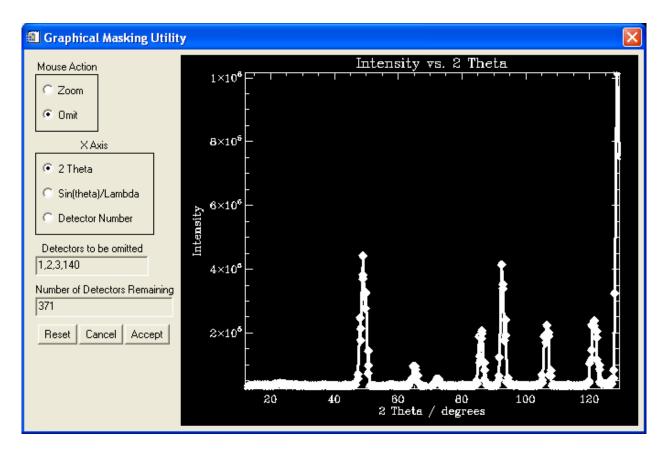
We also wish to calculate the diffraction pattern. The elastic line falls in the time window 3550 to 3850 microseconds, and we specify this time window accordingly.

| E Focus Data Reduction | | |
|--|---|--|
| BASIC | INPUT | |
| Vanadium Normalisation | Signal File(s) \$2005Signal8.hdf \$2005Signal1.hdf \$2005Signal2.hdf | |
| Background Subtraction | s2005Signal3.hdf | |
| Single Sample Runs Sum All Runs | s2005Signal4.hdf | |
| Select Signal File(s) | 5VanadiumCan.hdf | |
| Select Vanadium File | Background File(s) | |
| Select Background File(s) | Background13.hdf | |
| ?Help? | Background02.hdf Background03.hdf | |
| Quit REDUCE DATA | Background04.hdf | |
| ADVANCEI Output Binning Elastic Line Generate Diffraction Patter Sum Over All Channels Gum Between Channels (Min 3550 Max 3850 Output File Automatically Assigned Capecify Output File Nam Cadd to Stem: focus2005Signal8++ | ern (Vanadium Run Required) Microseconds): Name | |
| | | |

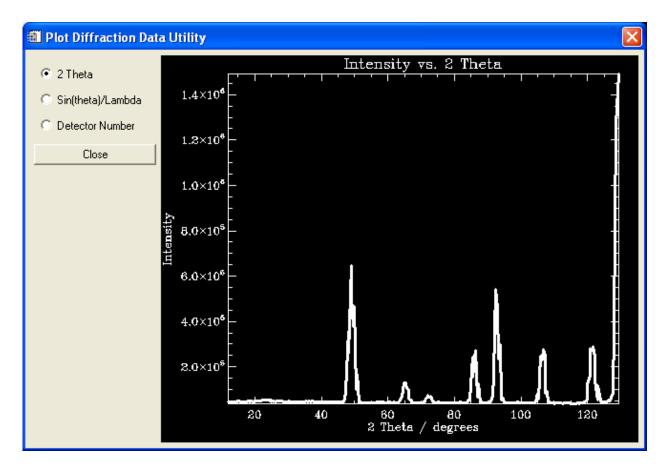
Neither the detailed balance nor the flat background option is selected. Note that if background files are selected and the flat baseline option is checked, background subtraction will proceed via subtraction of the empty can spectra only.



When the "REDUCE DATA" button is depressed, we are first invited to remove the bad detectors using the Graphical Masking Utility. From the screen shown below, the reader will note that detectors 1,2,3 and 140 have been removed. Detector 140 is a dead detector. If this detector is not removed at this stage, the program will crash.



The program then reduces the detectors to a set with unique values of 2 theta. The diffraction pattern is calculated, which the user is invited to inspect:



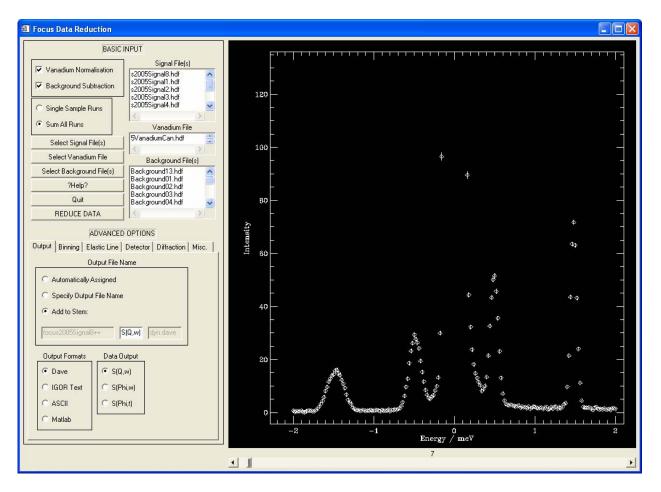
All data files are then converted to energy, and the spectra of the empty can subtracted. At this point, the user must specify the energy binning. The energy binning screen below shows that we have specified the units to be meV, with a step size of 0.02 in the range -2.0 to +2.0. The program tells us that in the binning process, points will be created in the energy range -2.0 to -1.62 meV.

| 🎒 Energy Rebinning Utility | |
|---|---|
| Range Without Binning # of channels 428 Lower limit -330.45902 Upper limit -2.0 | Binned Range With Interpolation # of channels 19 Lower limit -2.0000000 Upper limit -1.62101 |
| Range With Binning # of channels 201 Lower limit -2.0 Upper limit 2.0000000 Step Size 0.0200000 Accept Cancel | Range to Include C All Binned Binned Without Interpolation Units T meV C cm-1 C GHz |

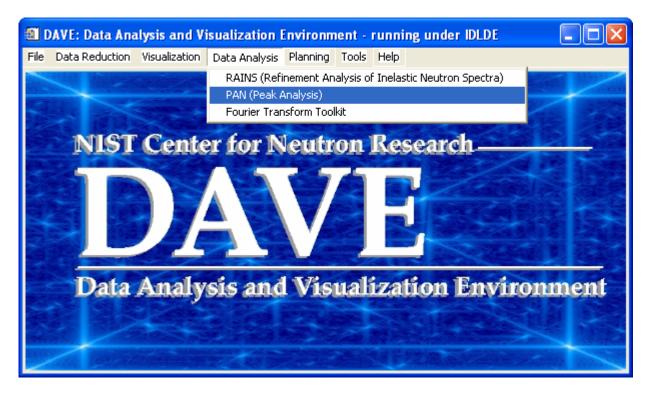
On pressing the "Accept" button, the Q Rebinning Utility appears. The default Q range is the maximum possible for the specified energy window. Since, however, we are interested in examining the Q-dependence of the transition that falls between 1.0 and 2.0 meV, we must set the minimum and maximum Q value to 0.5 and 2.0 Å⁻¹ respectively, as previously determined by the neutron calculator. We specify a step size of 0.2 Å⁻¹, resulting in 7 Q bins.

| 💷 Q Rebinning Utility | |
|--------------------------------|--------------------------------|
| Current Binning # of bins 5 | Desired Binning # of bins 7 |
| Lower Q limit 0.26836589 | Lower Q limit 0.5 |
| Upper Q limit 2.6371385 | Upper Q limit 2.0 |
| Step Size 0.47375452 | Step Size 0.2 |
| Ac | cept |
| Ca | ncel |

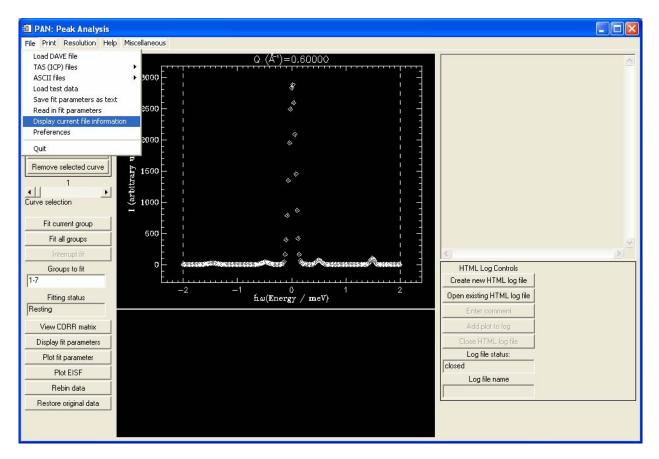
The data are then reduced. Seven spectra are calculated in total, corresponding to the seven Q bins. At this stage we are able to zoom in on the portion of the spectrum of interest, and view each spectrum utilising the group slider.



We have now done all we can with FDR. The data analysis program, PAN, is then initialised after quitting FDR.

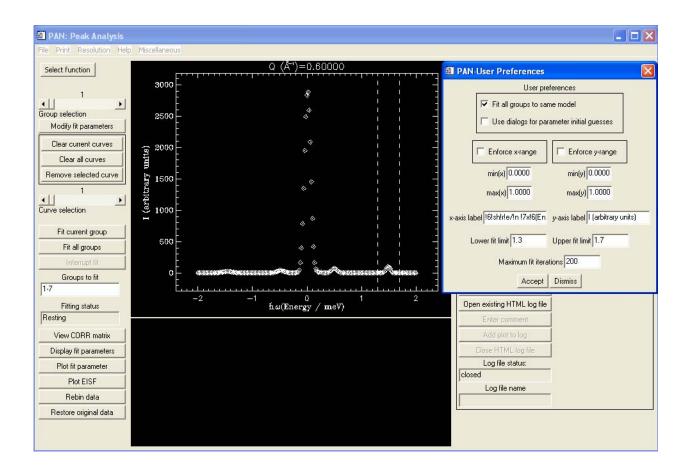


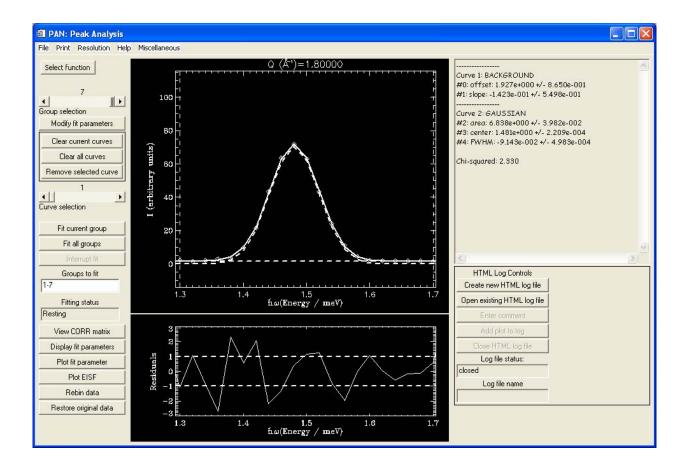
The data just reduced are read in automatically by PAN. The details of the reduction process just undertaken may be viewed by selecting "Display Current File Information":

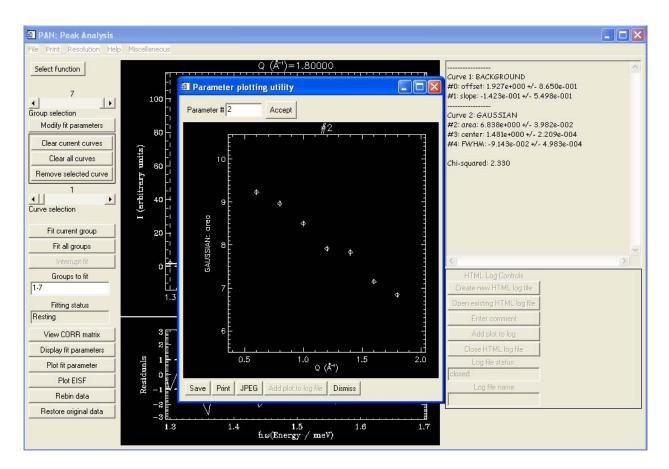


| PAN File Information | |
|--|----------|
| Data reduction/treatment details: | ~ |
| Files used in sum: | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Signal8.hdf | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Signal1.hdf | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Signal2.hdf | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Signal3.hdf | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Signal4.hdf | _ |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Signal5.hdf | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Signal6.hdf | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Signal7.hdf | |
| Total Number of Detectors Read From File: 375; | |
| Number of Detectors Excluded in the Reduction: 4; | |
| Detector numbers of Those Excluded: 1,2,3,140; | |
| 371 Detectors Reduced to 265 With Unique Two Theta Angles. | |
| Elastic Line Calculated From Vanadium File, TOF Data | |
| Diffraction Pattern Calculated from Time Channels in the Range: 3550.00 to 3850.00 Microseconds. | |
| Background file(s) used in subtraction: | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Background13.hdf | |
| C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Background01.hdf C:\Documents and Settings\tregenna\My Documents\temp\dave\Raw Data Directory\focus2005Background02.hdf | |
| C: Documents and Settings/tregerina/wy Documents/temp/dave/haw Data Directory/tocus2005Background03.hdf | |
| C: Documents and Settings/tregerina/wy Documents/temp/dave/haw Data Directory/tocus2005ackg/ound04.hdf | |
| C: Documents and Settings tregerina wy Documents treinp dave hav Data Directory Viocus2005 ackground/b, hdf | |
| C: Documents and Settings/tregerna/My Documents/temp/dave/trav/Directory/locus2005Background06.hdf | |
| C: Documents and Settings tregerinality Documents temp date in an Directory Vocus2005Background07.hdf | |
| c. became it is a tradition of the strain of | <u> </u> |
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| Dismiss | |
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PAN may be used to model the spectra as a superposition of different functions. In order to examine the Q-dependence of the intensity of the high-energy peak, the peak is fitted to a superposition of a linear background and Gaussian function, and the area of the Gaussian function plotted as a function of Q. The area of the peak in the seven files can be exported to an ascii file as shown. Please consult the PAN help file for futher details.







The data may be displayed using the data browser program program, which is found under visualization.

